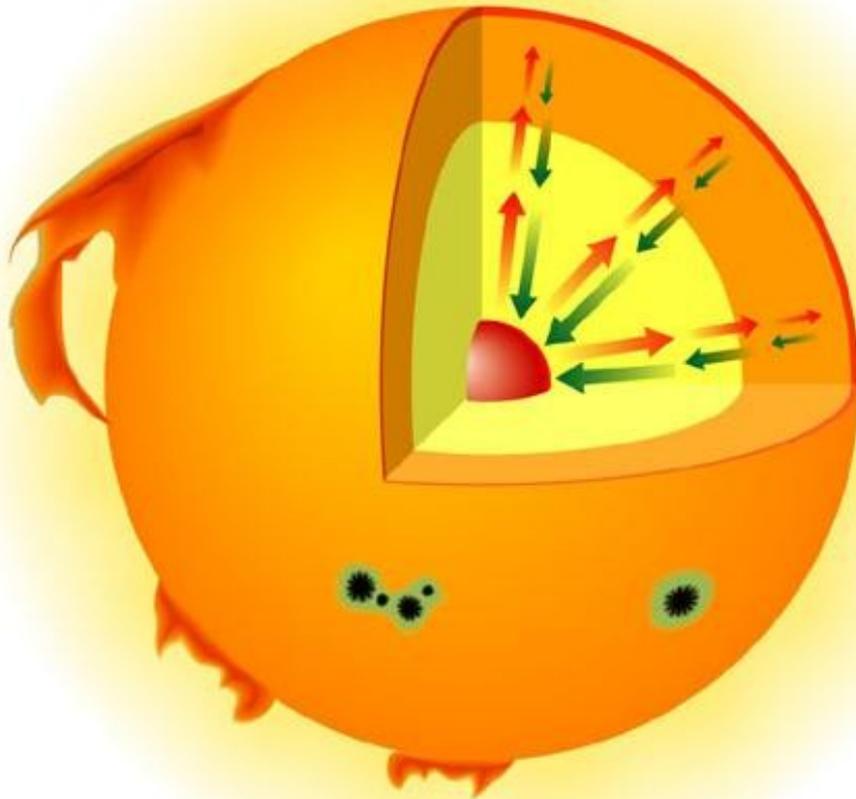


# Stellar Graveyards



# Protracted battle with gravity

pressure →  
gravity ←



To support weight:

⇒ need high pressure

⇒ need high temperature

⇒ will lose energy

⇒ need energy source:  
- Gravitational contraction  
- Nuclear fusion

Ultimately,

***fuel exhaustion***

# Mira ('the wonderful') -- the ceremonial dance for death

An old star (asymptotic giant), ~ 10 Billion years

change brightness drastically over ~ year, pulsing  
*(Mira: discovered 1596, visible at bright phase)*

