

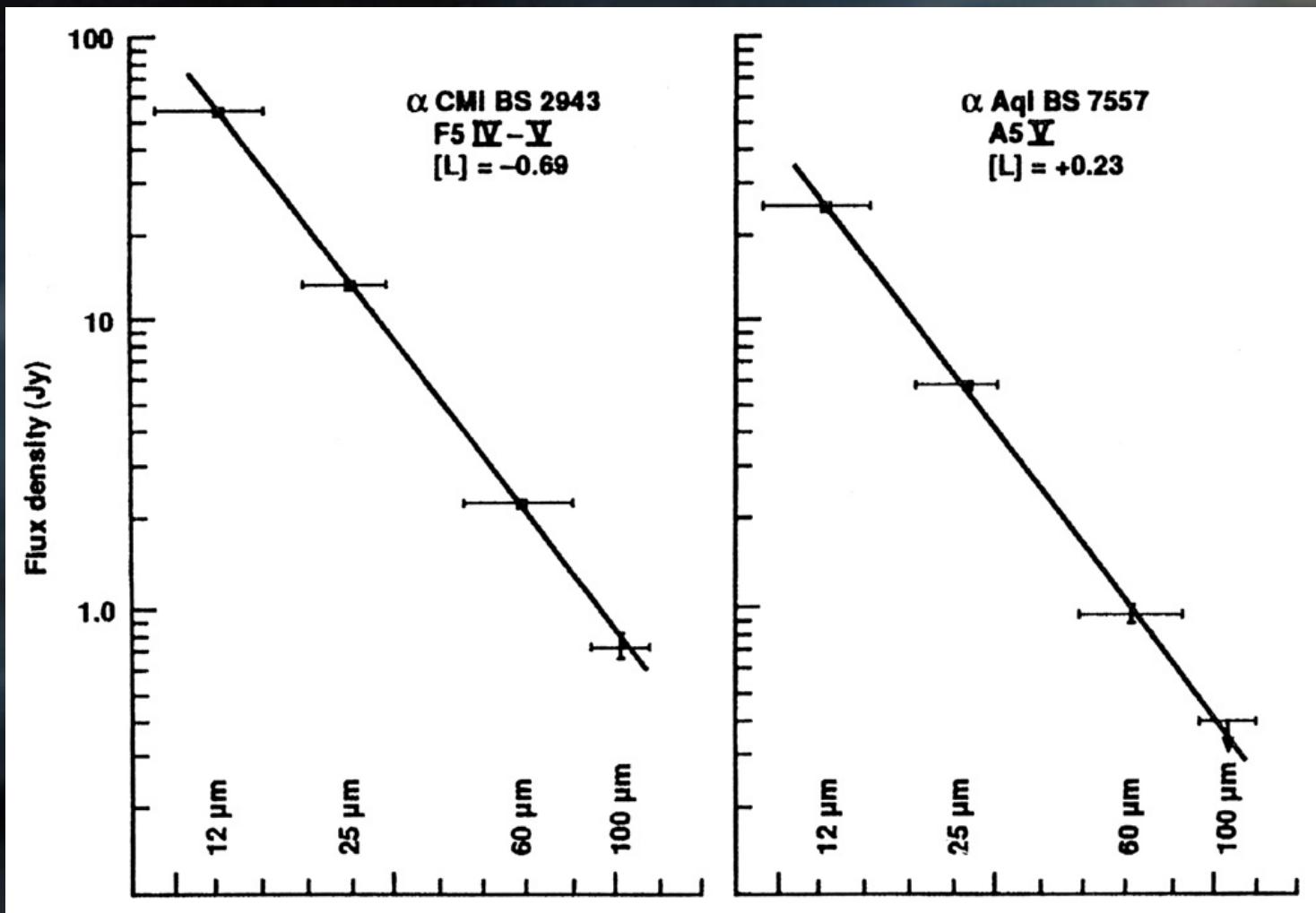
Debris disks

- Basic properties
- Observational methods & constraints
- Why "debris"?
- Structures

Alexis Brandeker, Lecture 6 in star formation mini-course

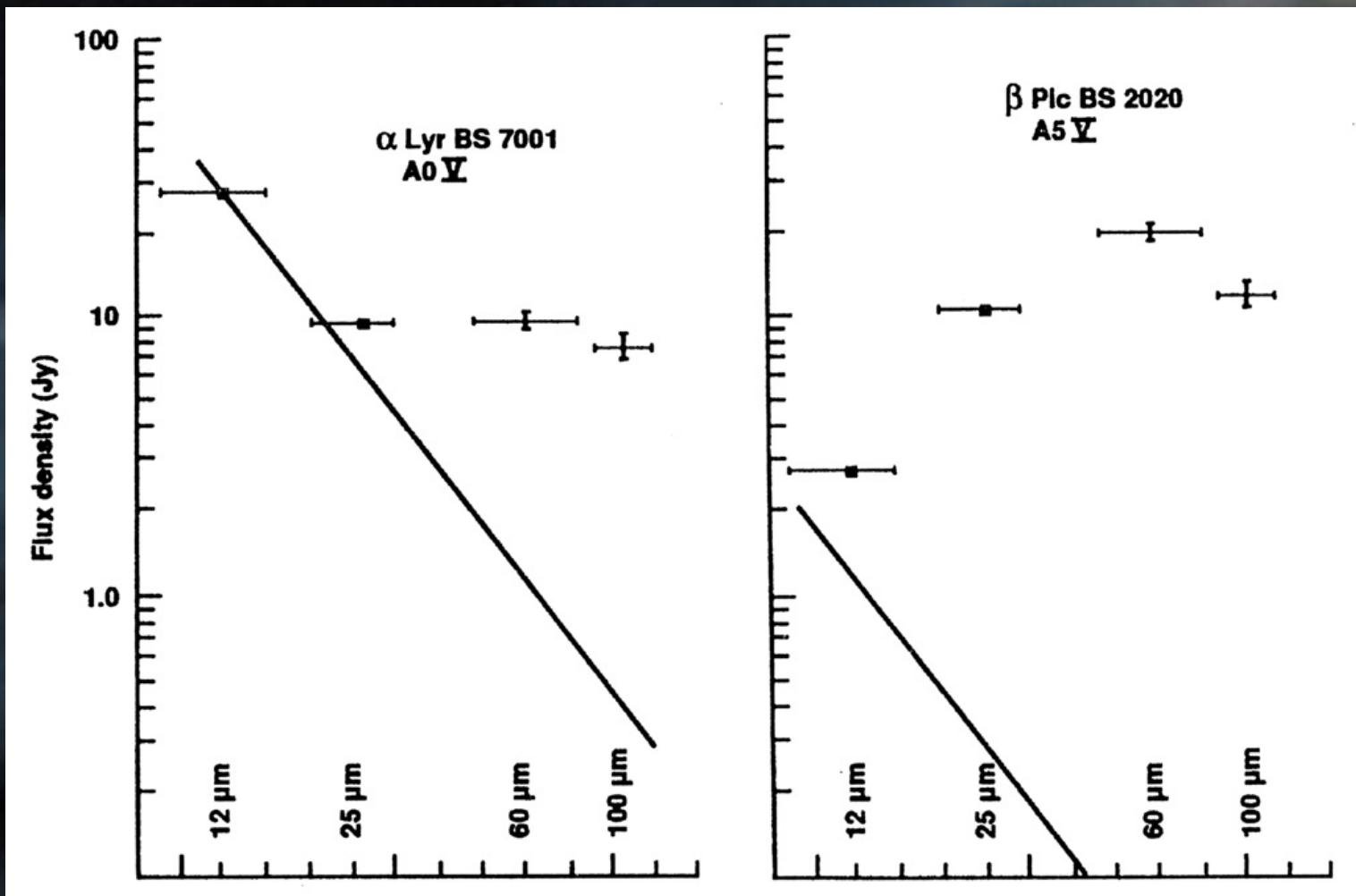
Discovery of debris disks

From Backman & Parece 1993, PPII, 1253

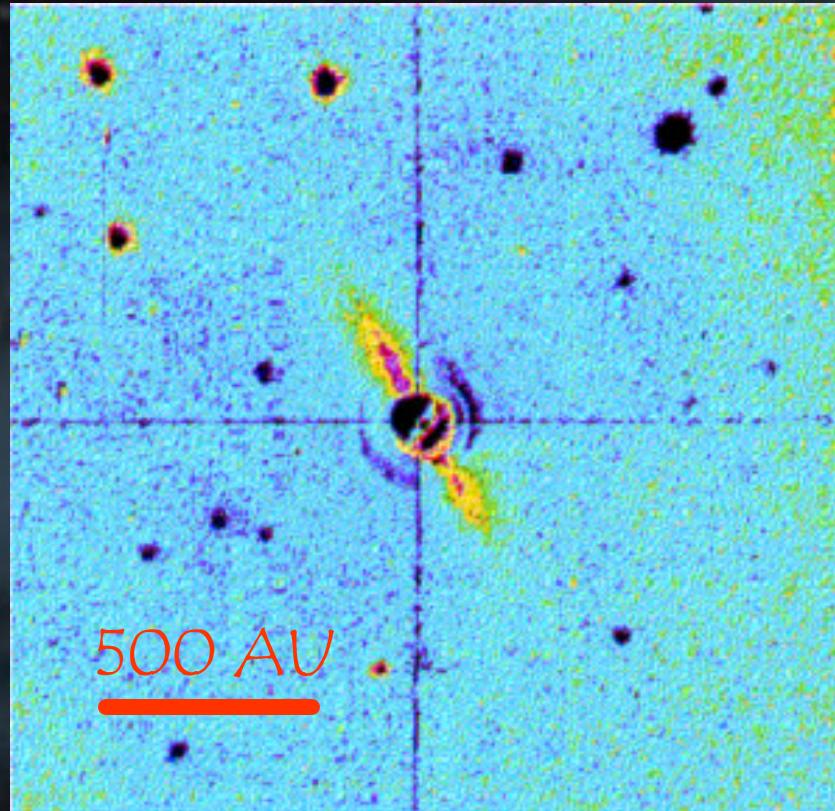


Discovery of debris disks

From Backman & Parece 1993, PPIII, 1253



First disk spatially resolved: β Pictoris

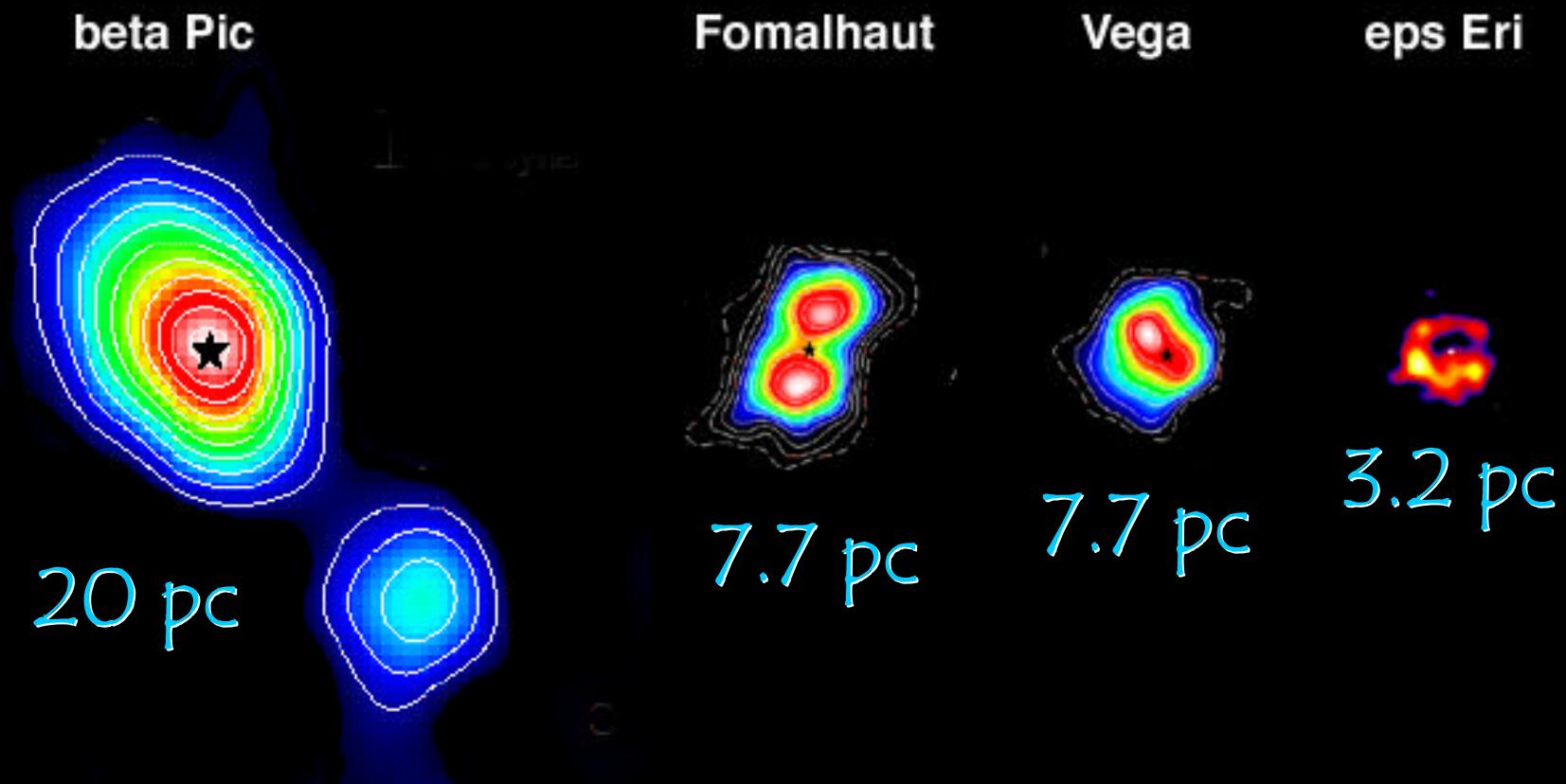


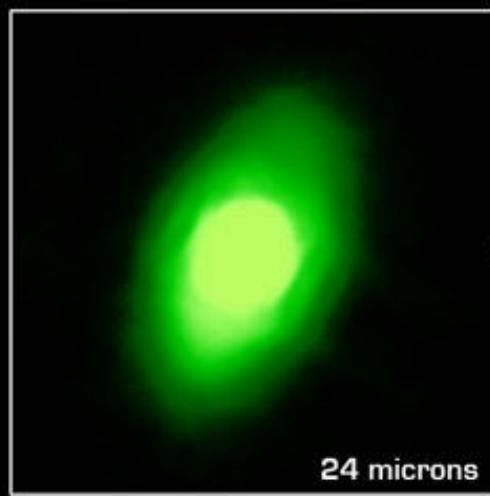
Discovery image by
Smith & Terrile 1984
(Sci. 226, 1421)

Properties

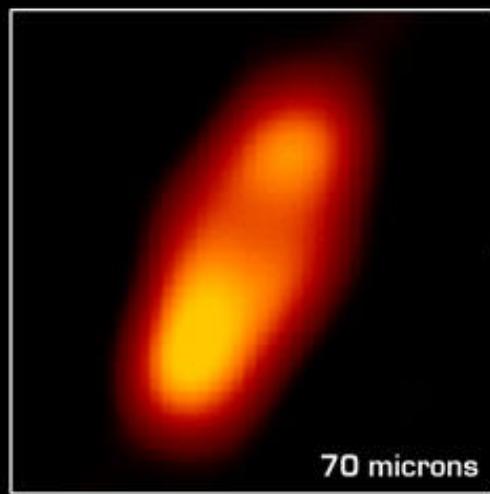
- Large (up to 1000 AU) dusty disks found around young main-sequence stars ~ 10 Myr – 1 Gyr
- Dust disk mass $M \sim 10^{-3}$ to a few M_{Earth} .
- Essentially free of gas
- Cold; typically 30 – 300 K
- Dominating dust emission from μm to mm sized grains

