

Close binaries in young systems

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How sensitive is star formation to environment?

Diagnostics:

- Mass function (stars/BDs?)
- Multiplicity (fraction of multiples?
Configurations?)

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Determining multiplicity

- Adaptive optics surveys down to ~ 60 mas
- Optical interferometry: ~ 1 mas
- Radial velocity surveys: period $\sim$ years (depending on patience)
- Completeness? @ 150 pc, 60 mas ~ 27 yr
@ 50 pc, 60 mas ~ 5 yr!

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Close companion programme

Radial velocity survey of 141 stars in young (< 100 Myr), nearby (< 100 pc) groups:

η Cha	~ 8 Myr	97 pc
TWA	~ 8 Myr	20-100 pc
BPMG	10-20 Myr	10-40 pc
TucHor	20-30 Myr	40-60 pc
AB Dor	50-100 Myr	20-50 pc

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Close companion programme

Target selection:

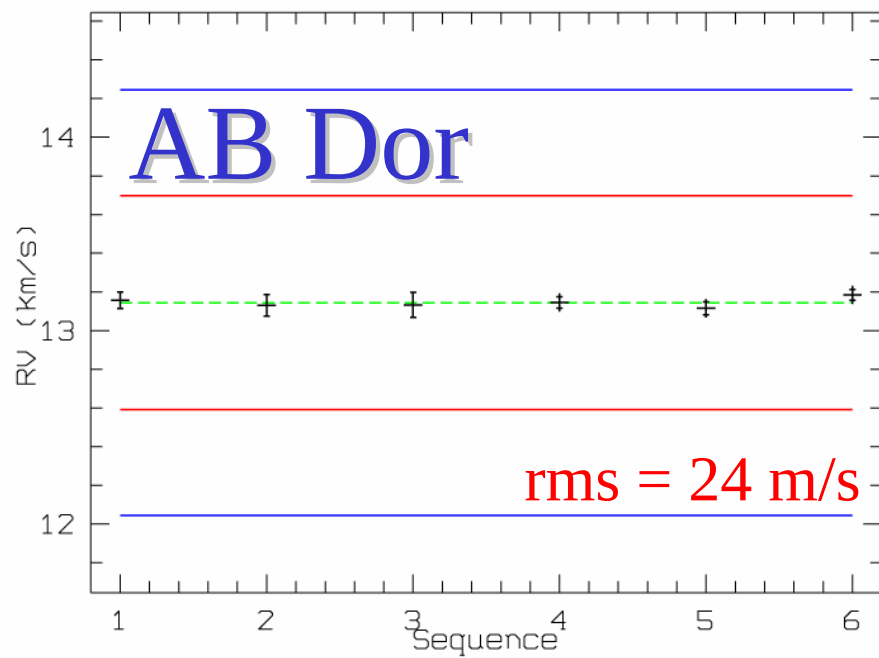
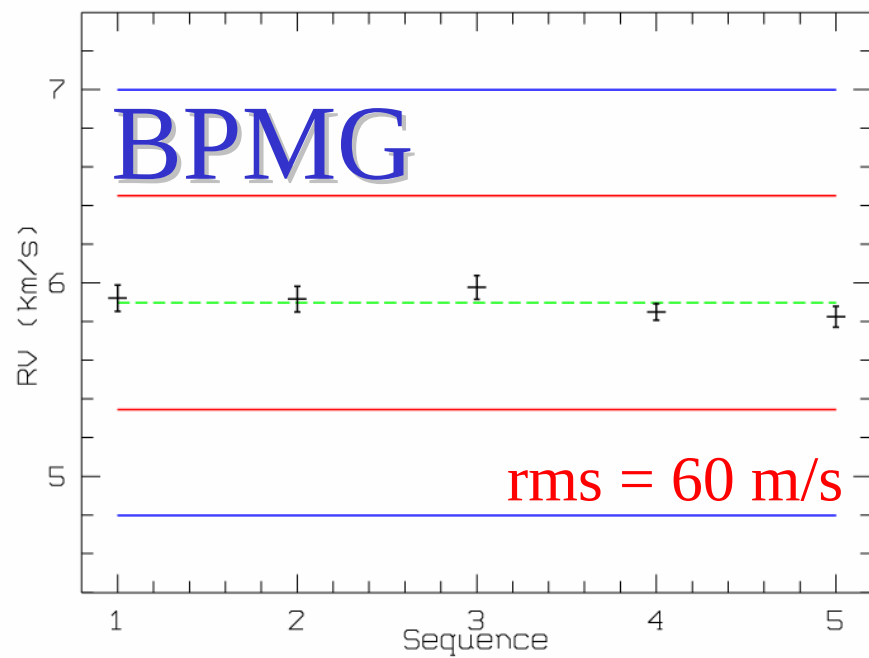
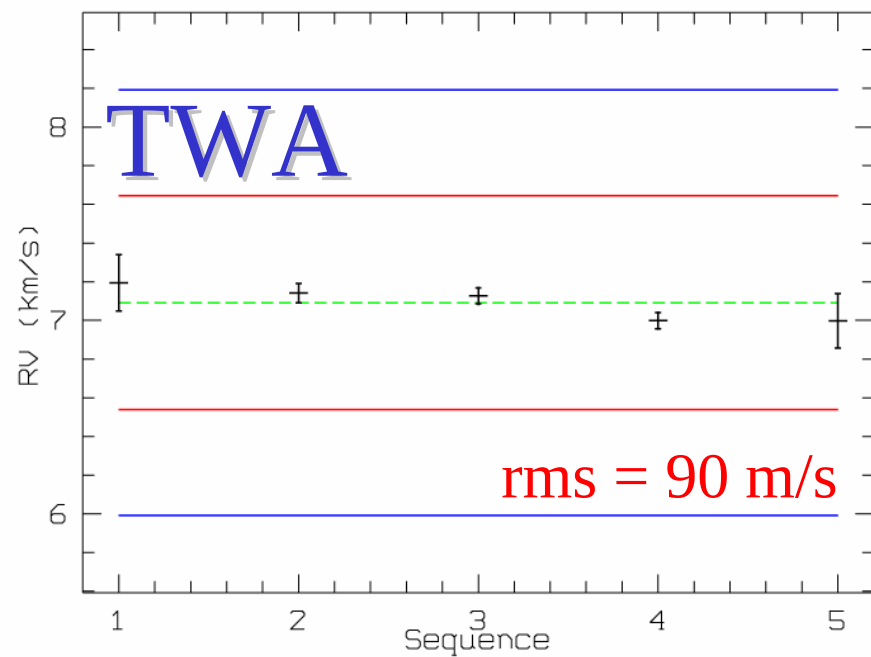
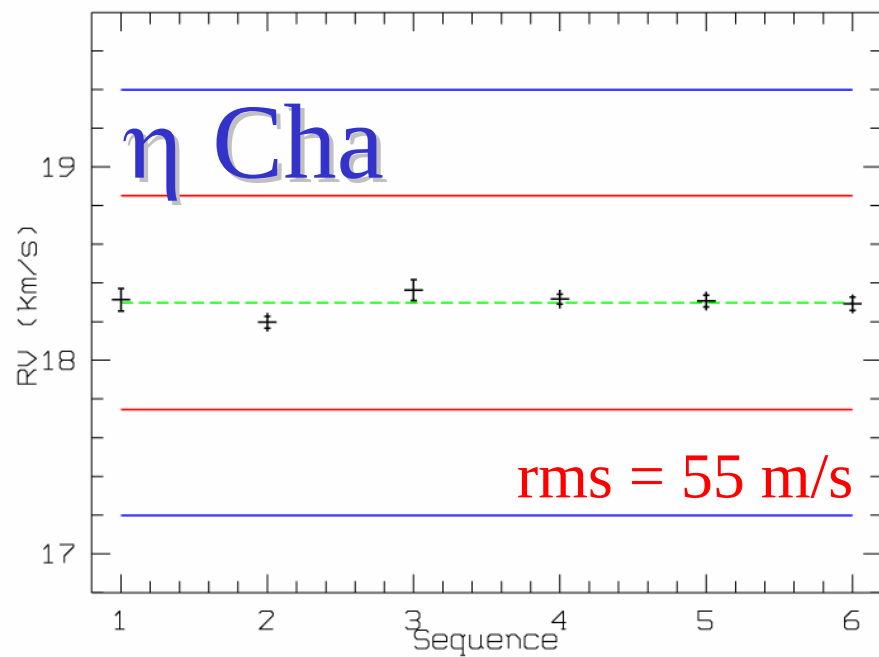
- All known members of SpT later than F5 ($< 2 M_{\text{Sun}}$)
- Visible from Las Campanas ($\delta < +30^\circ$).
- Brighter than R=14 (earlier than M5)