losing ~ Earth mass every pulsing

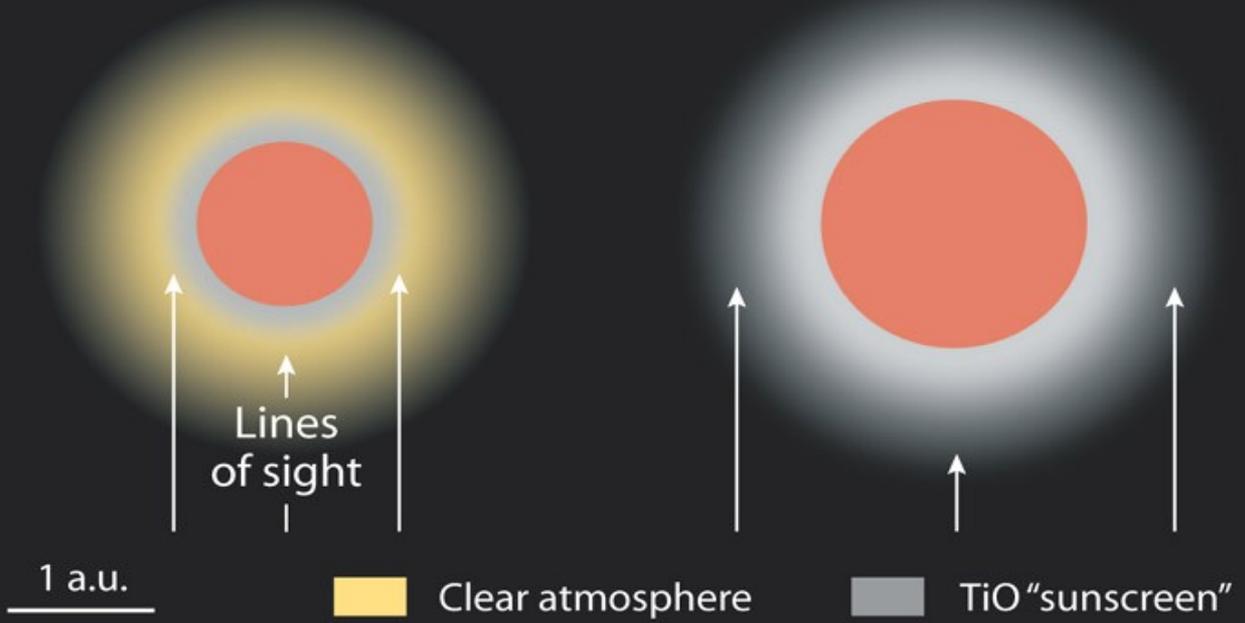
live for ~ few Million years



## Pulsing Mira-Type Star

Maximum light

Minimum light

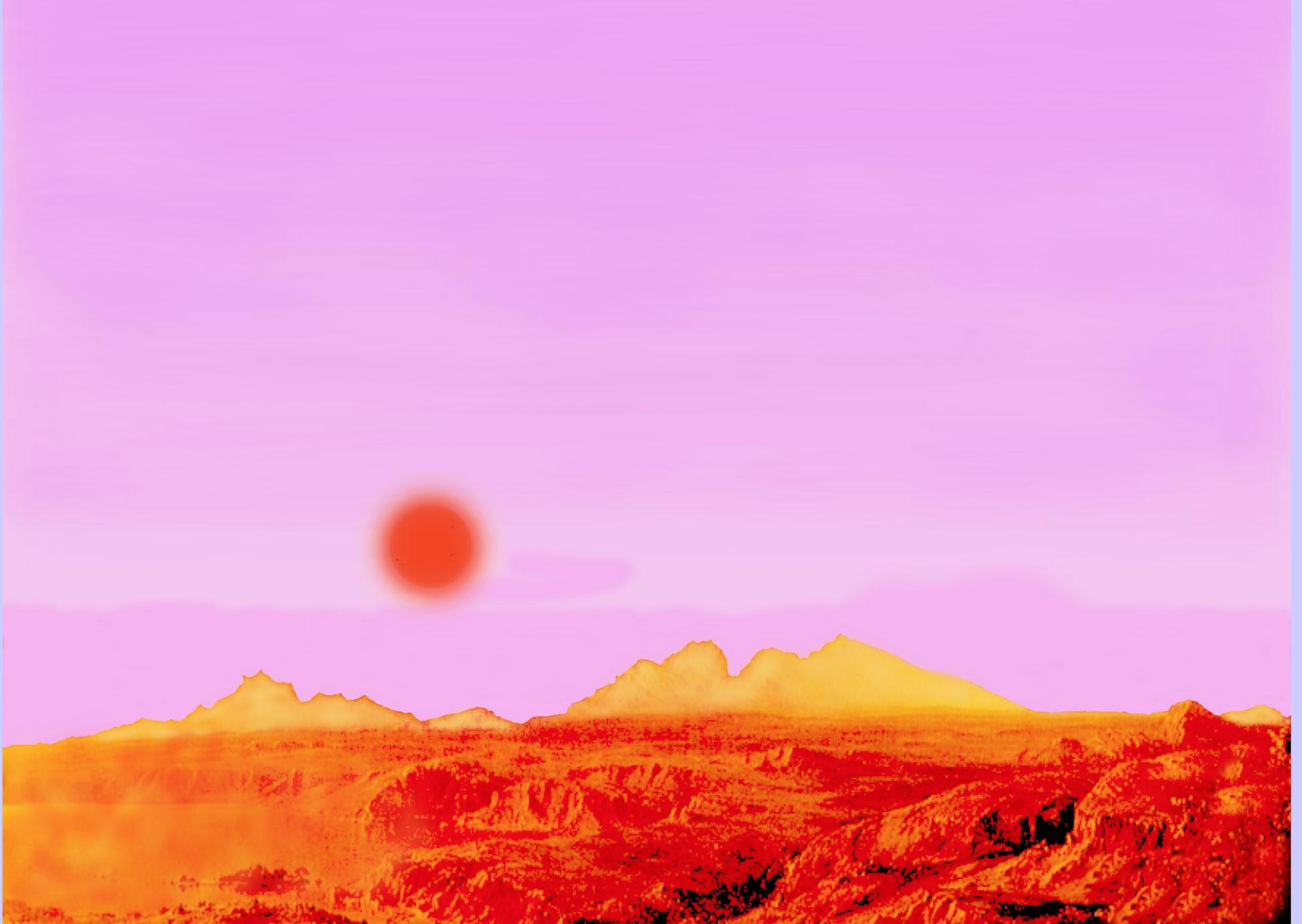