"The fabulous four" at 850μm

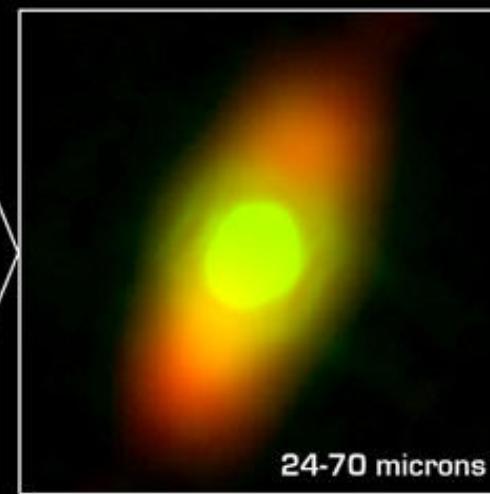




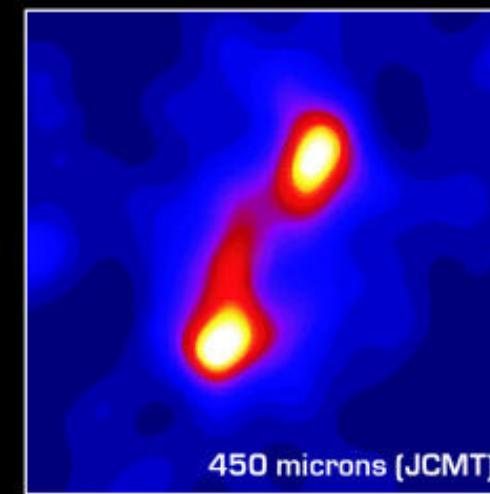
24 microns



70 microns



24-70 microns



450 microns (JCMT)

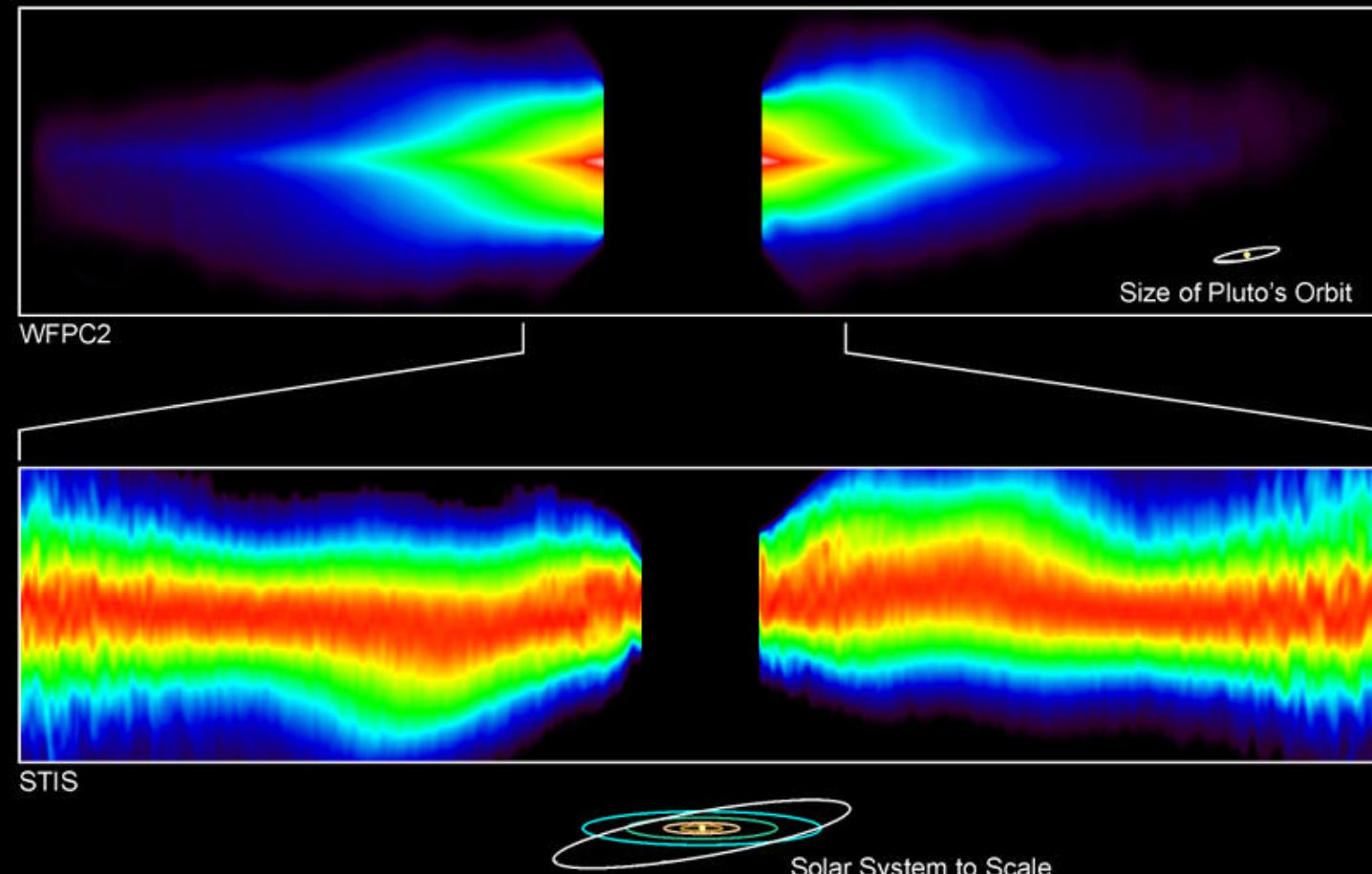
Fomalhaut Circumstellar Disk

NASA / JPL-Caltech / K. Stapelfeldt (JPL)