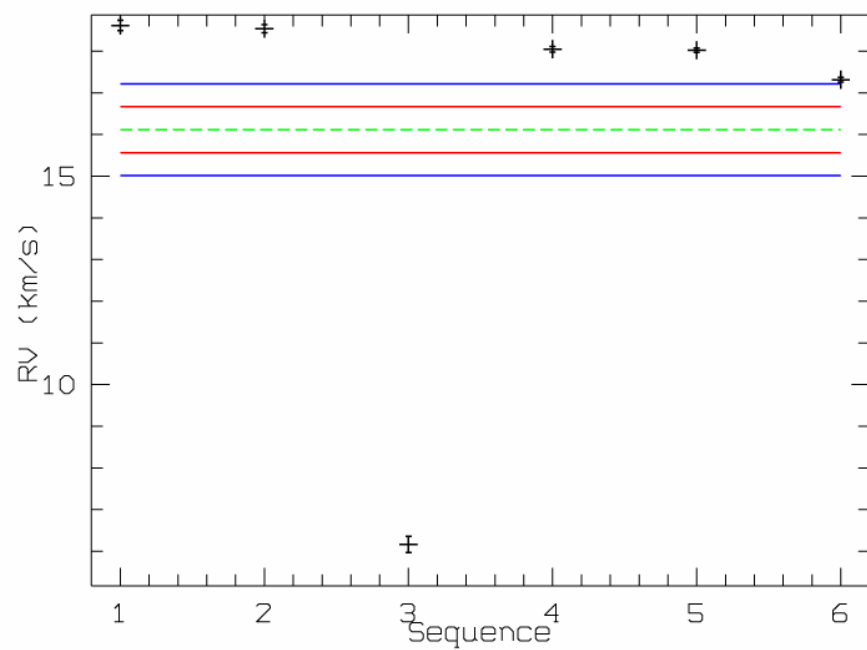
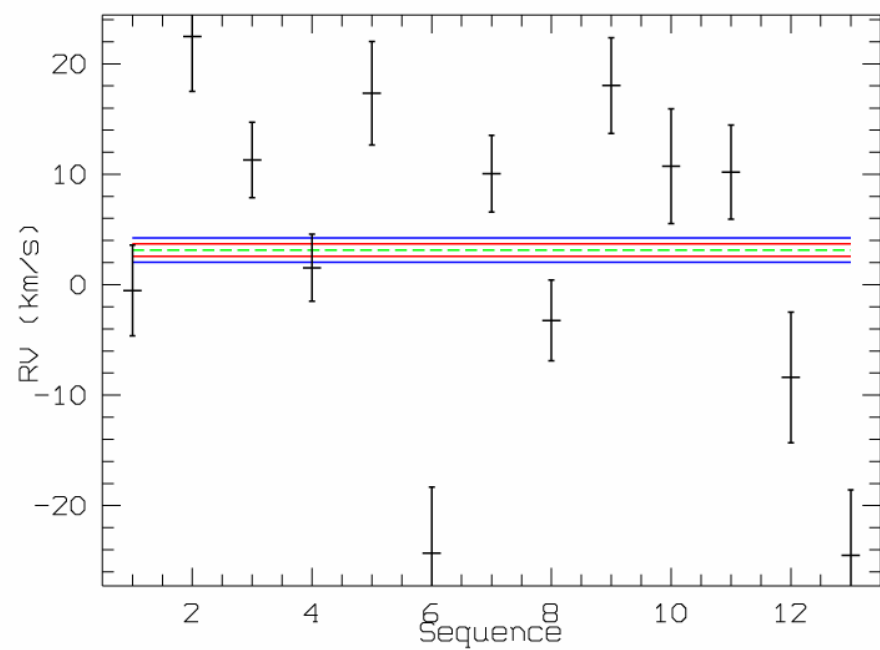
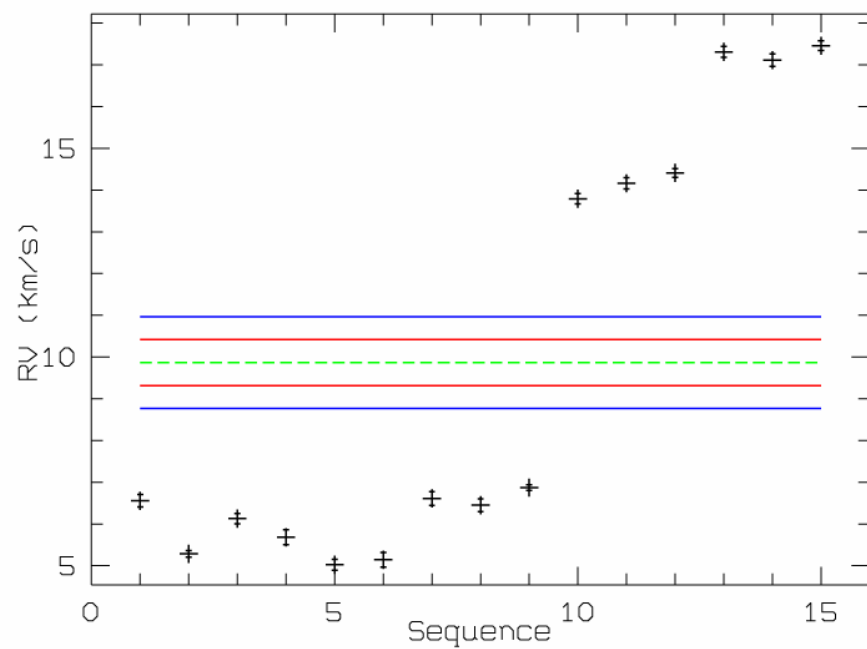