*Dimmest when biggest*

Mira at minimum (viewed from a resident planet)



Mira at maximum (viewed from a resident planet)

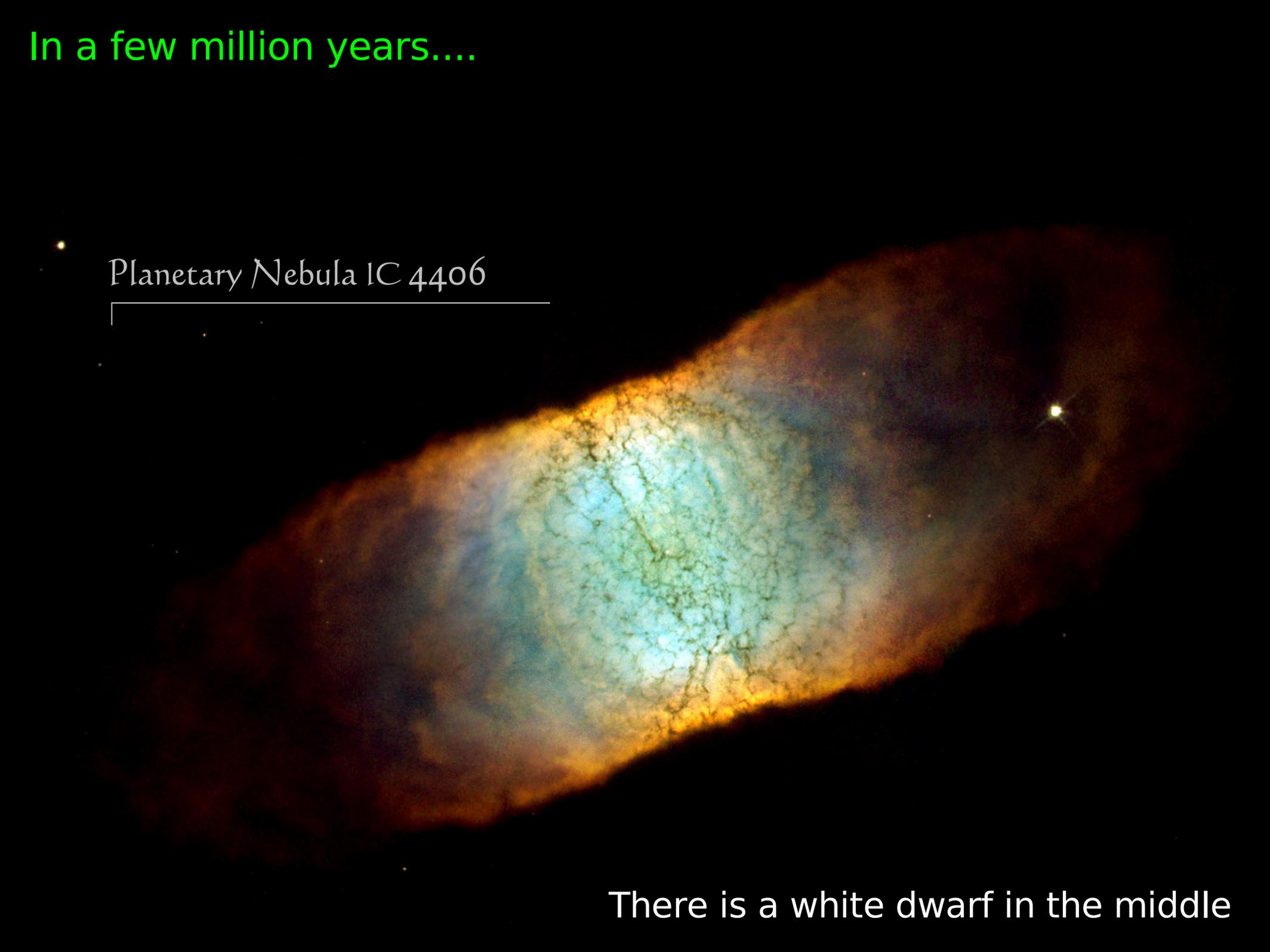


In a few million years....