Spitzer Space Telescope • MIPS

ssc2003-06i

β Pictoris in scattered light

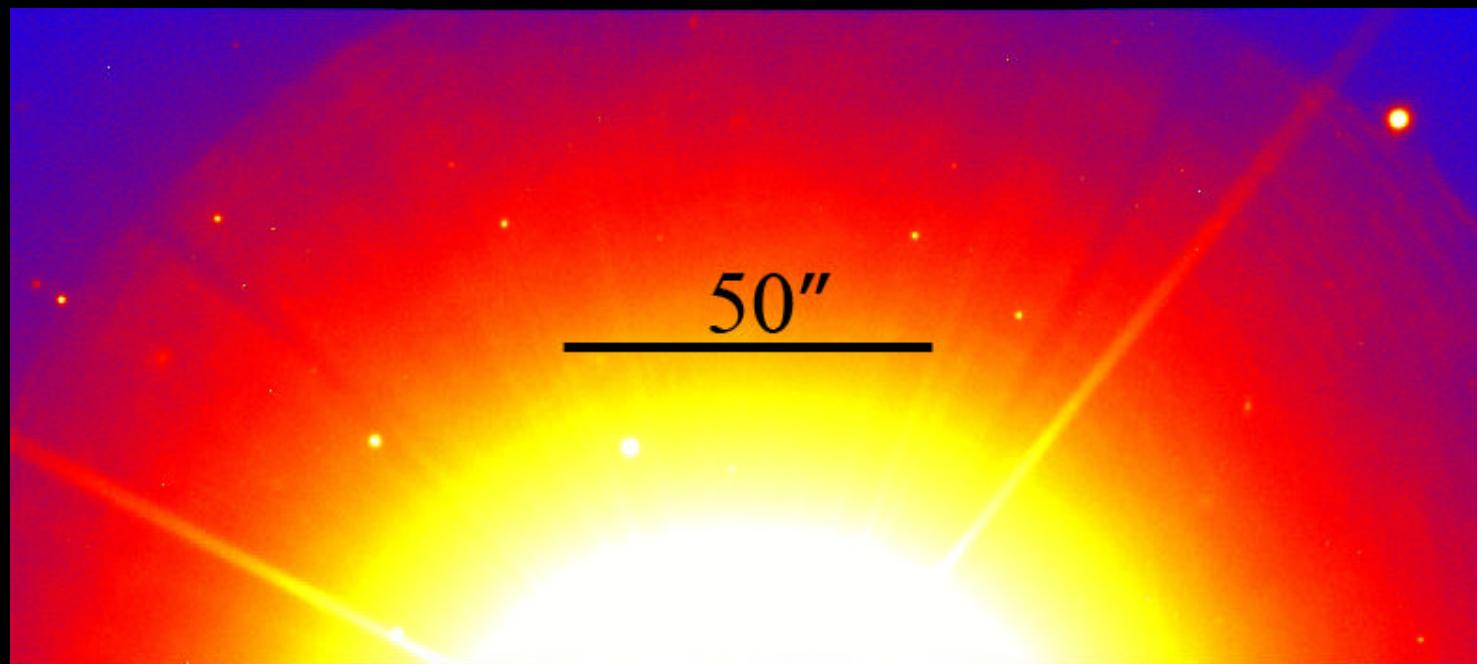


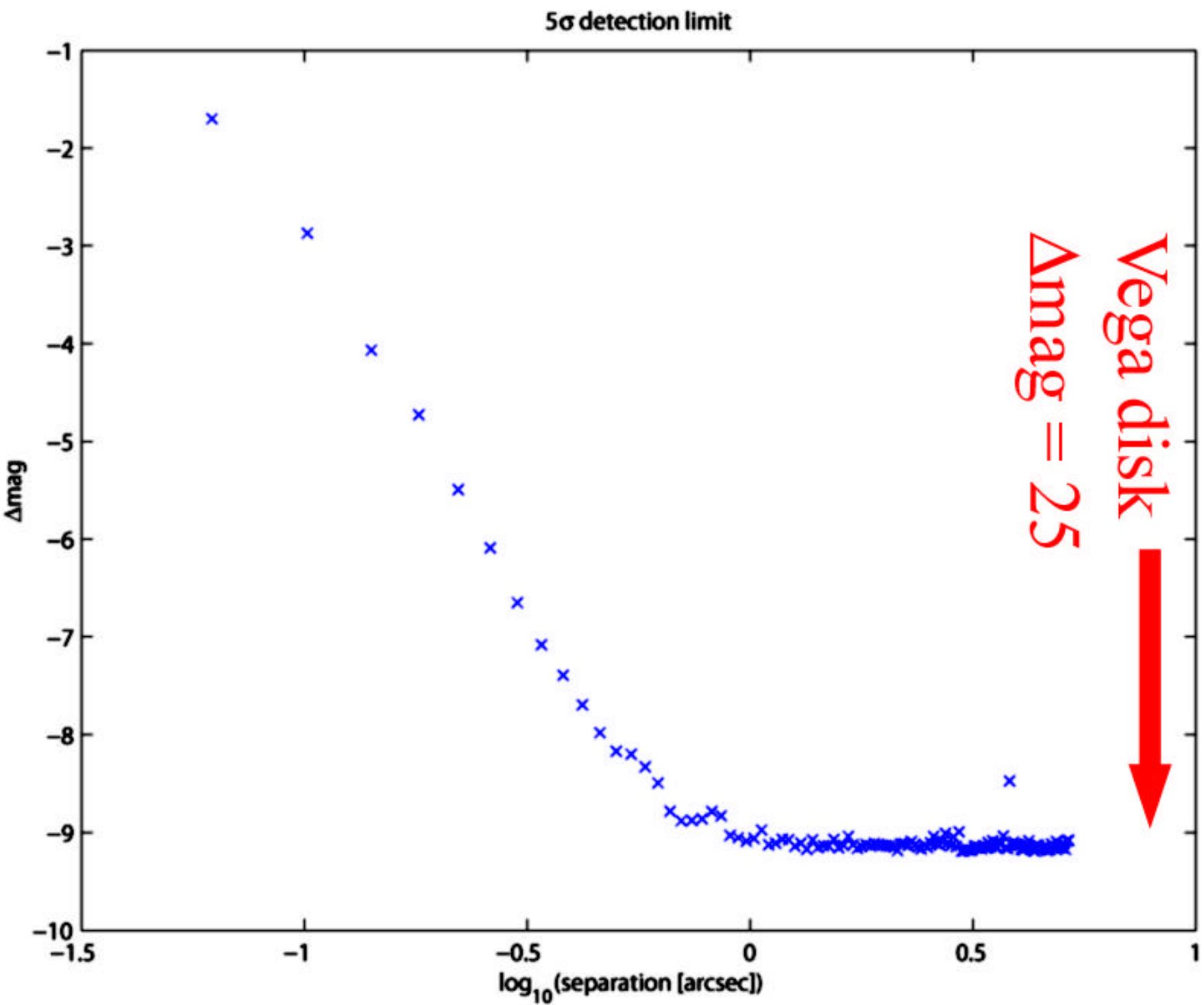
Beta Pictoris

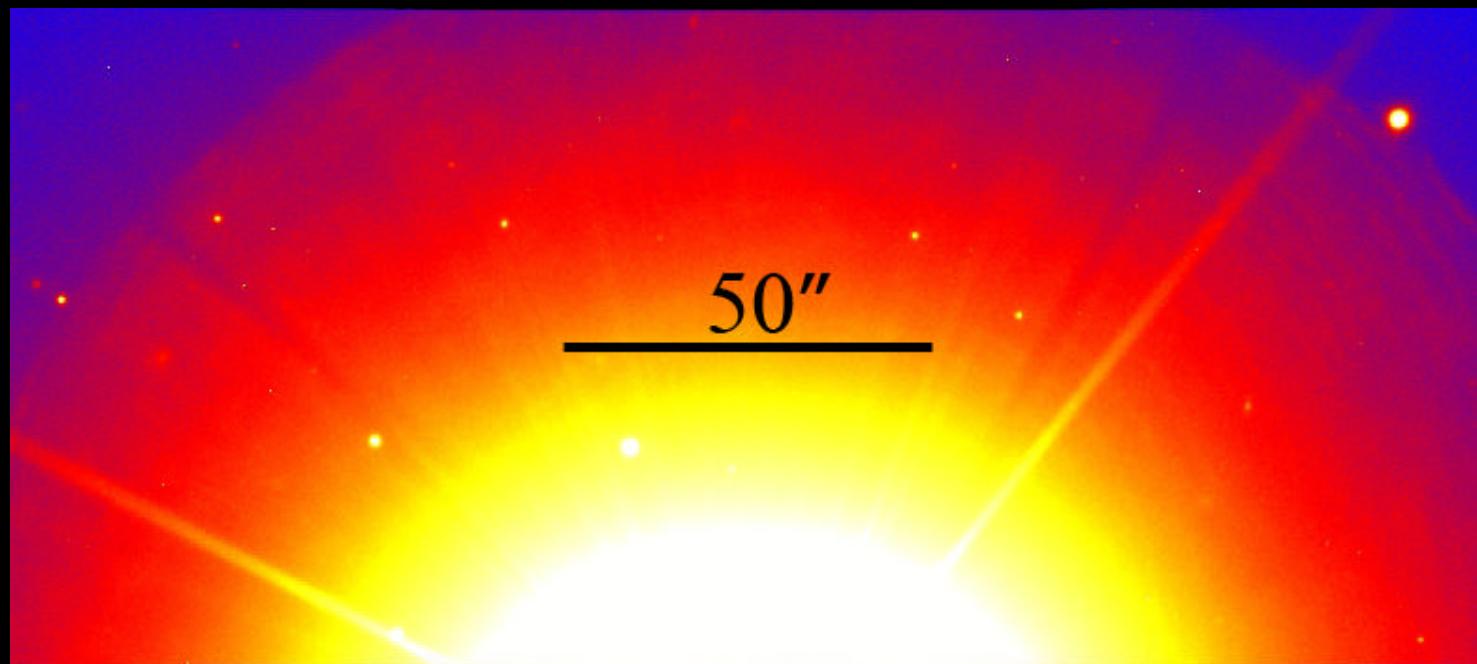
PRC98-03 • January 8, 1998 • ST Scl OPO

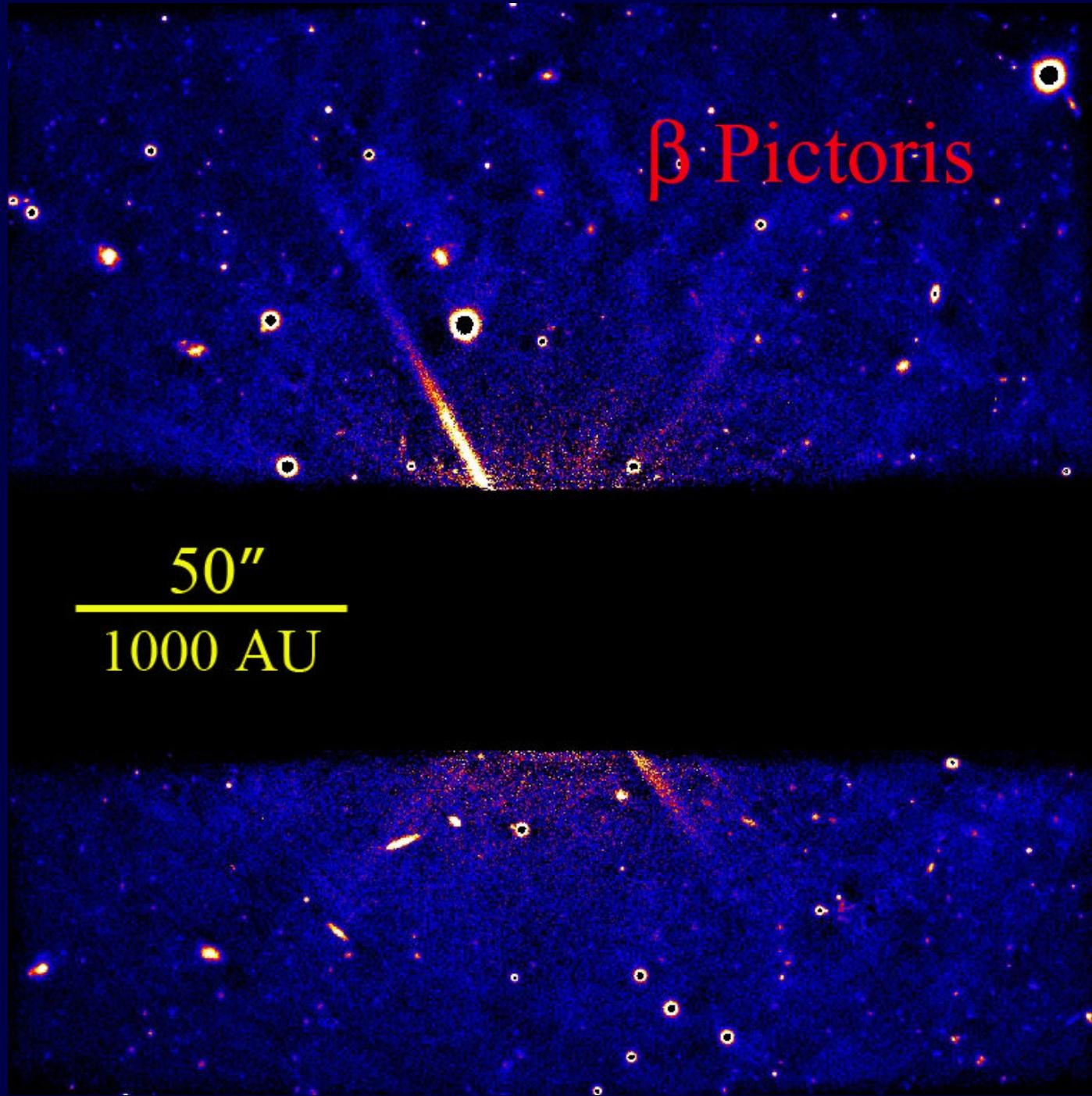
A. Schultz (Computer Sciences Corp.), S. Heap (NASA Goddard Space Flight Center) and NASA