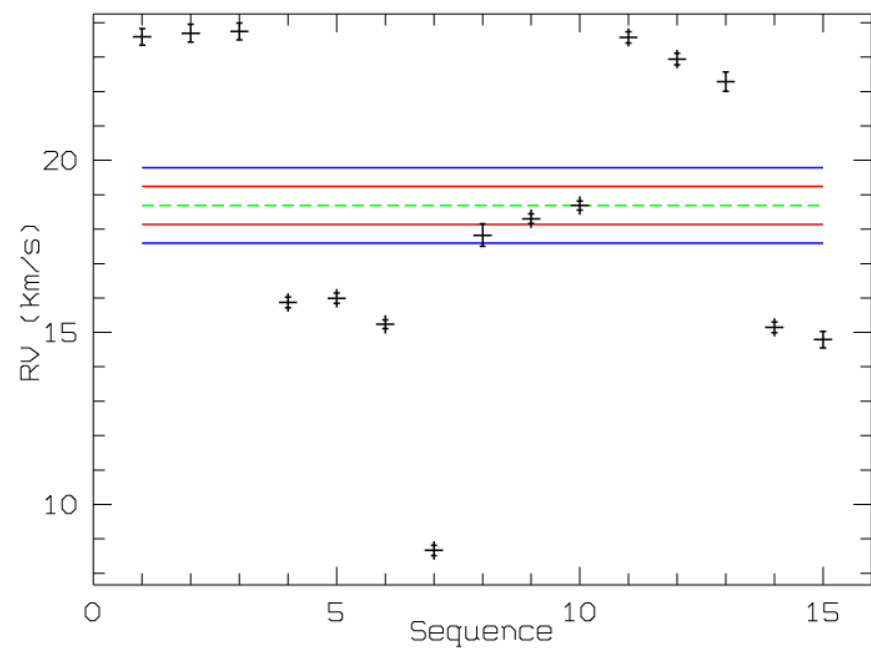
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Close companion programme