Planetary Nebula IC 4406

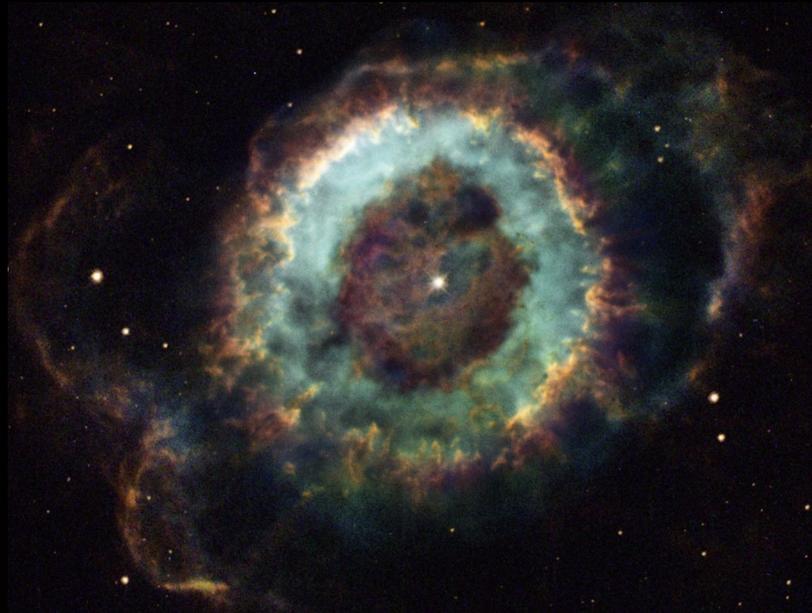
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There is a white dwarf in the middle