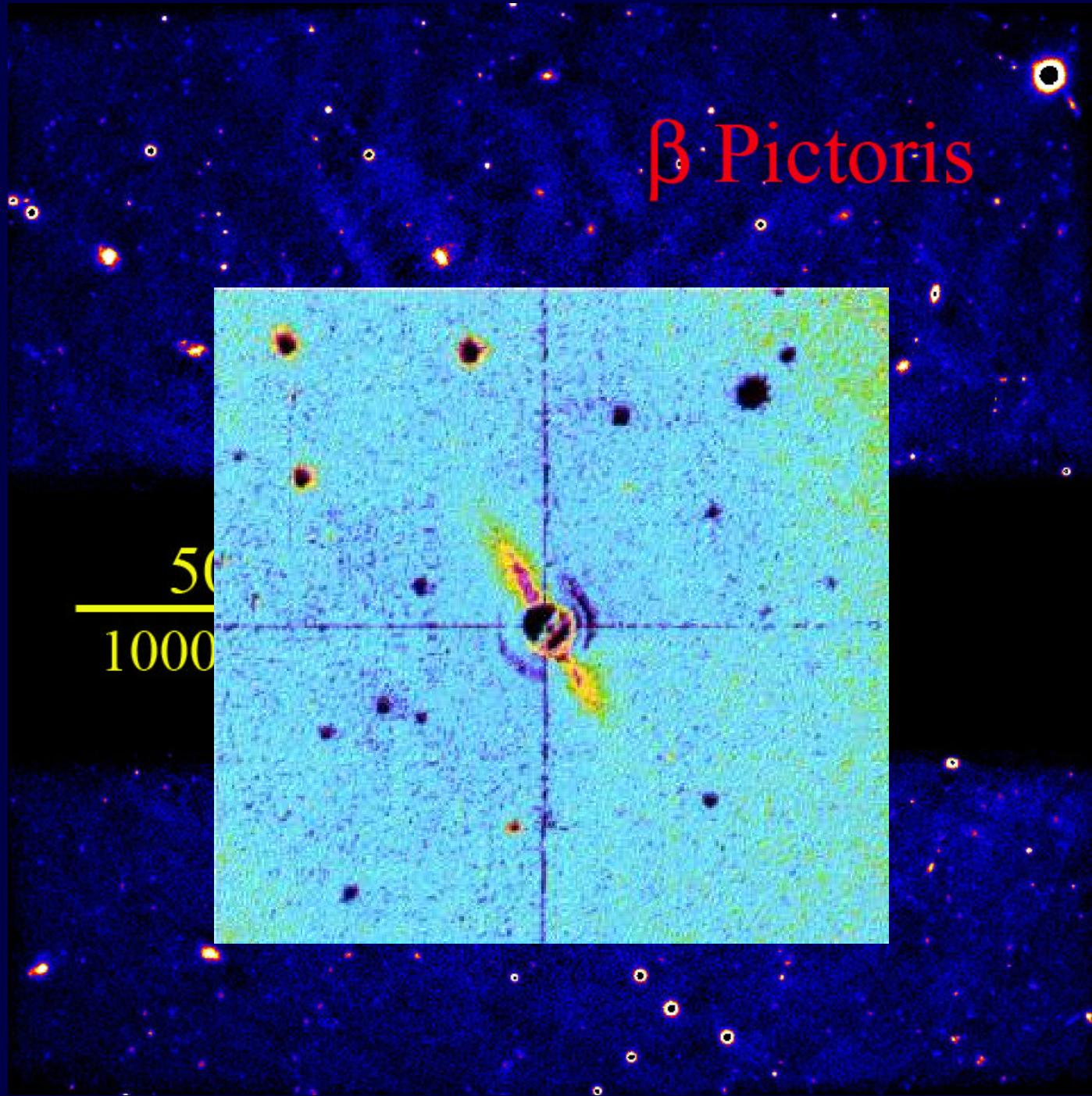
HST • WFPC2 • STIS

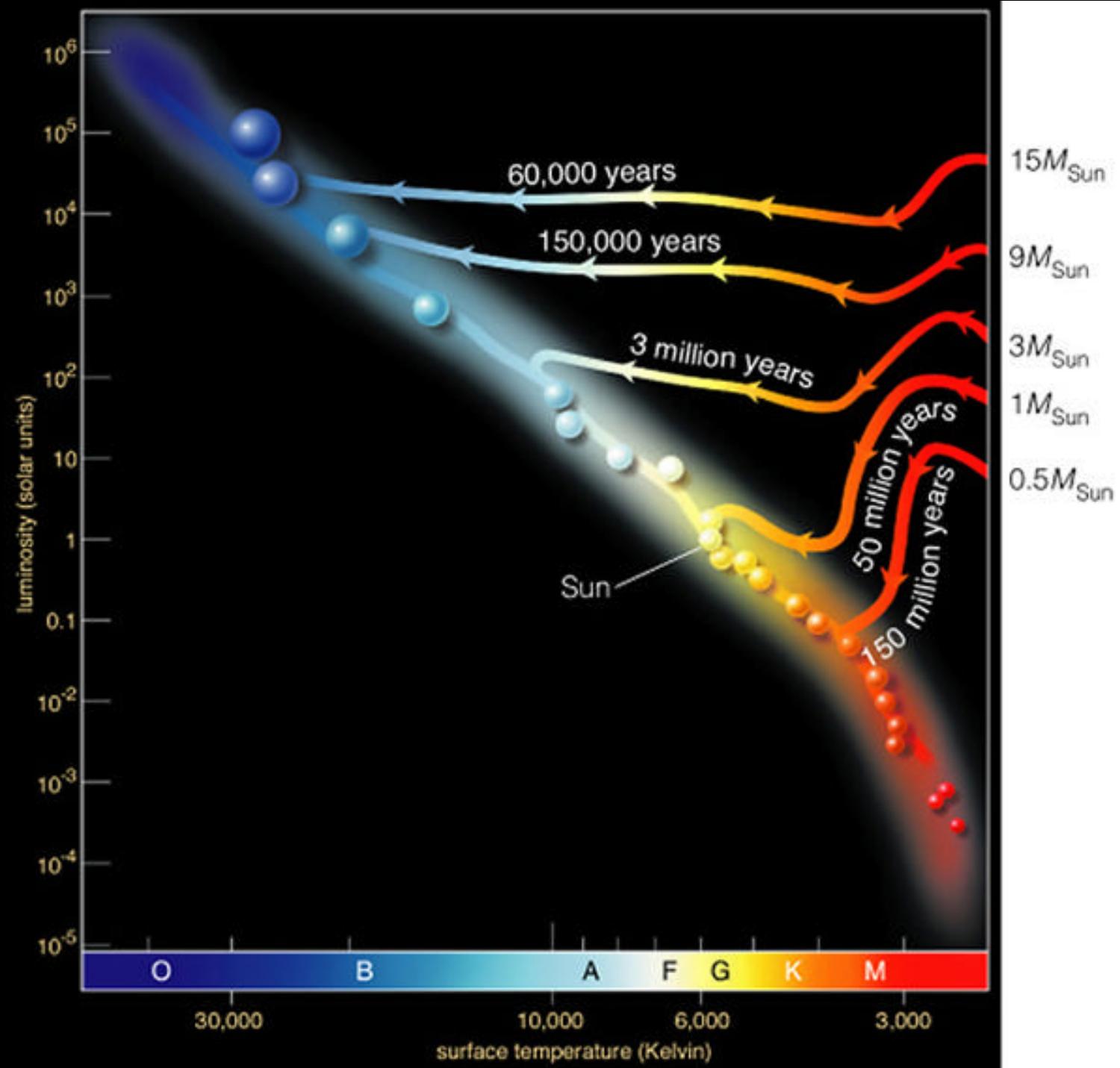




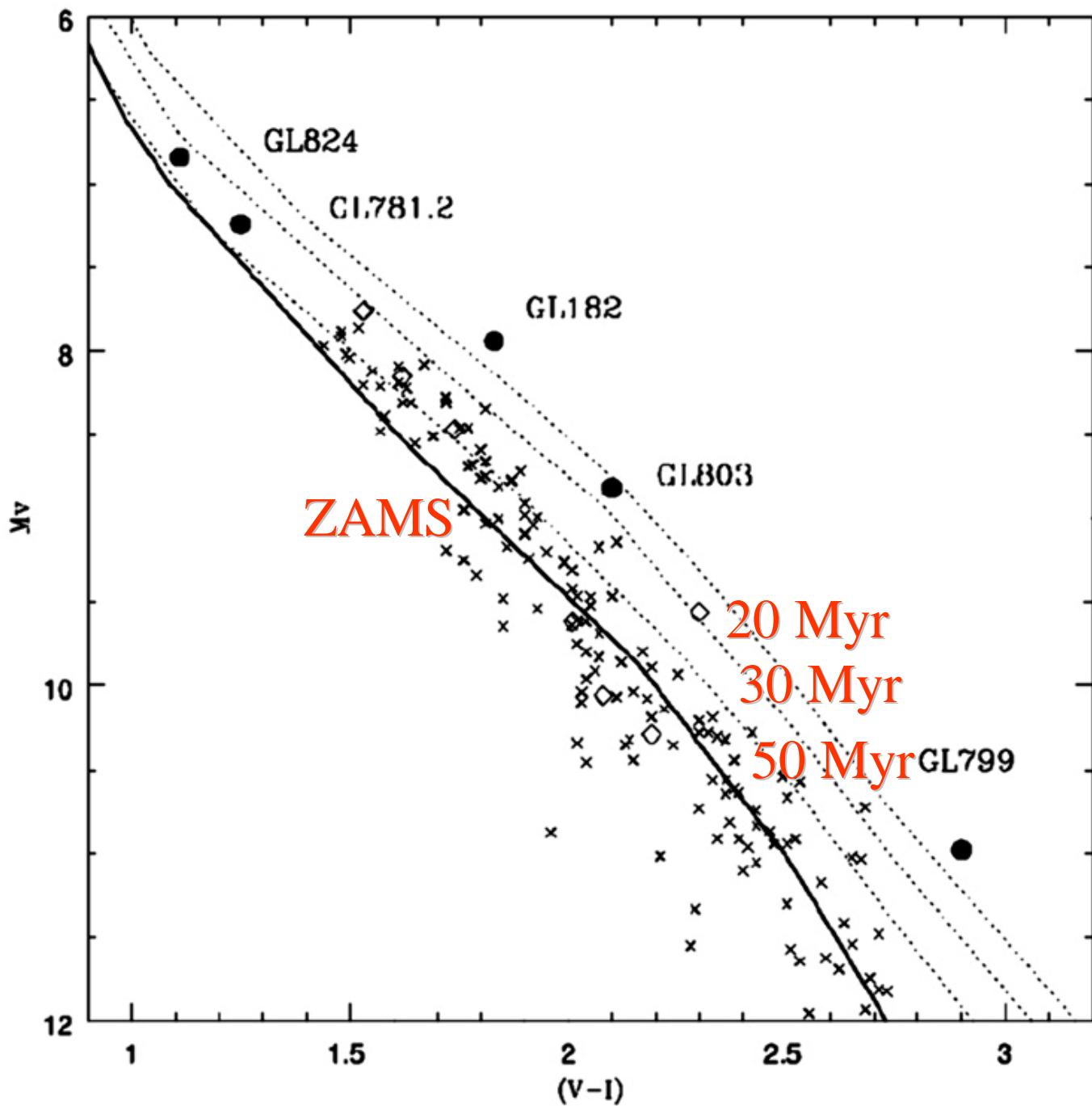




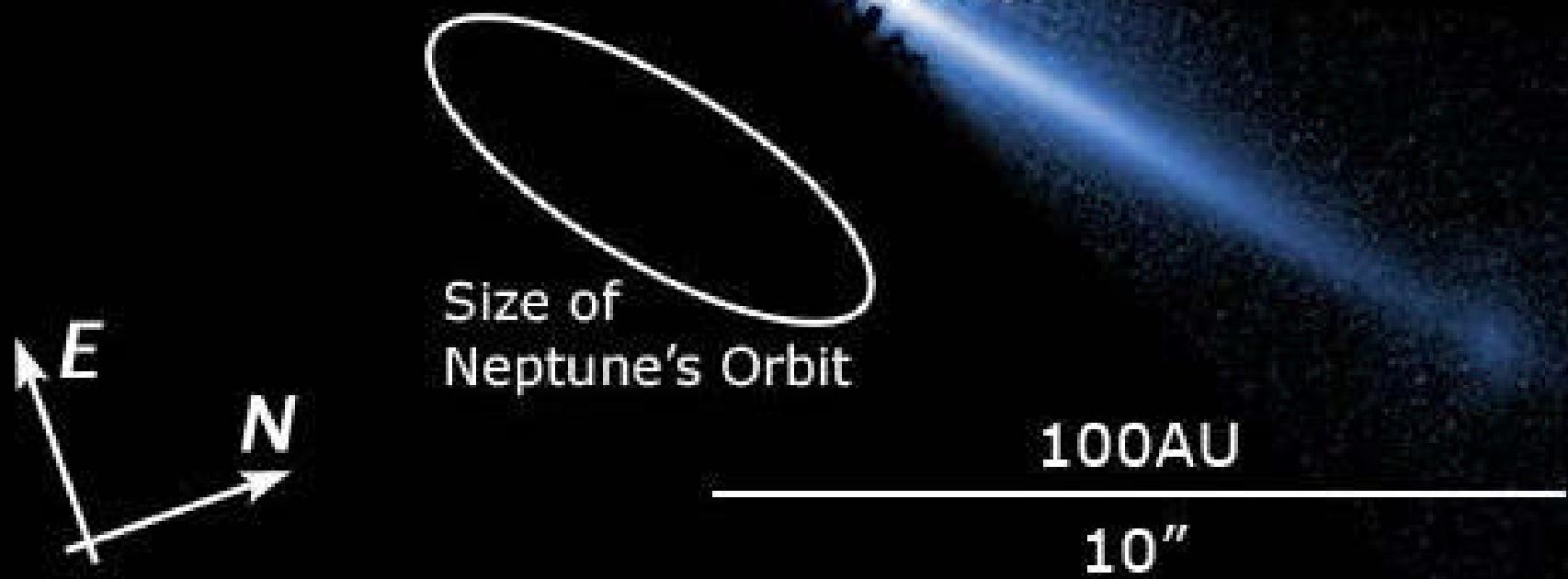




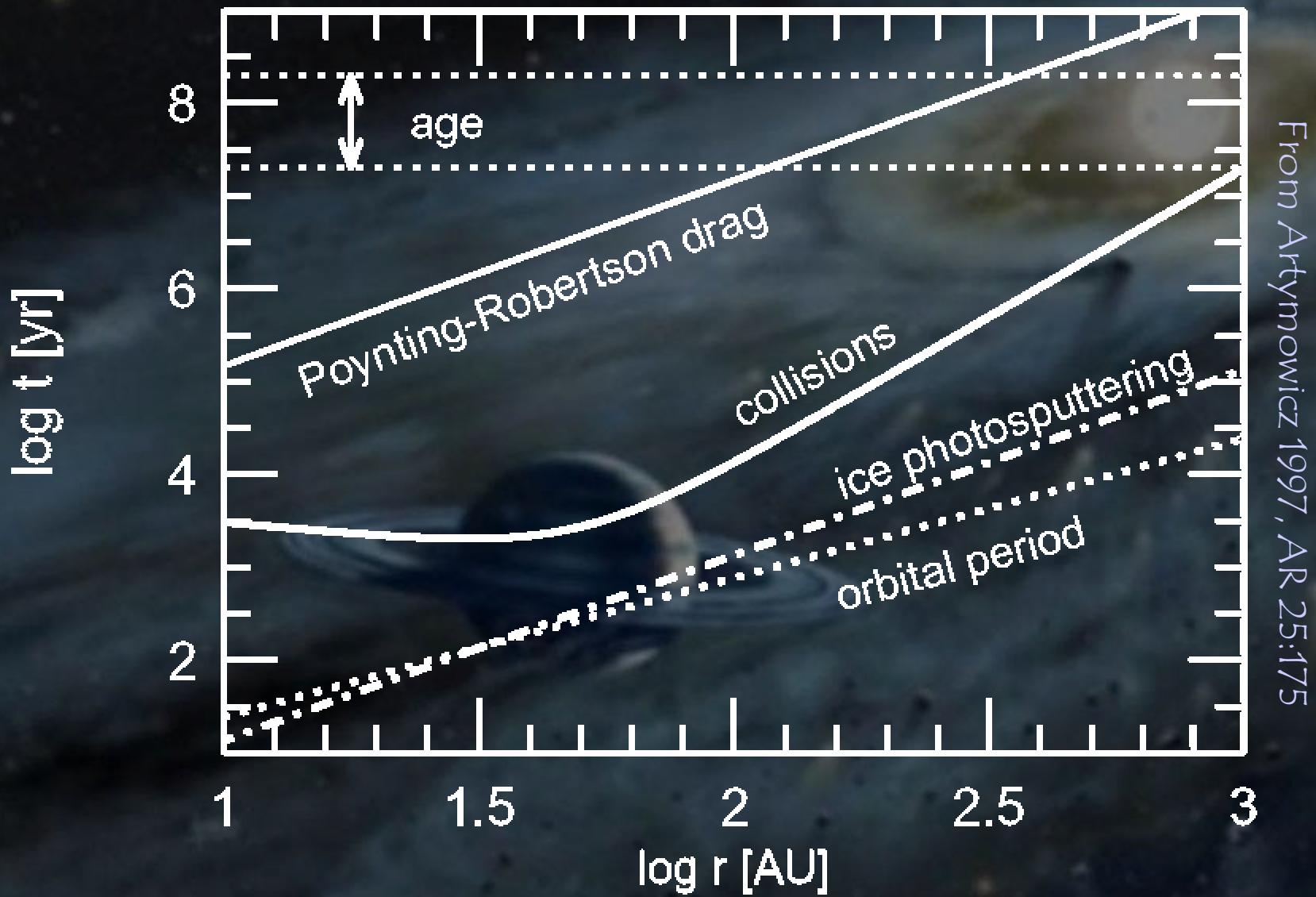
Barrado y Navascués 1999



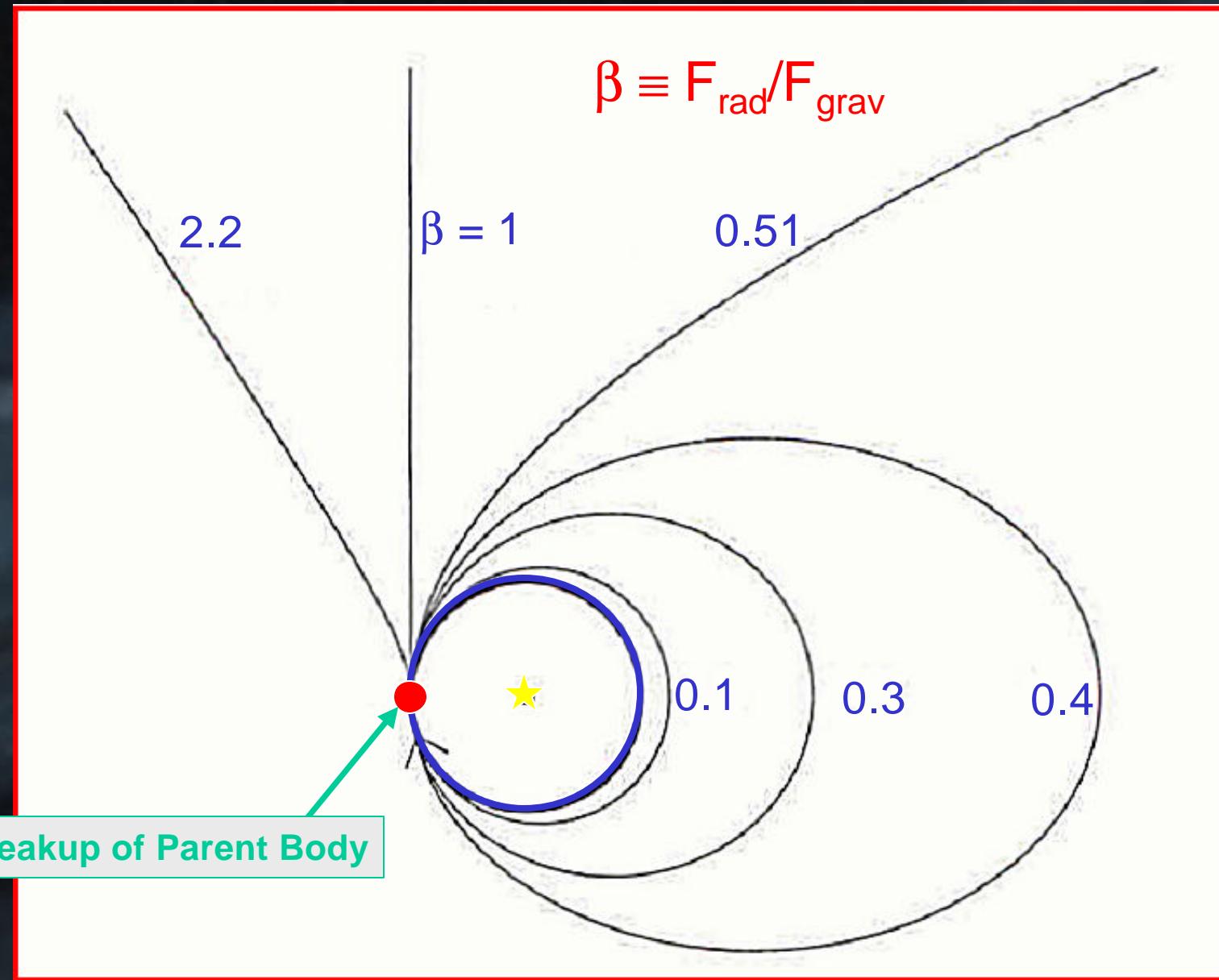
AU Mic
HST ACS/HRC
J. Krist (STScI/JPL)



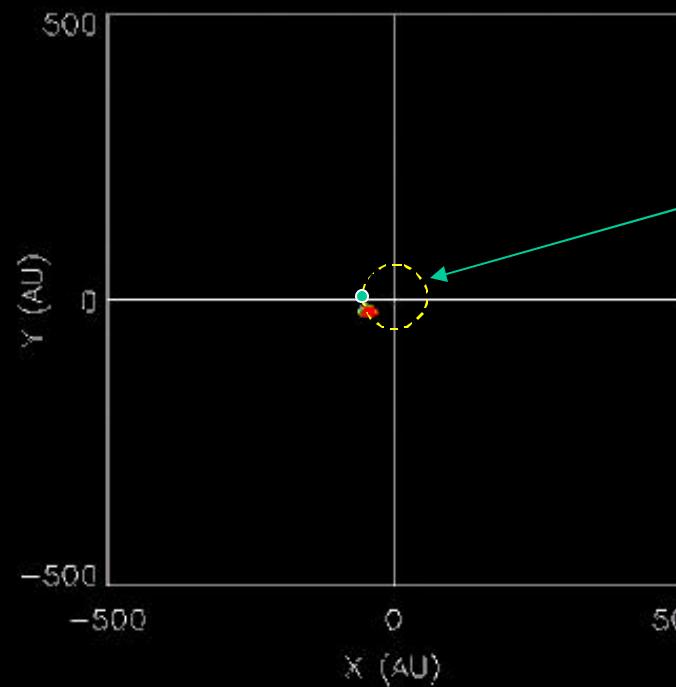
Why "debris"?



From Artymowicz 1997, AR 25:175

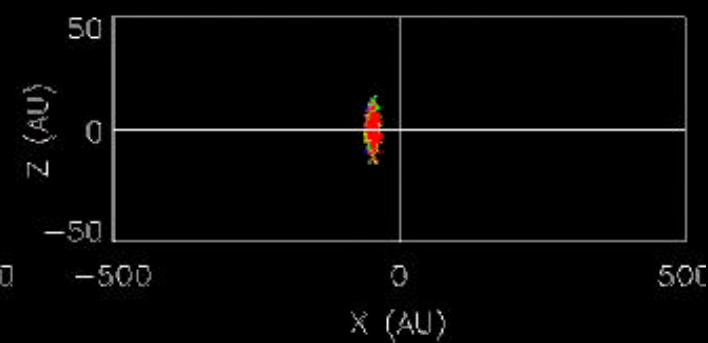


Top view



Orbit of Pluto-mass
parent body before breakup

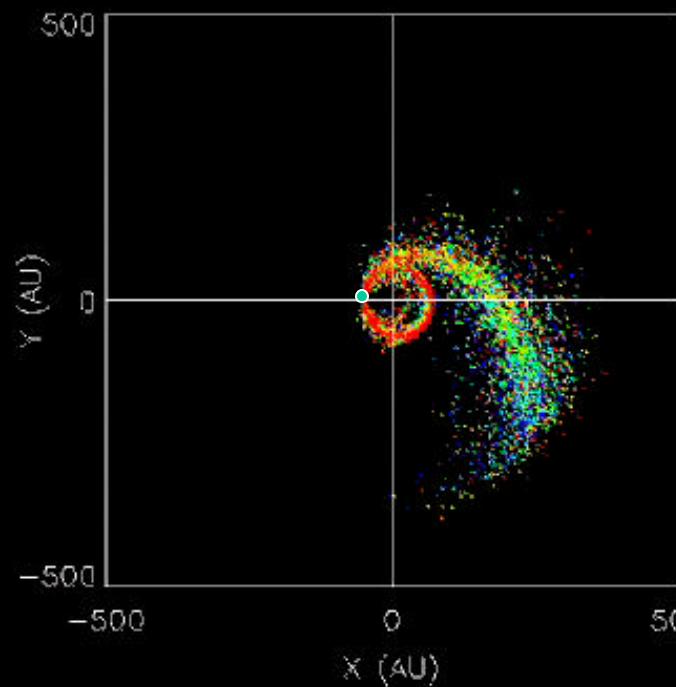
Side view



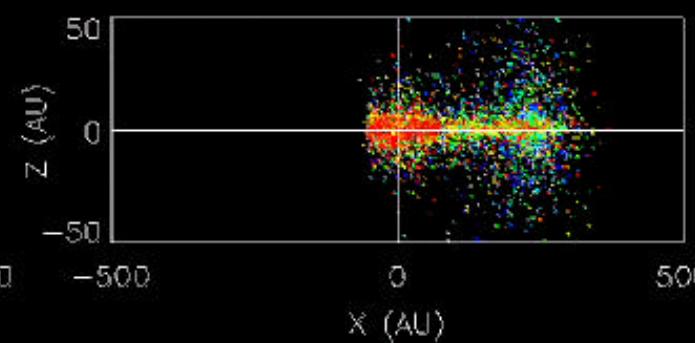
$\beta = 0.1 \quad 0.2 \quad 0.3 \quad 0.4 \quad 0.5$

Elapsed time: 21 years

Top view



Side view



$\beta = 0.1 \quad 0.2 \quad 0.3 \quad 0.4 \quad 0.5$

Elapsed time: 569 years

Subaru/COMICS

(Y.Okamoto et al. 2004)

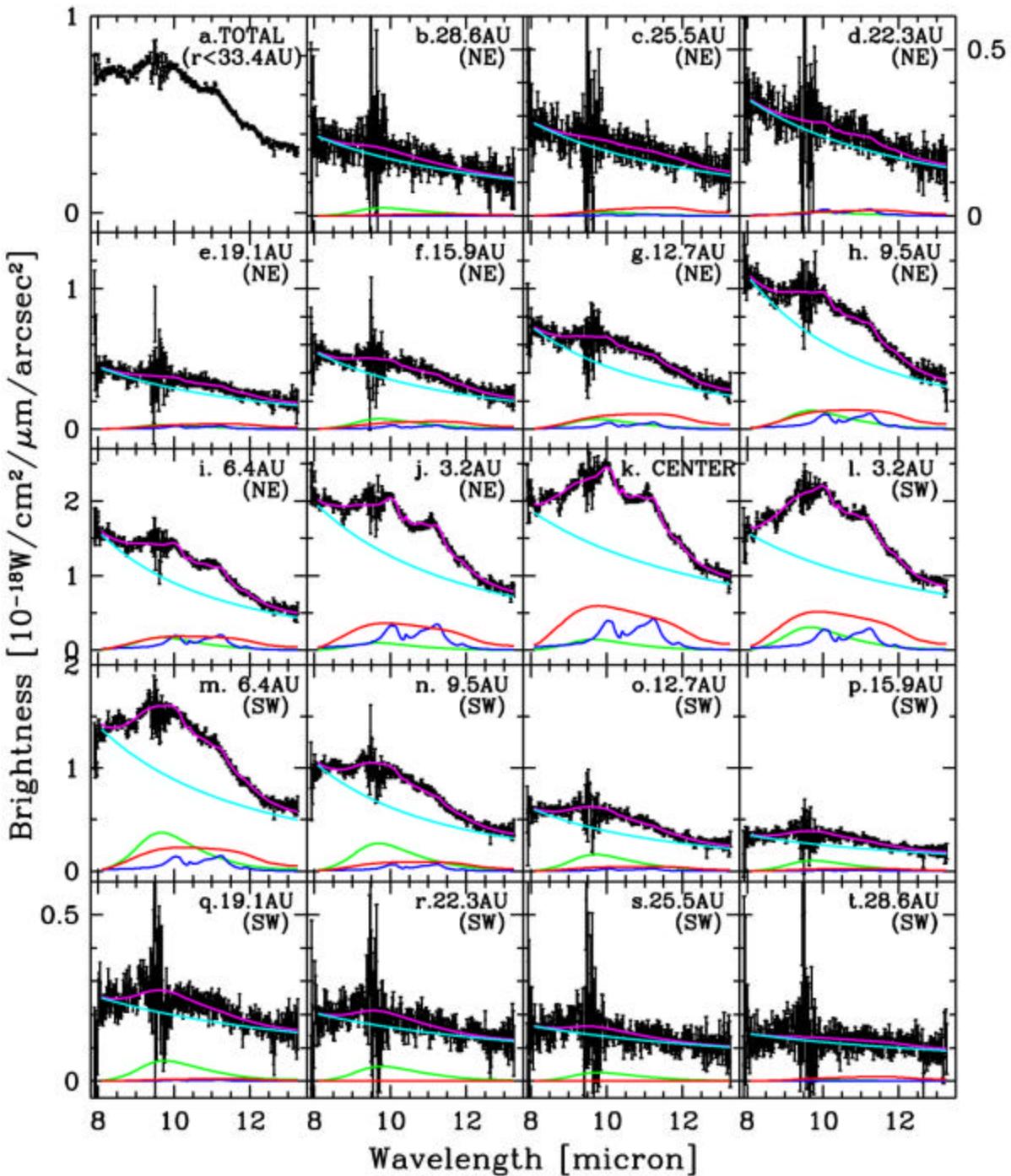
Total (magenta)