- Telescope:
Clay/Magellan 6.5m at
Las Campanas, Chile
- Instrument:
MIKE, High resolution
echelle spectrograph
($R \sim 55000$)
- 30 bright nights
distributed over 2 years
- ~ 2200 spectra obtained

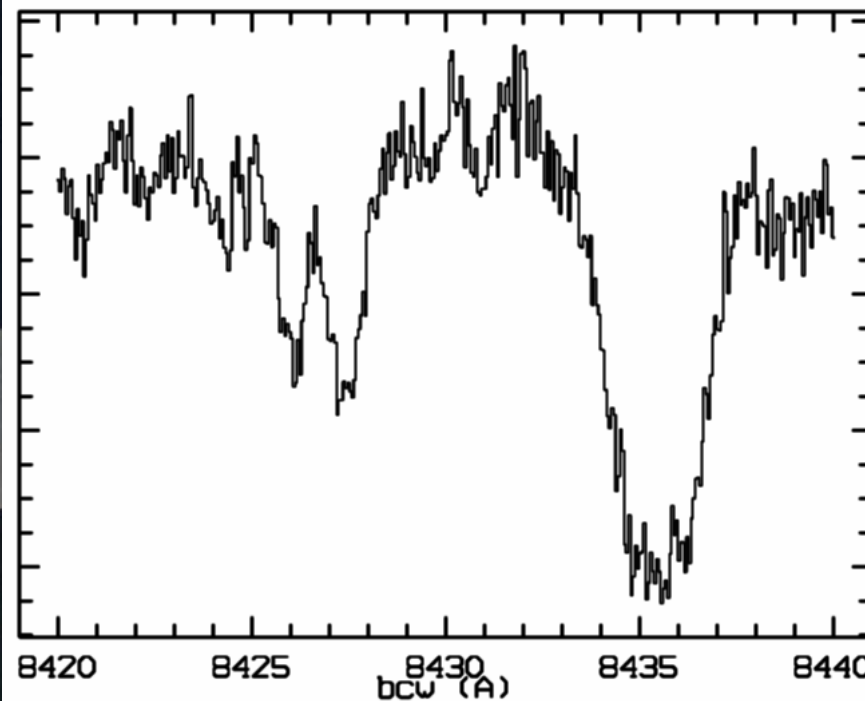
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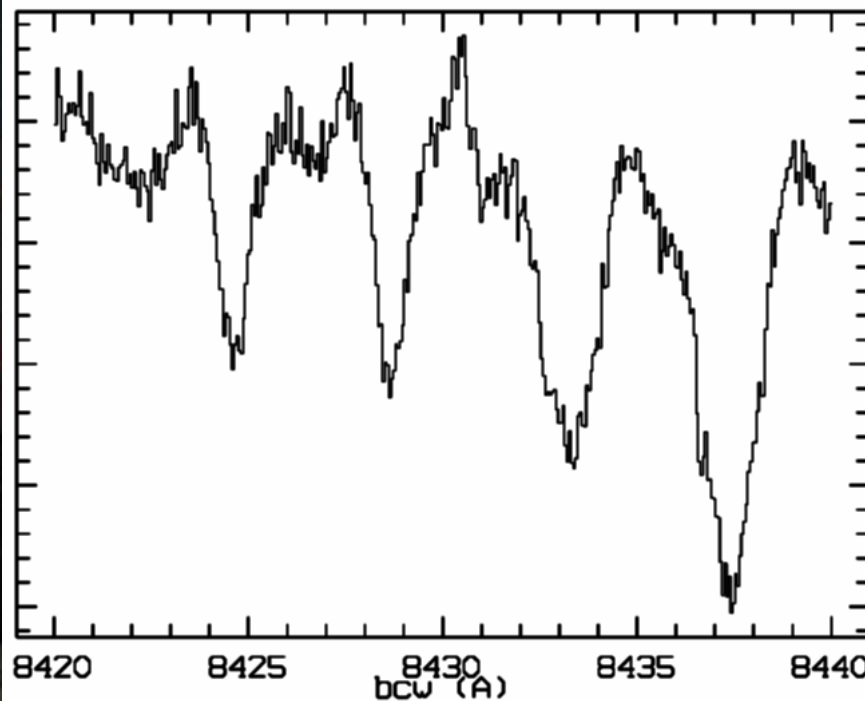