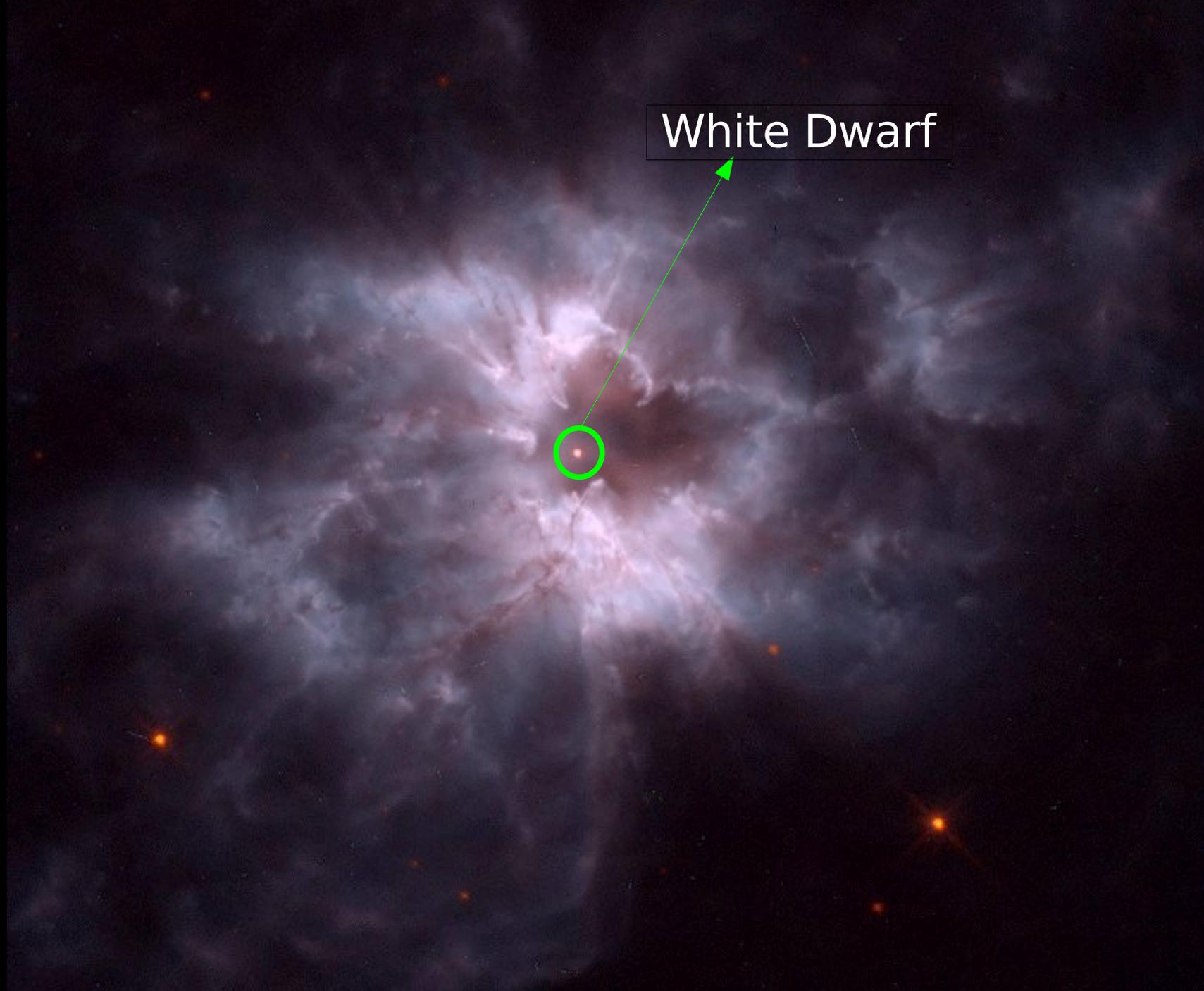


So will the Sun (in another 5 billion years...)

And how will we look like in death?