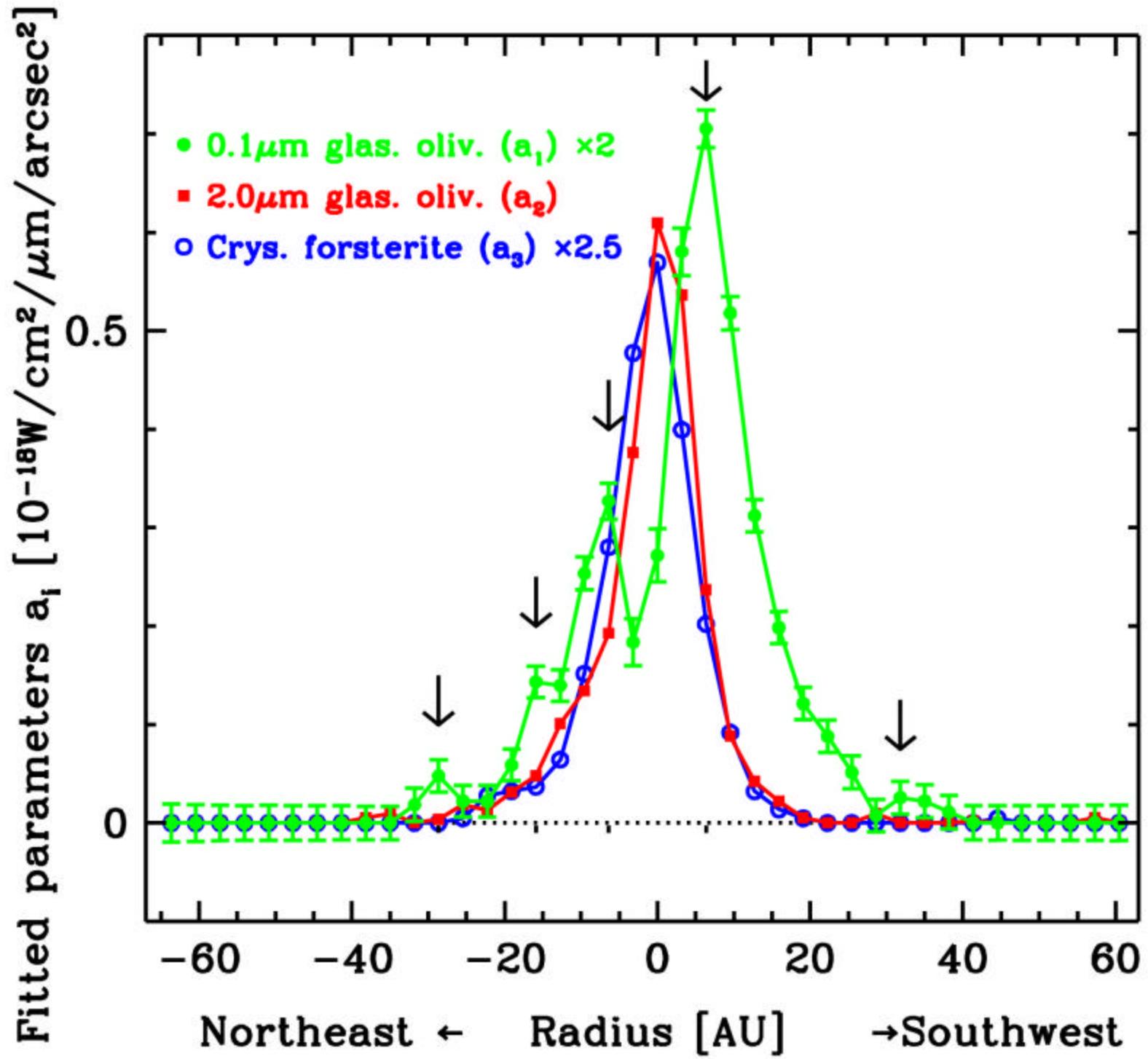
0.1mm amorphous
olivine (green)

2mm amorphous olivine
(red)

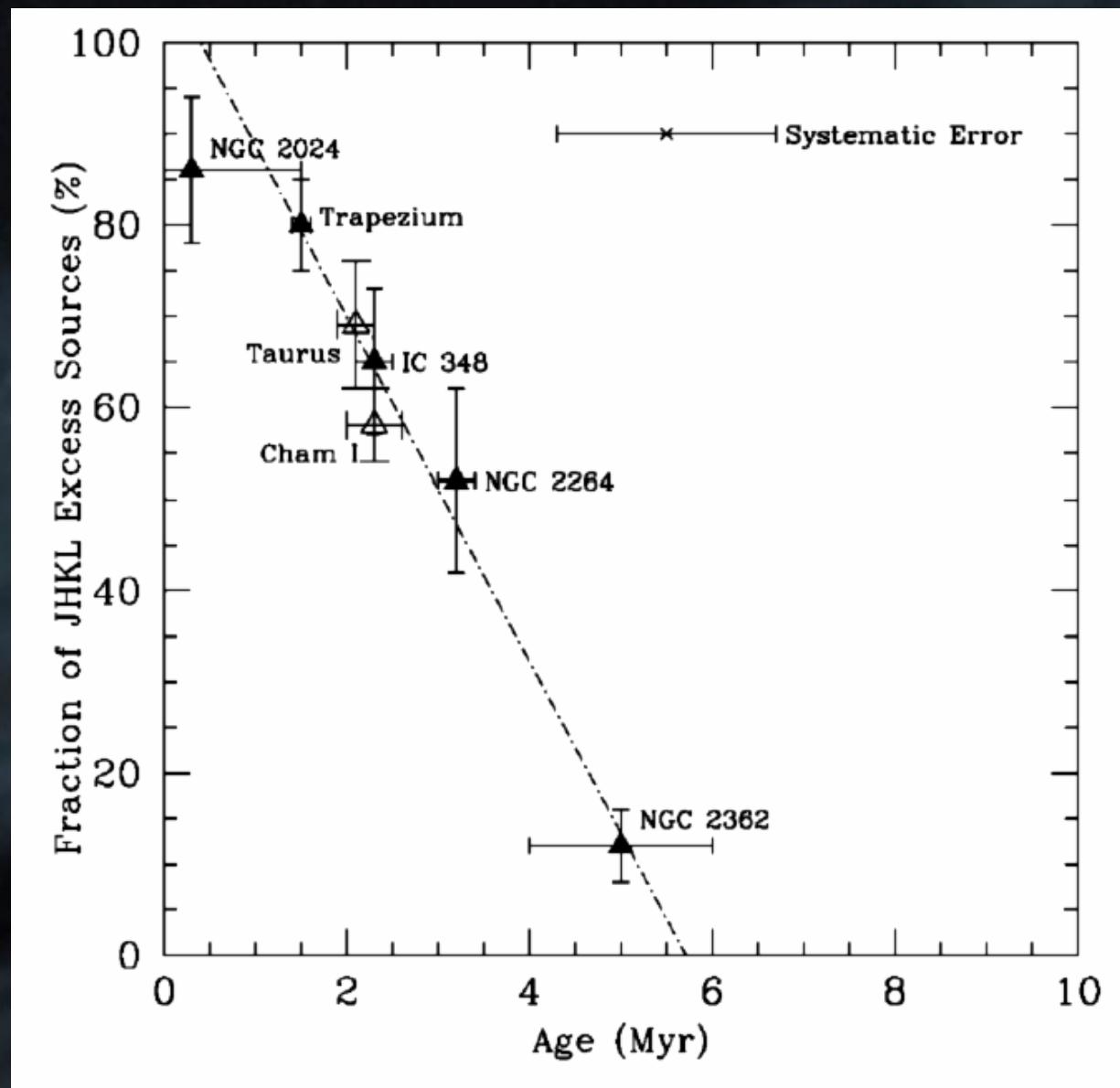
Crystalline forsterite
(blue)

Power-law continuum
(cyan)





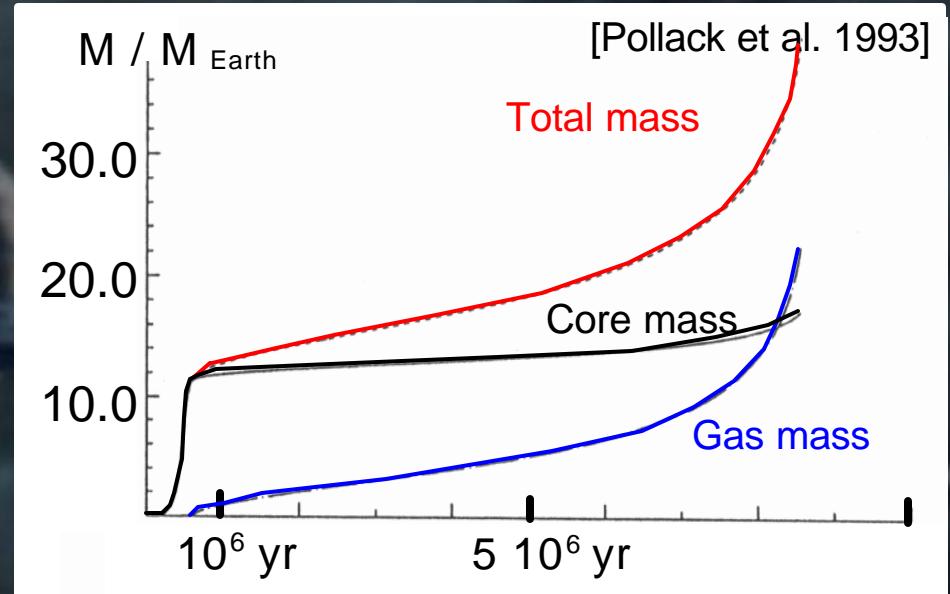
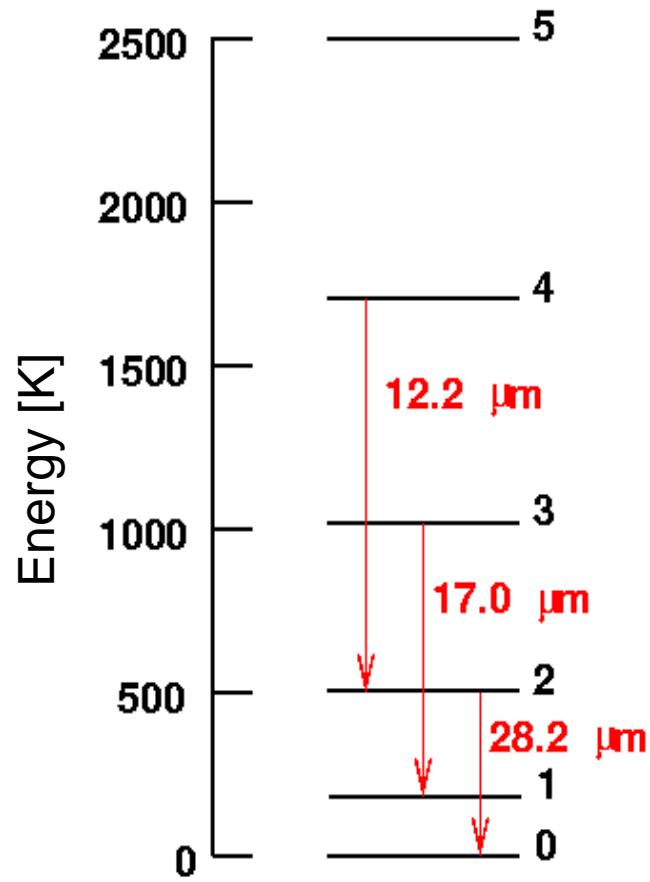
Dust disk lifetimes



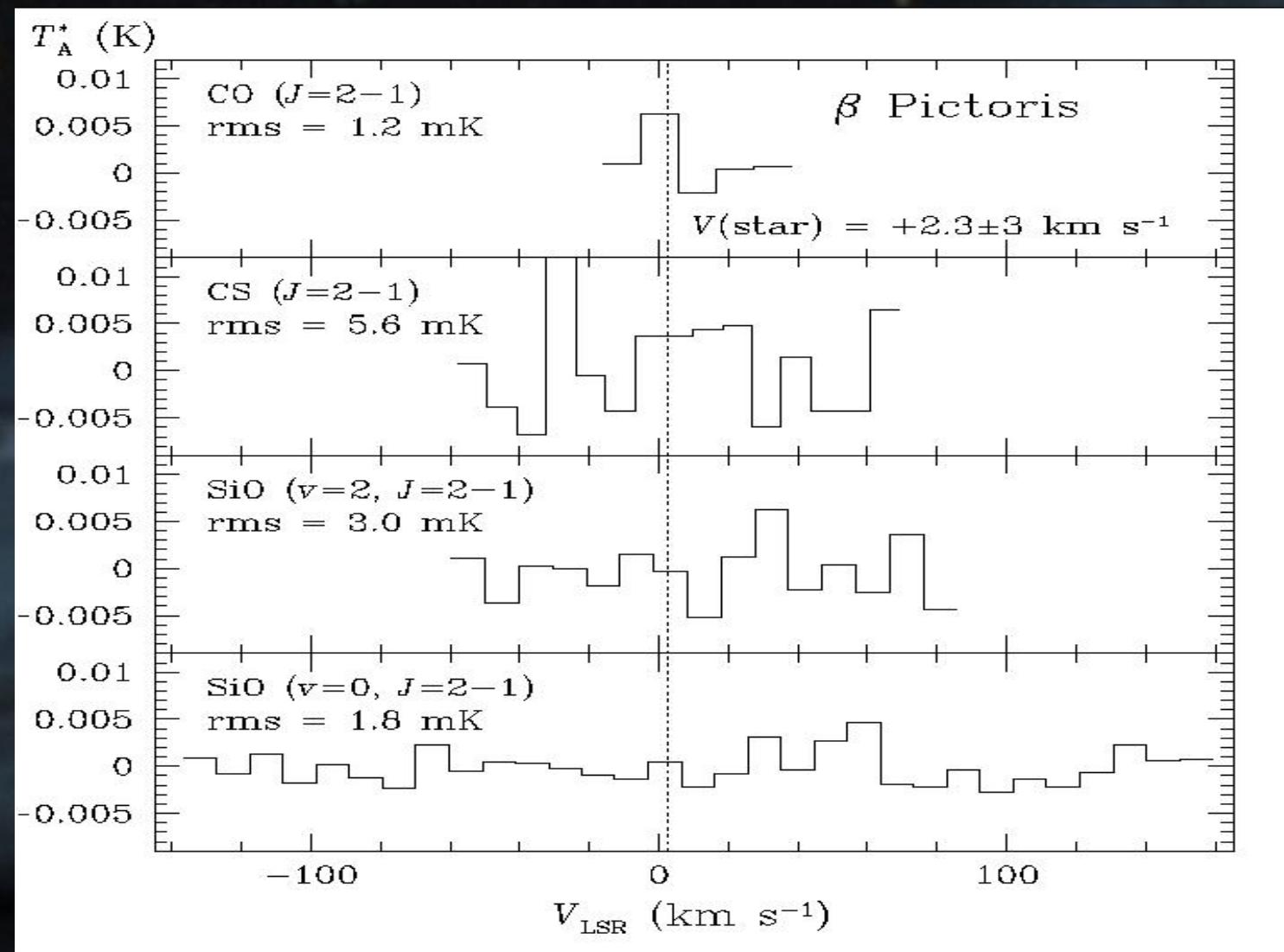
From Haisch, Lada & Lada 2001, ApJ, 553, 153

What about gas?