Example: TWA 22

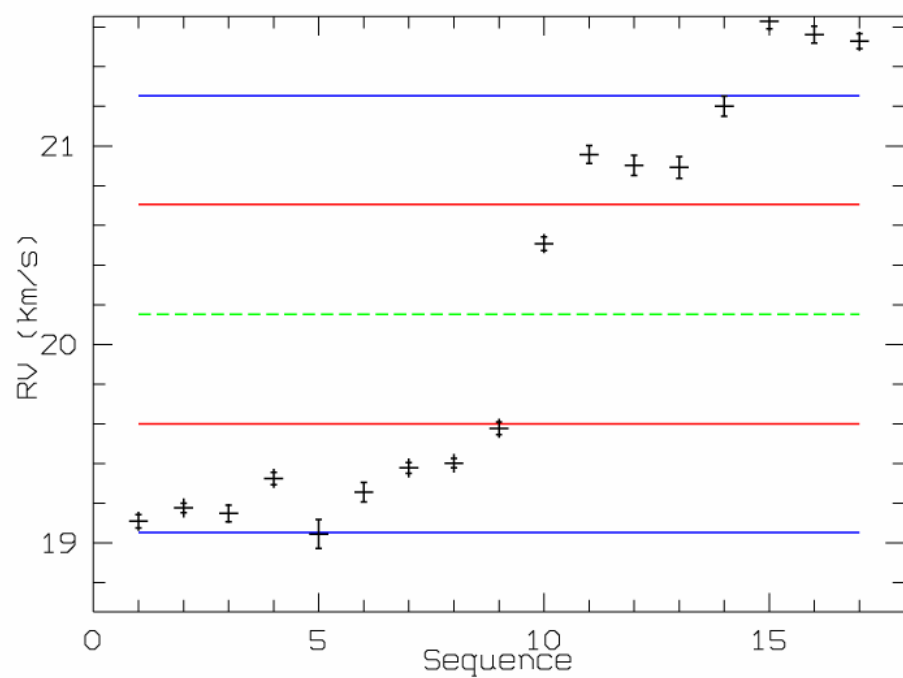
Epoch 1



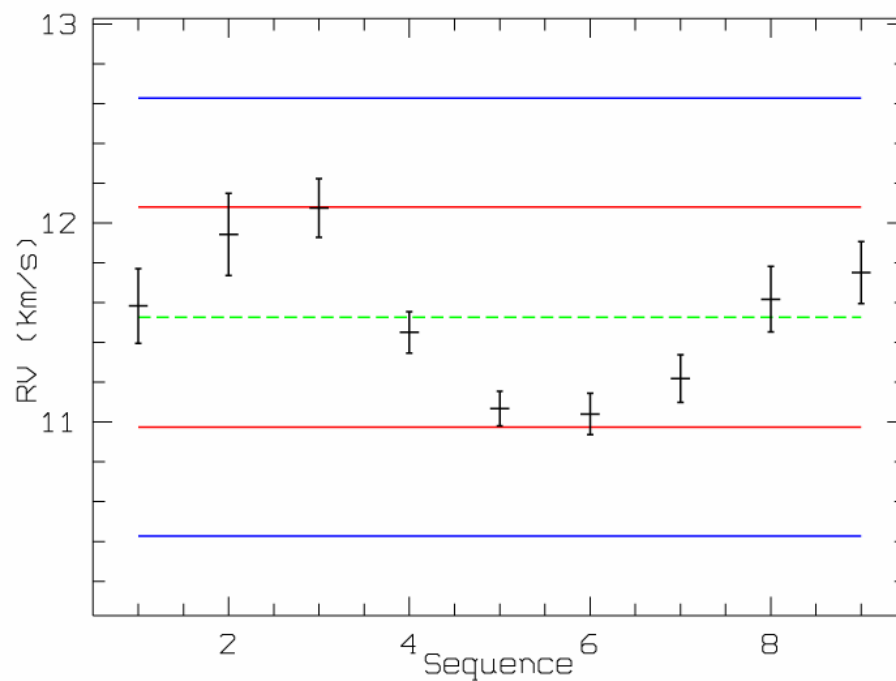
Epoch 2



BD candidate (P~4 yr)



Planet candidate (5 M_J)