Energy diagram of H_2

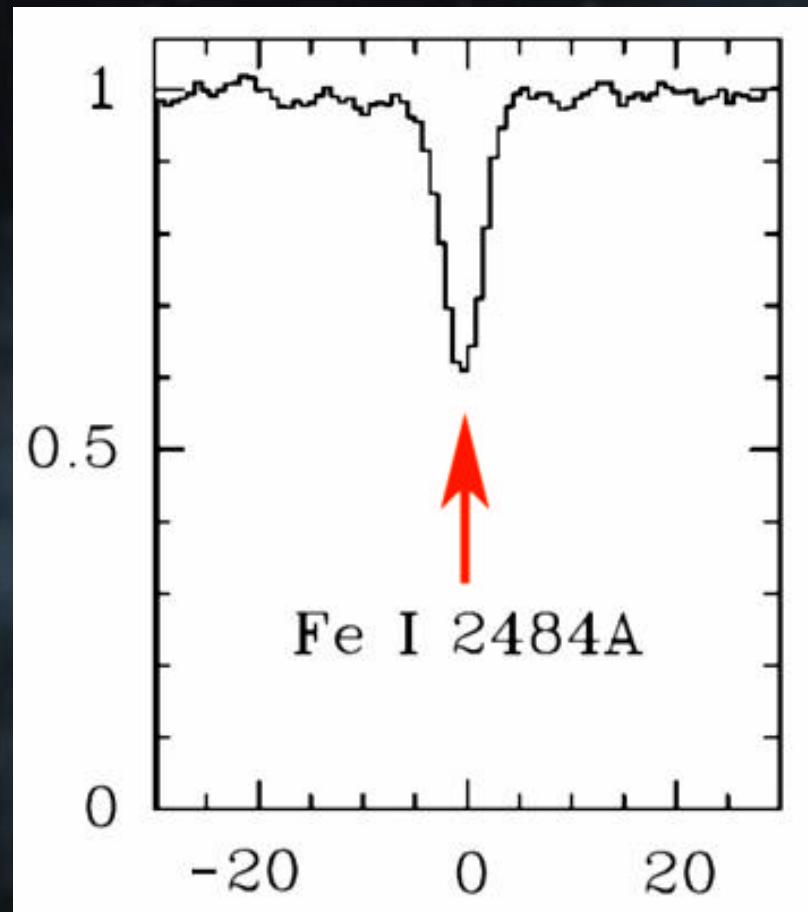


mm/submm observations



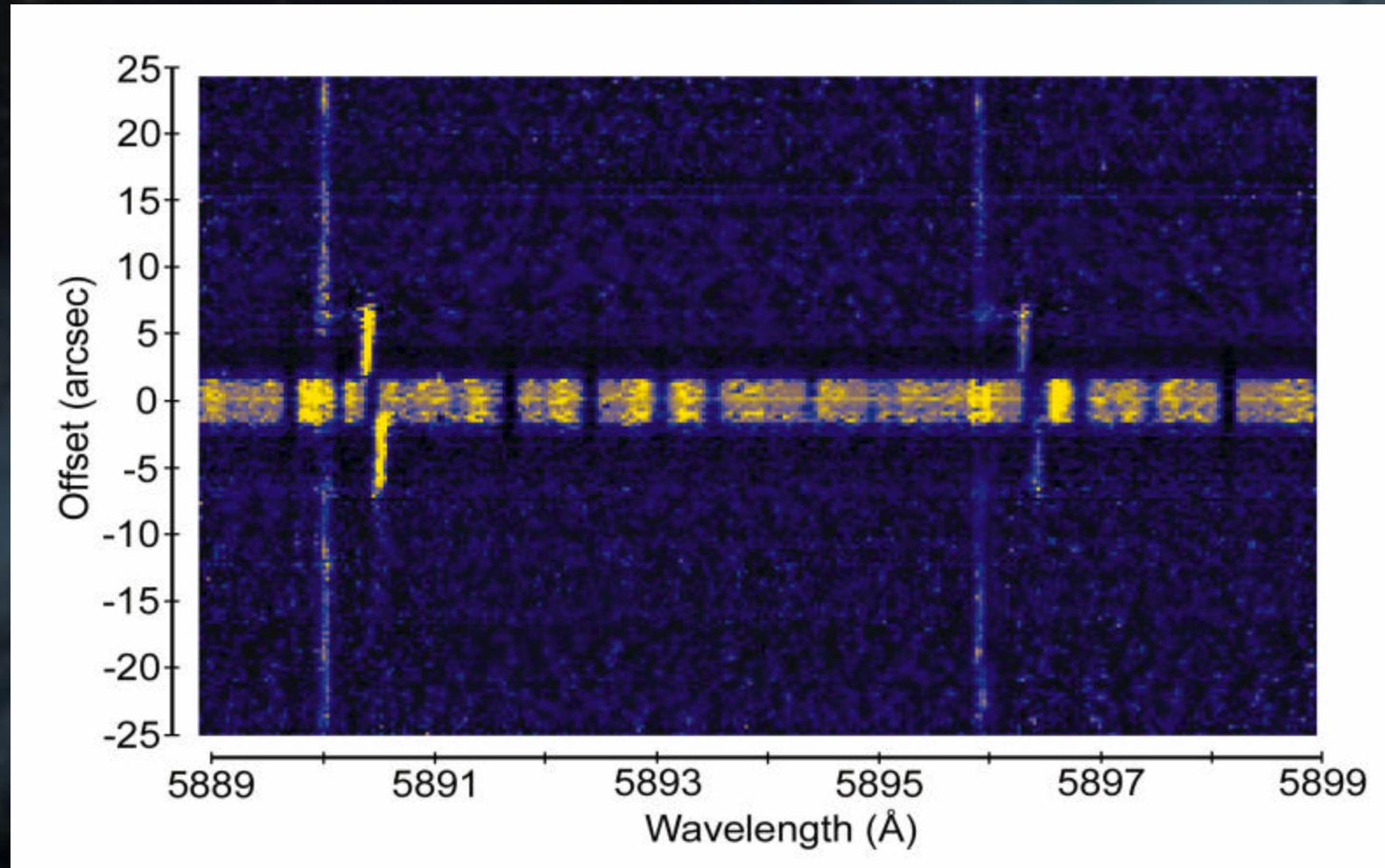
Liseau & Artynowicz 1998, A&A 335, 935

Edge-on P absorption!



| Ion | Δv km s^{-1} | $\sigma_{\Delta v}$ km s^{-1} |
|------|----------------------------------|---|
| FeI | -0.0 | 0.3 |
| NaI | -1.2 | 0.3 |
| CaII | -0.3 | 2.1 |
| TiII | 0.2 | 0.8 |
| NiI | 0.4 | 0.4 |
| NiII | 2.8 | 3 |
| CrII | 2.2 | 3 |

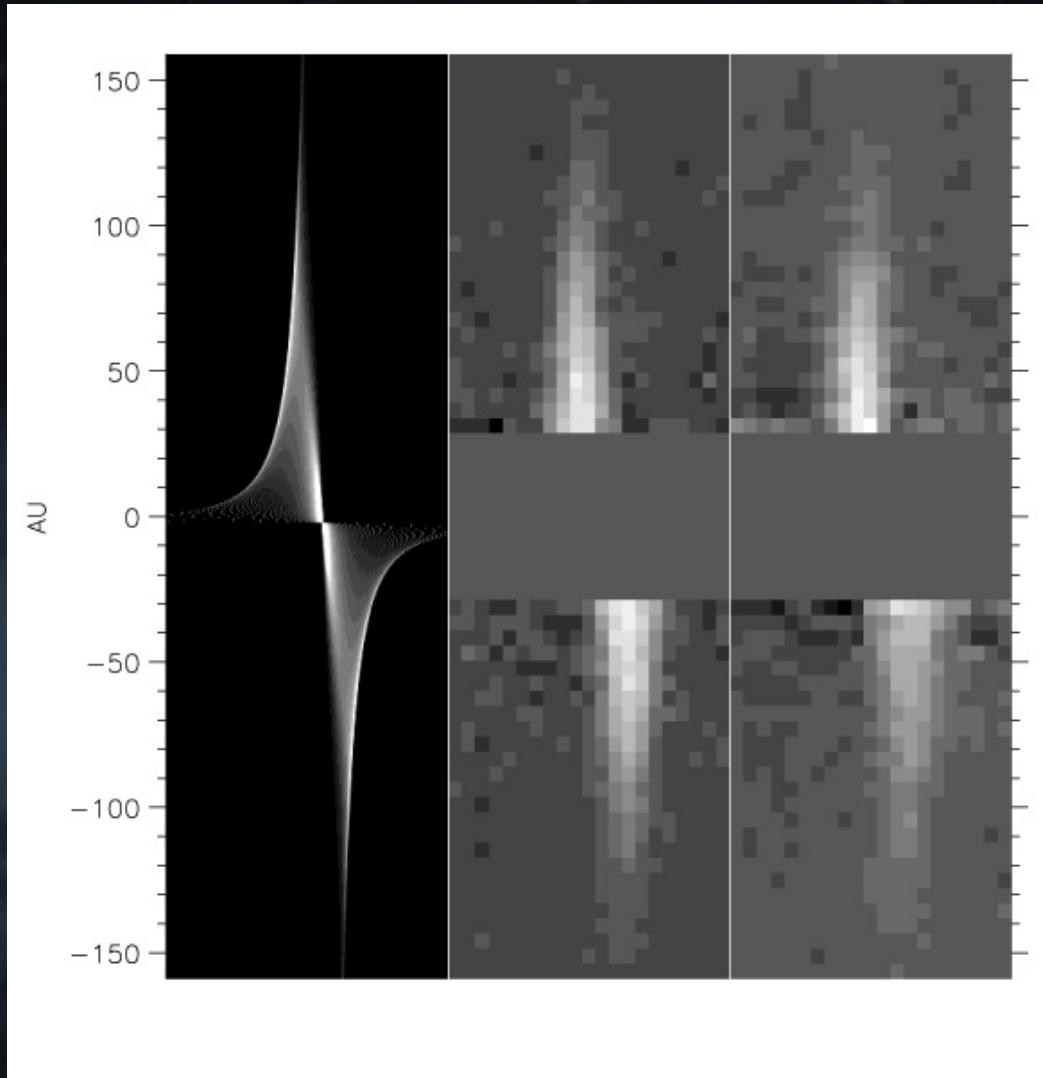
Gas in emission



Sodium D₁/2 lines toward b Pictoris

Olofsson, Liseau & Brandeker 2001, ApJL 563

Dynamical studies...



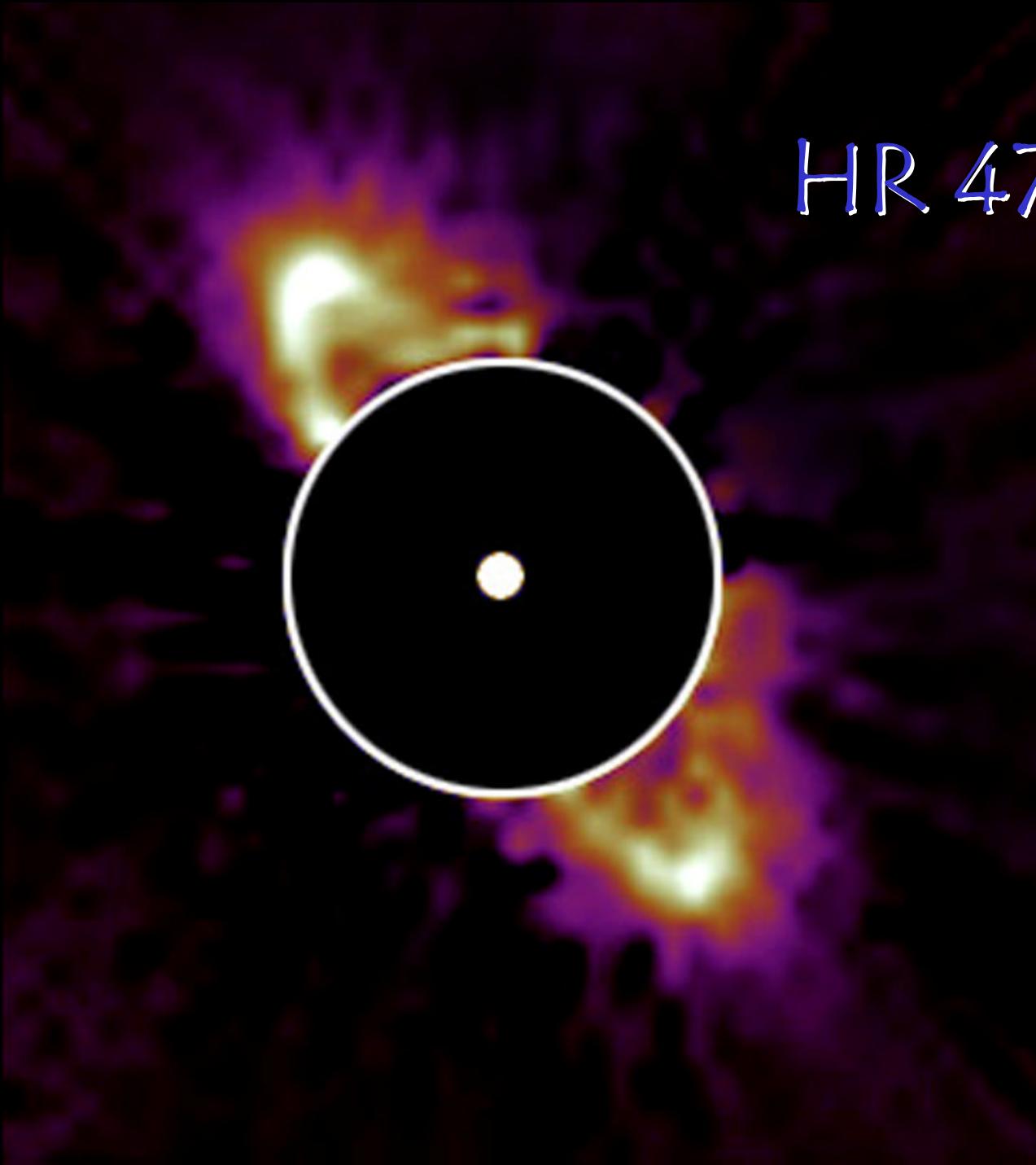
Stellar structure:

$$M_* = (1.75 \pm 0.1) M_\odot$$

Gas dynamics:

$$M_{\text{dyn}} = (1.40 \pm 0.05) M_\odot$$

Discrepancy probably due
to radiation pressure!

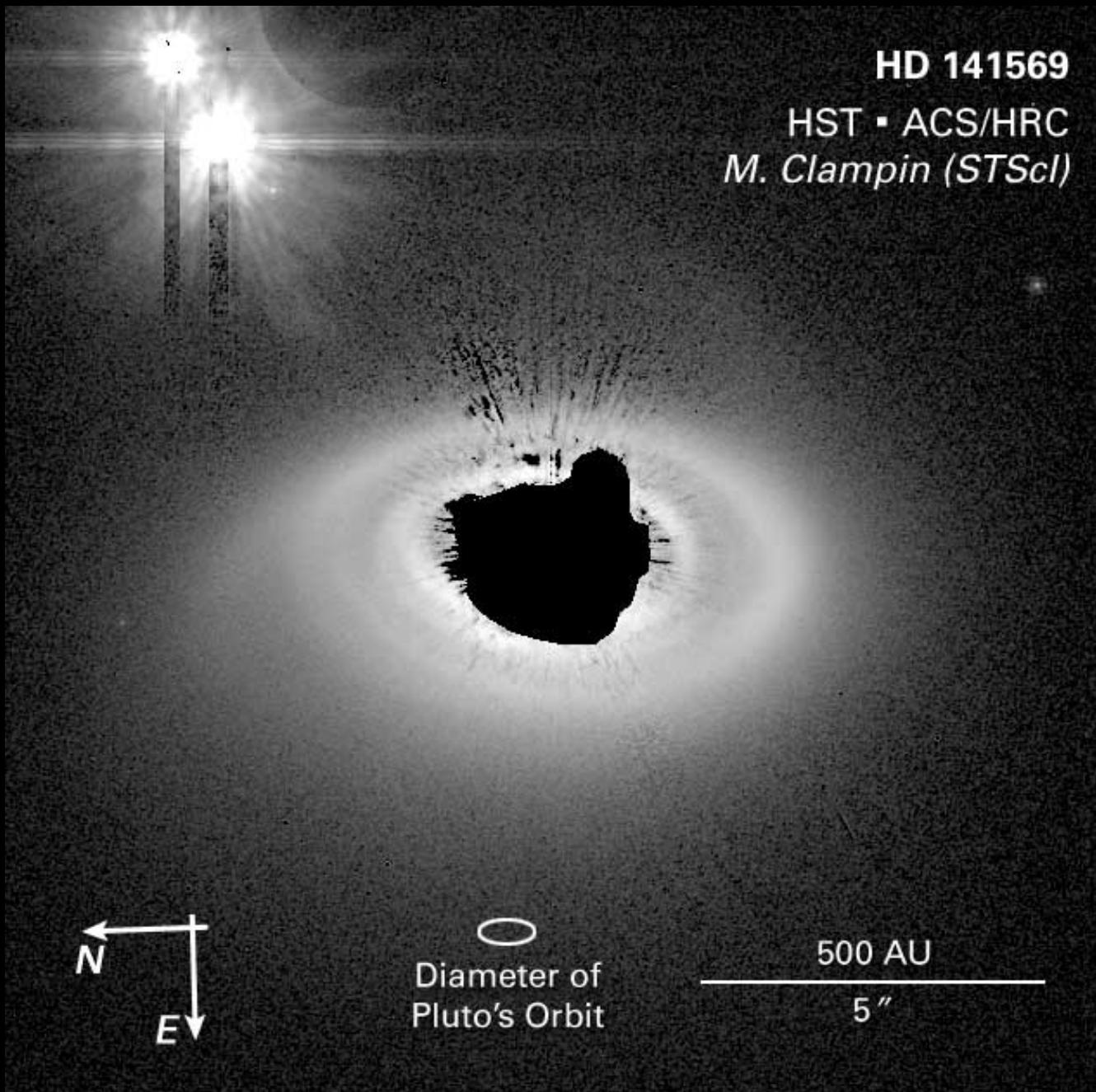


HR 4796A

HD 141569

HST • ACS/HRC

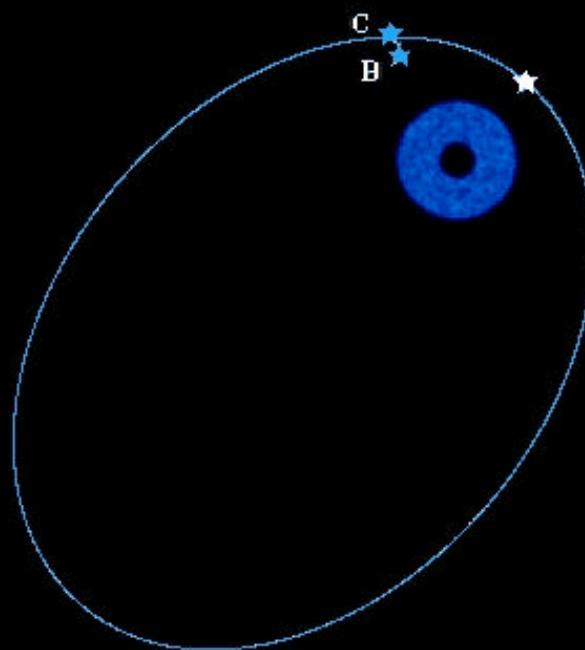
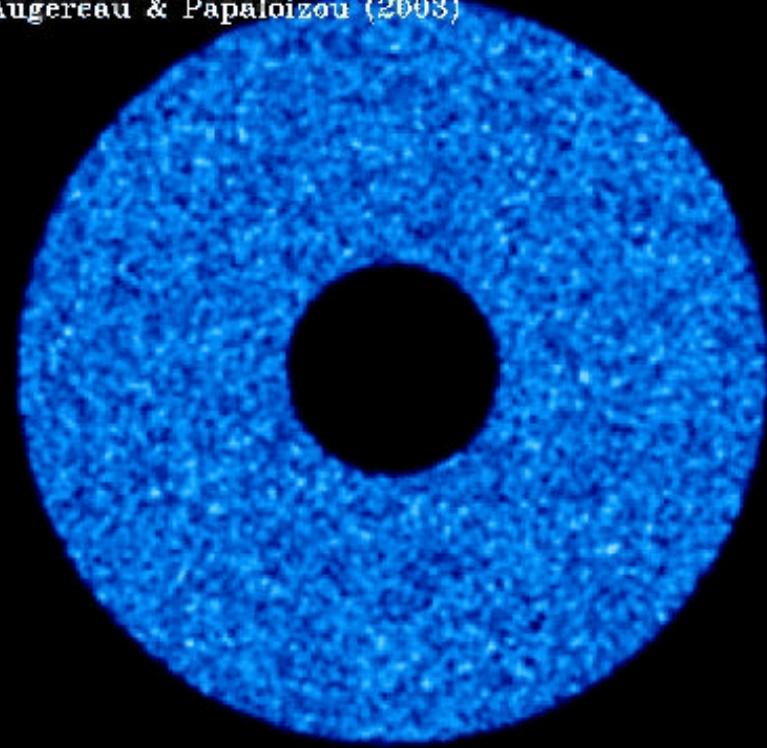
M. Clampin (STScI)



Diameter of
Pluto's Orbit

500 AU
5"

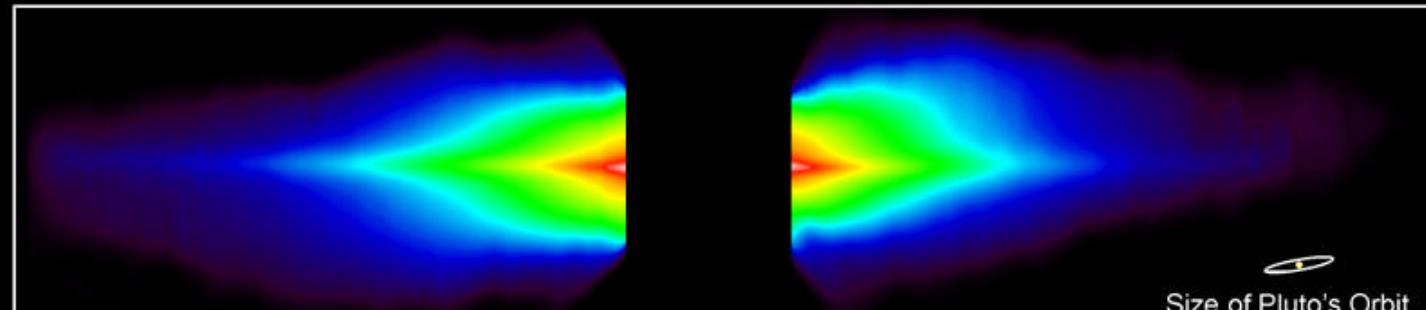
Augereau & Papaloizou (2003)