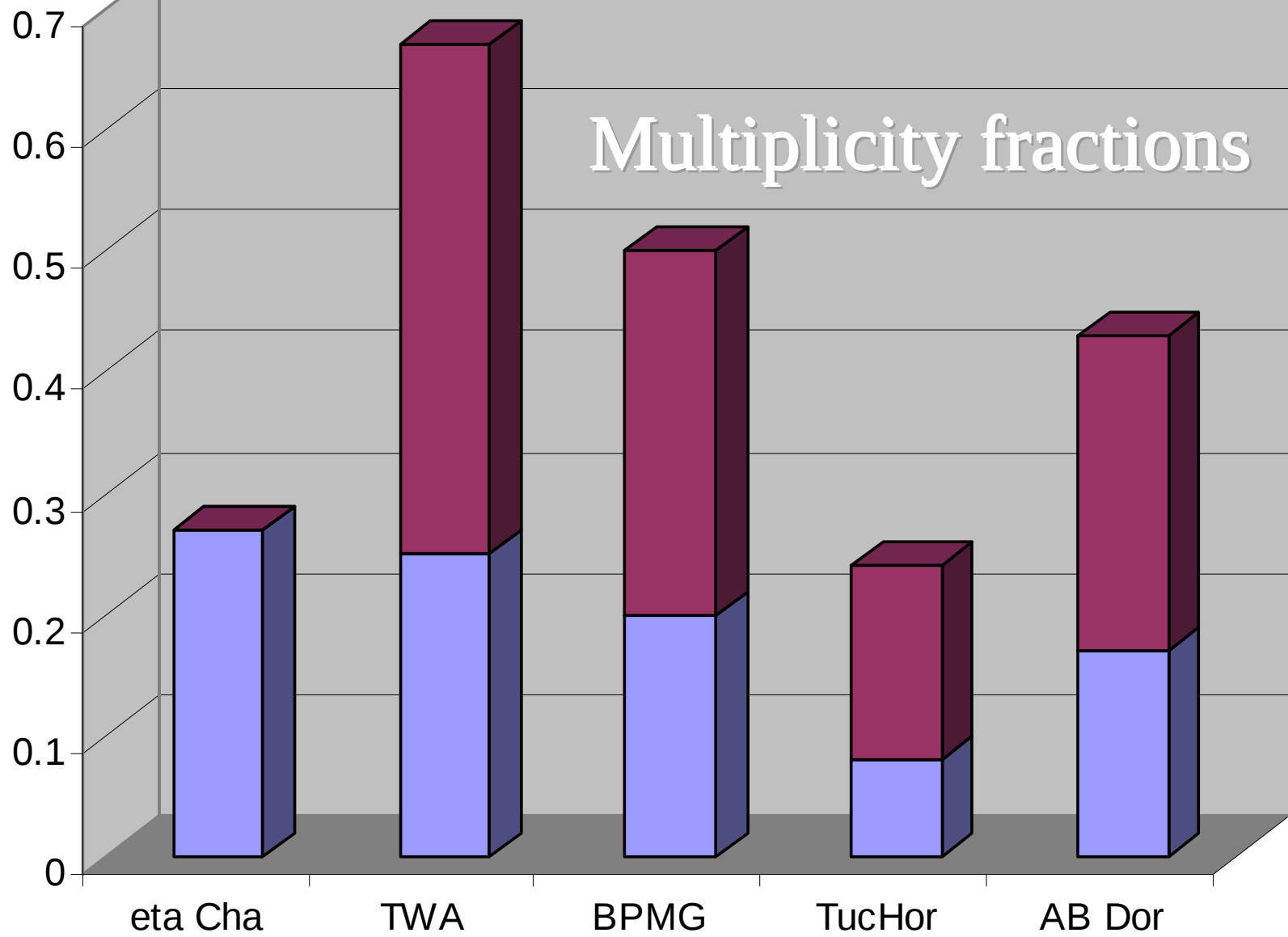
Combined result

$$\text{MSF} = (42 \pm 5) \%$$

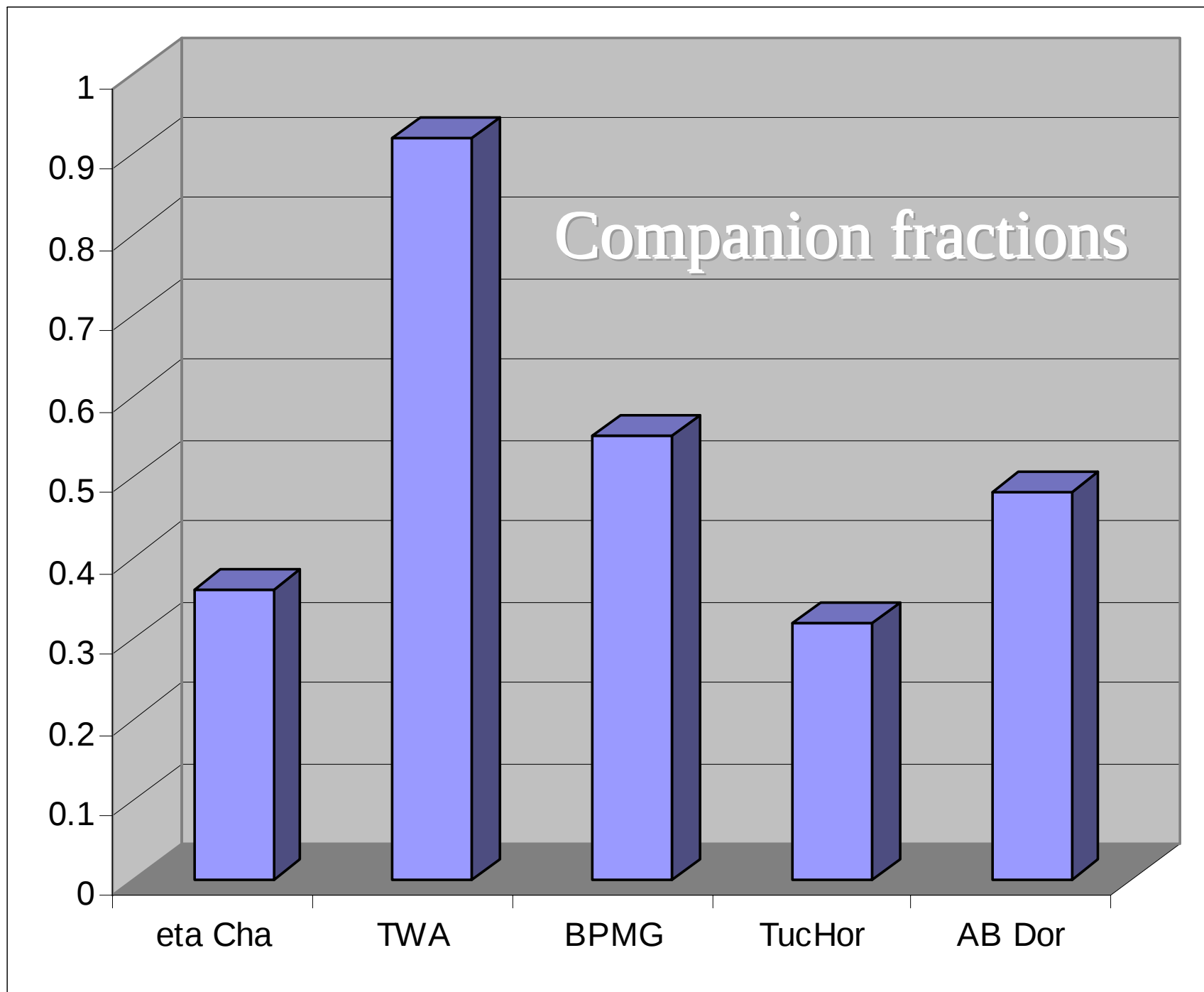
$$\text{CSF} = (52 \pm 4) \%$$

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Multiplicity fractions



Close Wide



Other results

- Lithium dating (Mentuch et al., in prep)
- Rotation-accretion connection (Jayawardhana et al. 2006)
- Rotation-activity connection (Scholz et al. 2007)
- Dynamics (N/A)

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Conclusion

- Multiplicity is environment dependent.

Immediate future

- Analysis in progress; more juicy statistics to come.

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