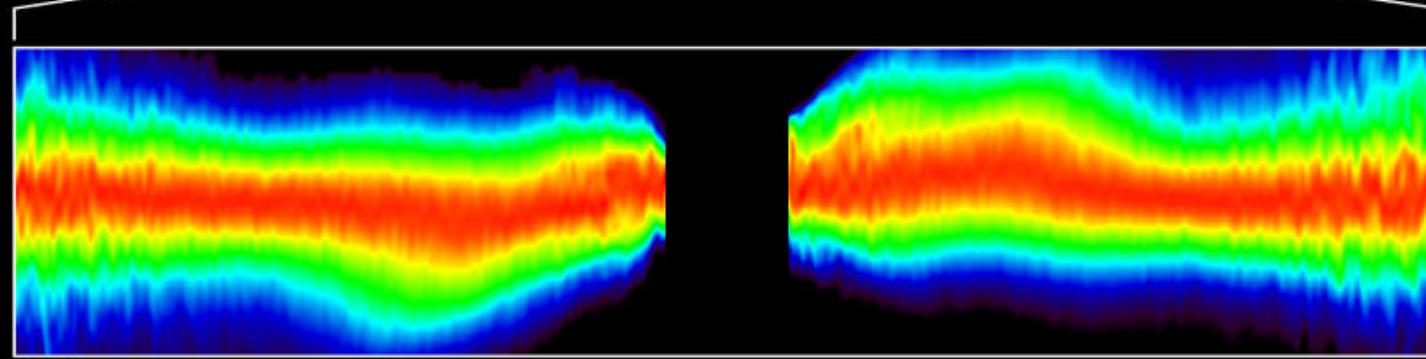
$t = 0.00000$

β Pictoris in scattered light

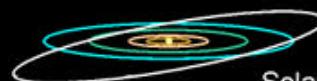
NE



SW



STIS



Solar System to Scale

Beta Pictoris

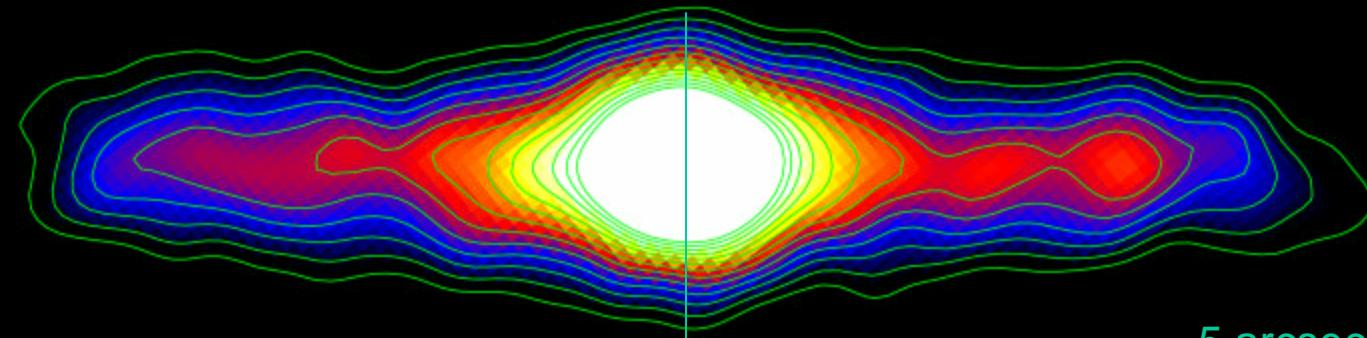
PRC98-03 • January 8, 1998 • ST Scl OPO

A. Schultz (Computer Sciences Corp.), S. Heap (NASA Goddard Space Flight Center) and NASA

HST • WFPC2 • STIS

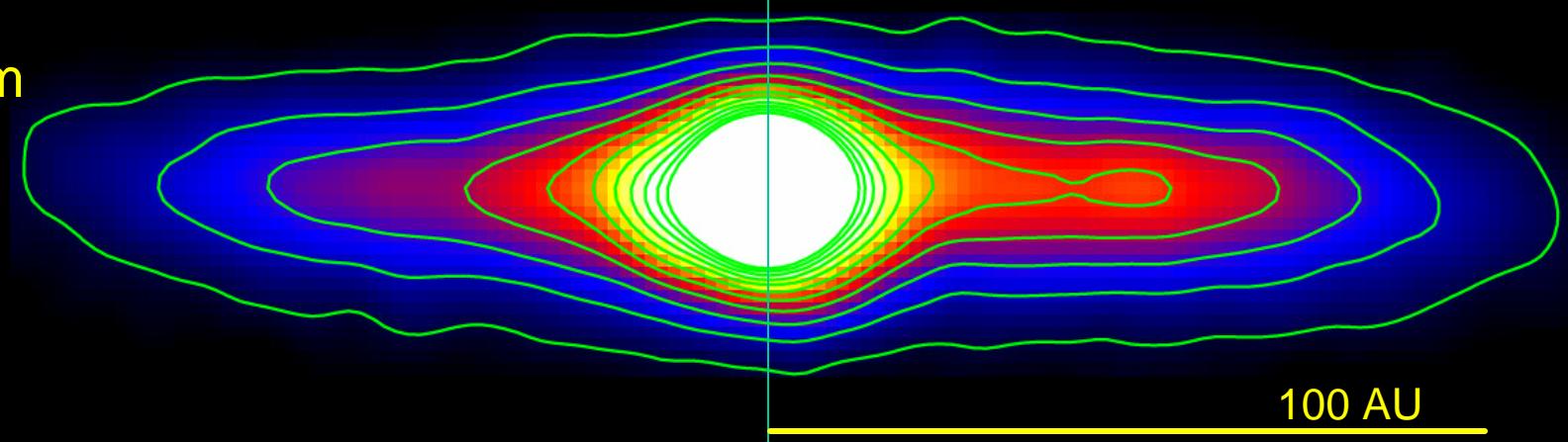
NE

24.3 μm

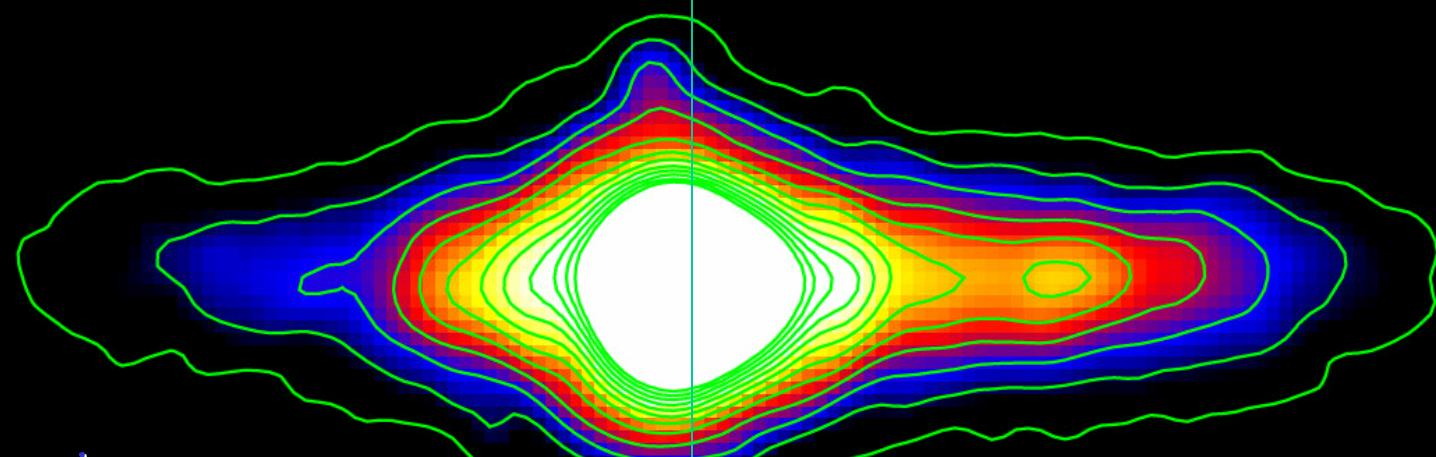


SW

18.3 μm



11.7 μm



Gemini south / T-ReCS

Telesco et al. 2005

pixel

60

80

100

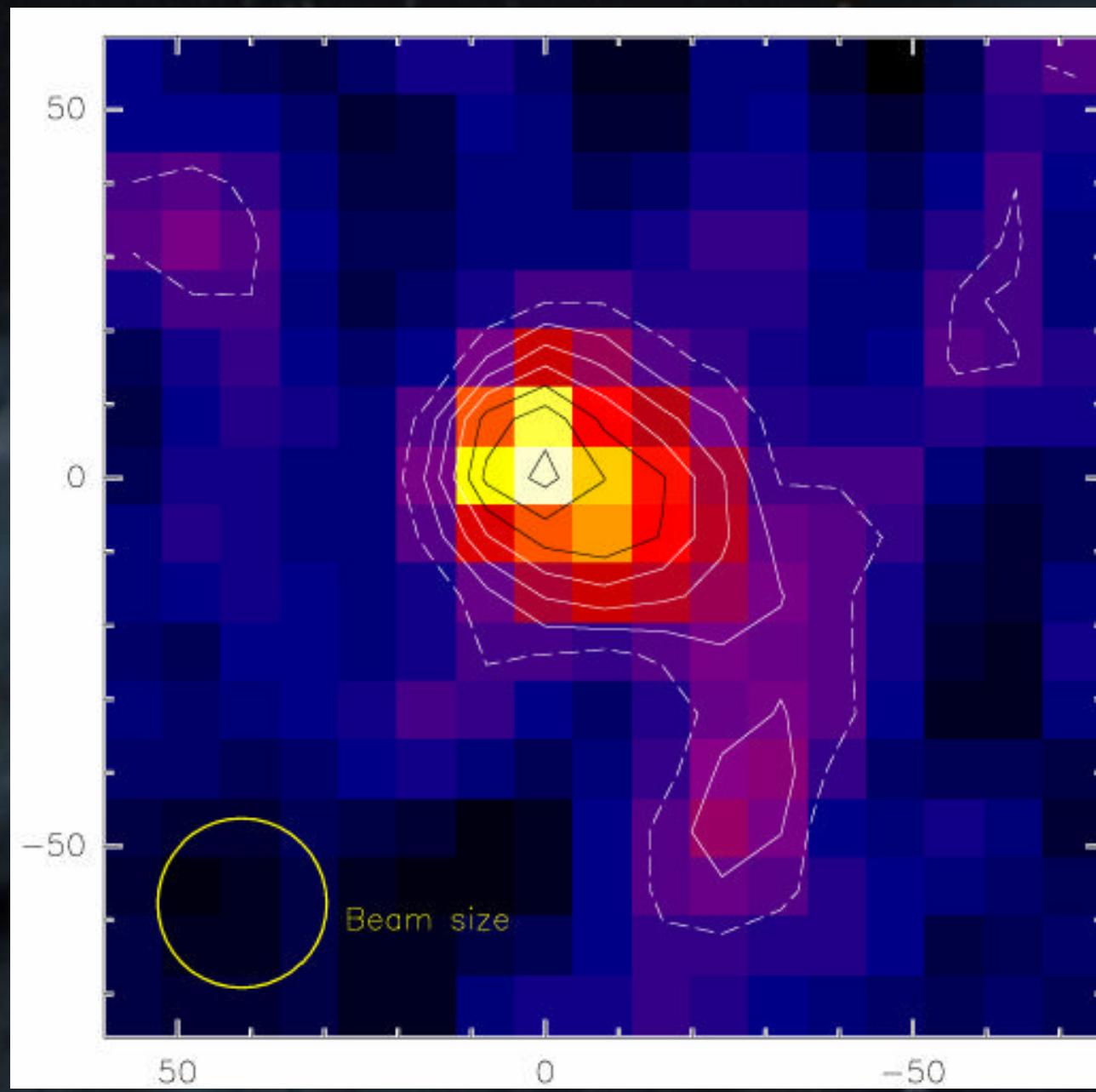
120

140

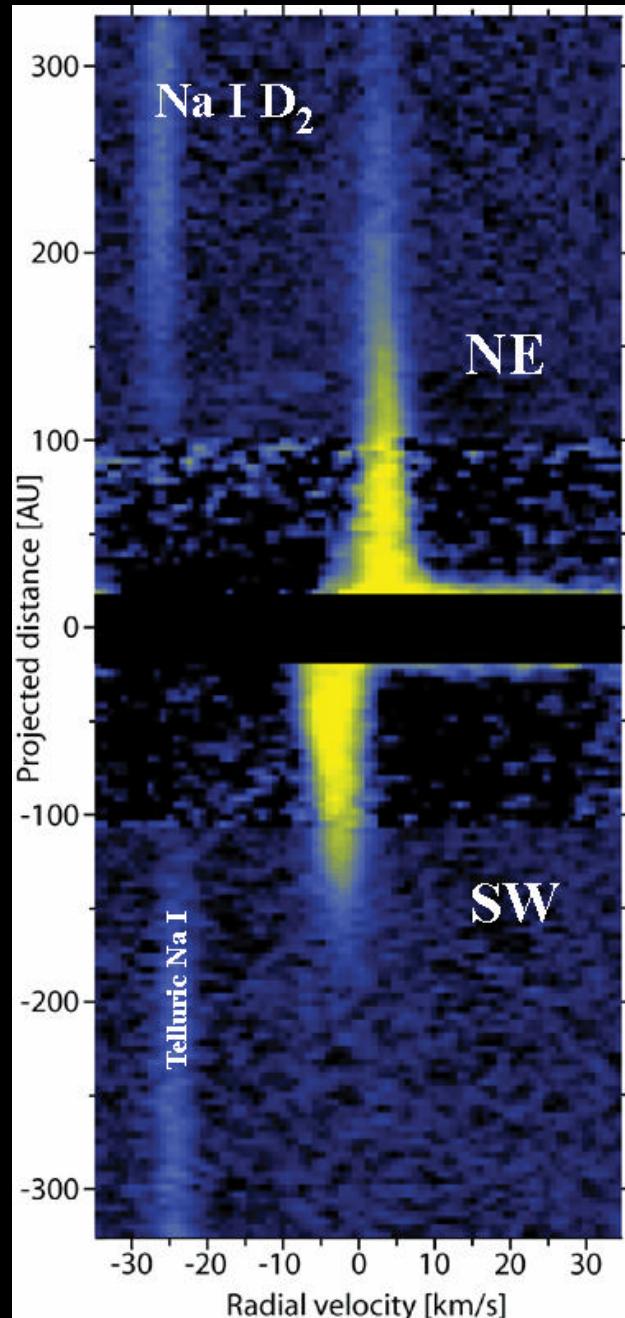
160

33

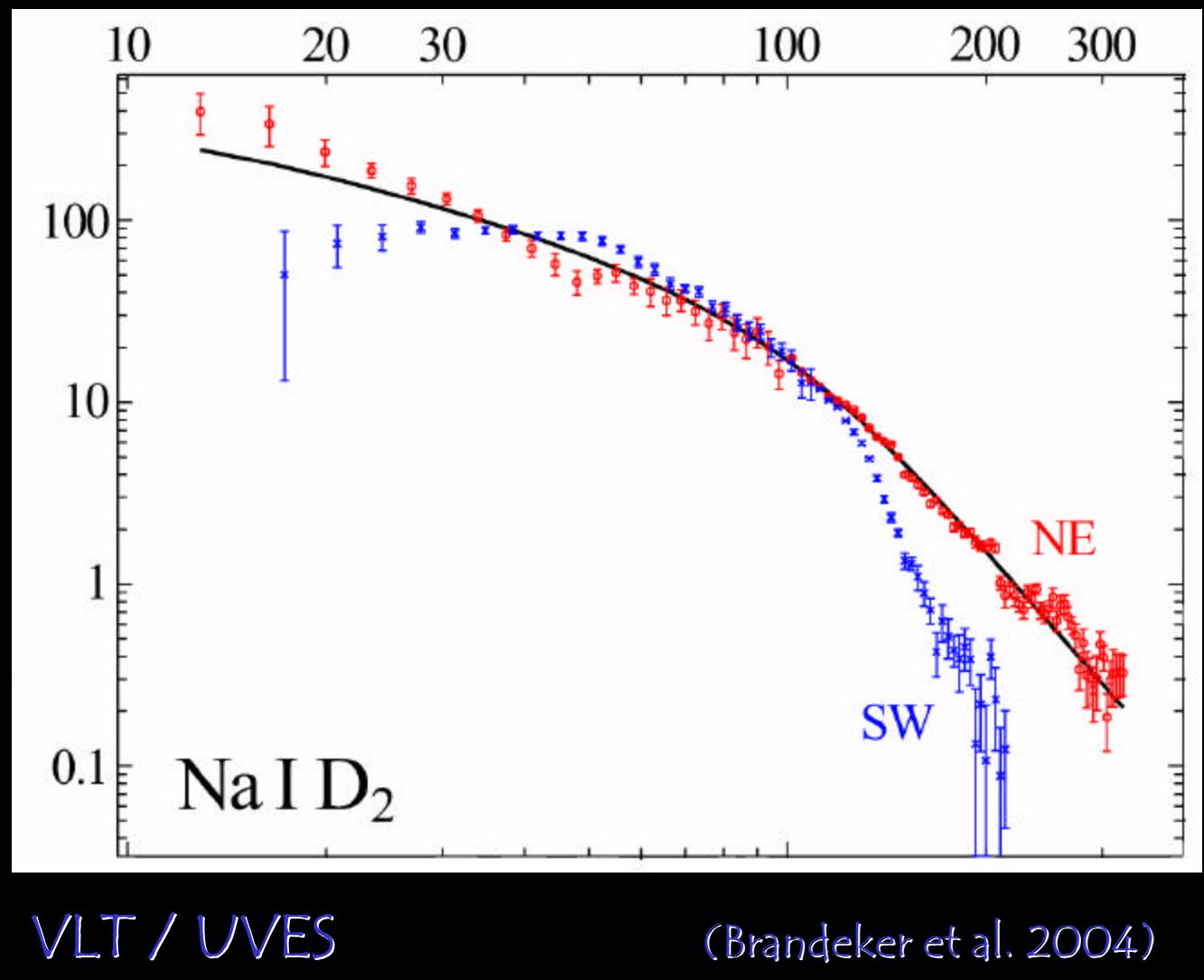
1.2 mm (SEST/SIMBA)



Liseau et al. 2003



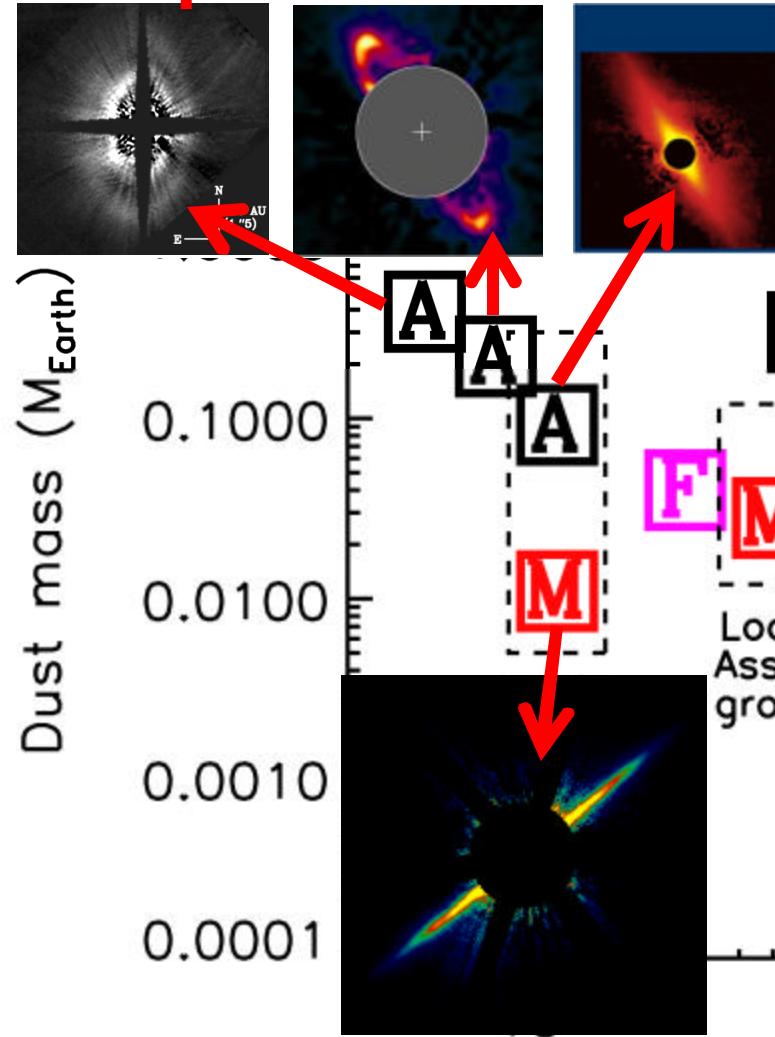
Asymmetries in gas disk



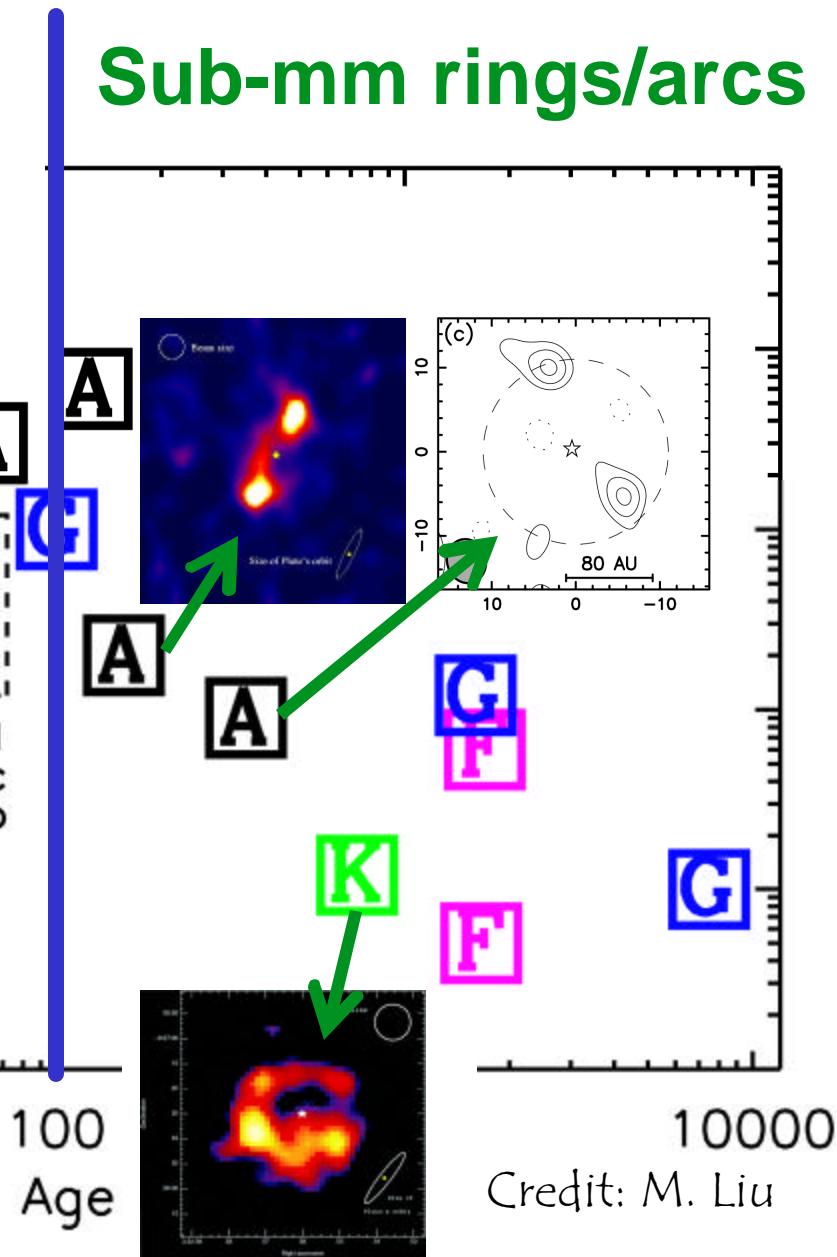
(Brandeker et al. 2004)

Evolution: young vs. old debris

Optical/IR disks



Sub-mm rings/arcs



StarStrider