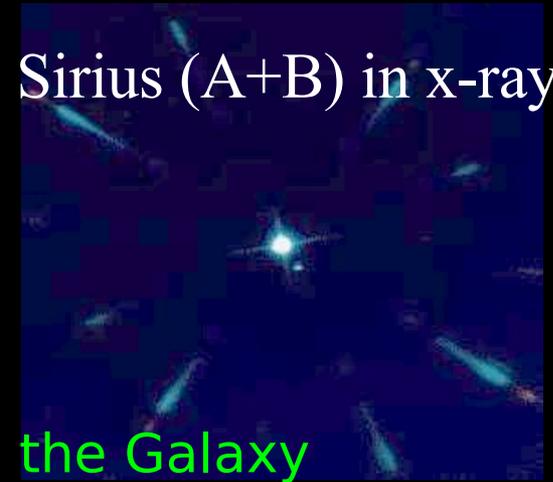


White Dwarf: remnant of a low mass ( $< 8 M_{\text{sun}}$ ) star

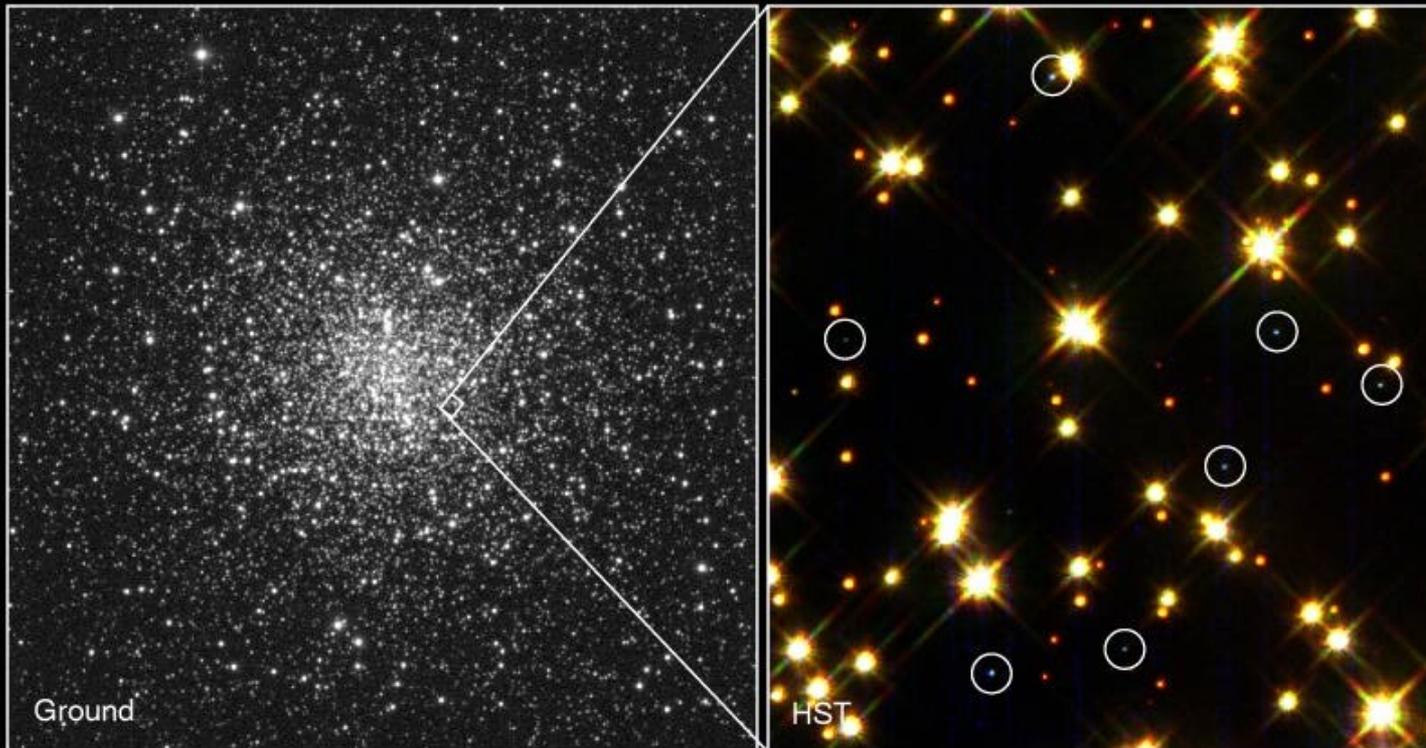


# White dwarfs

Sirius (A+B) in x-ray



There are as many white dwarfs as luminous stars in the Galaxy  
they are dead objects simply fading with time



**White Dwarf Stars in M4**

**HST · WFPC2**

Sirius: the brightest star on the sky  
also has a white dwarf companion →

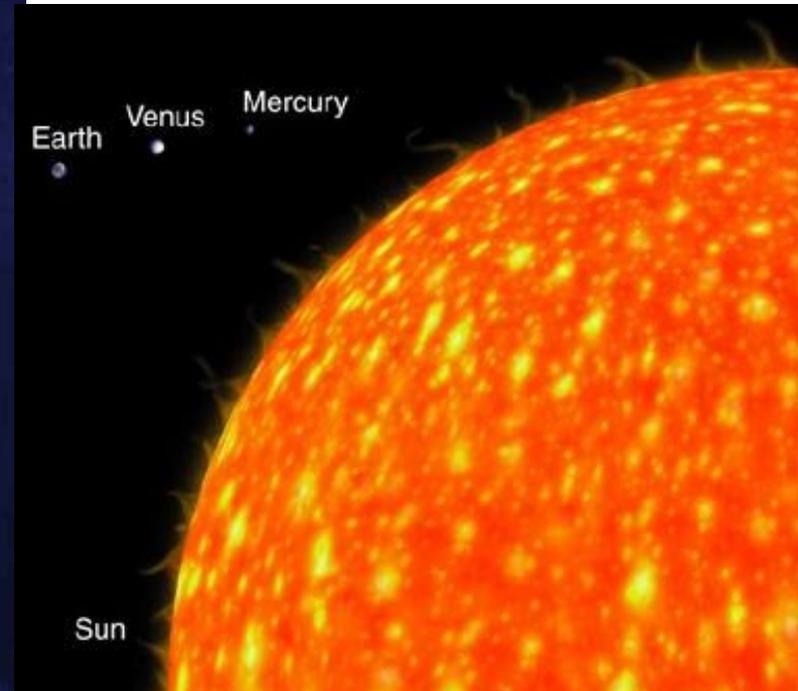
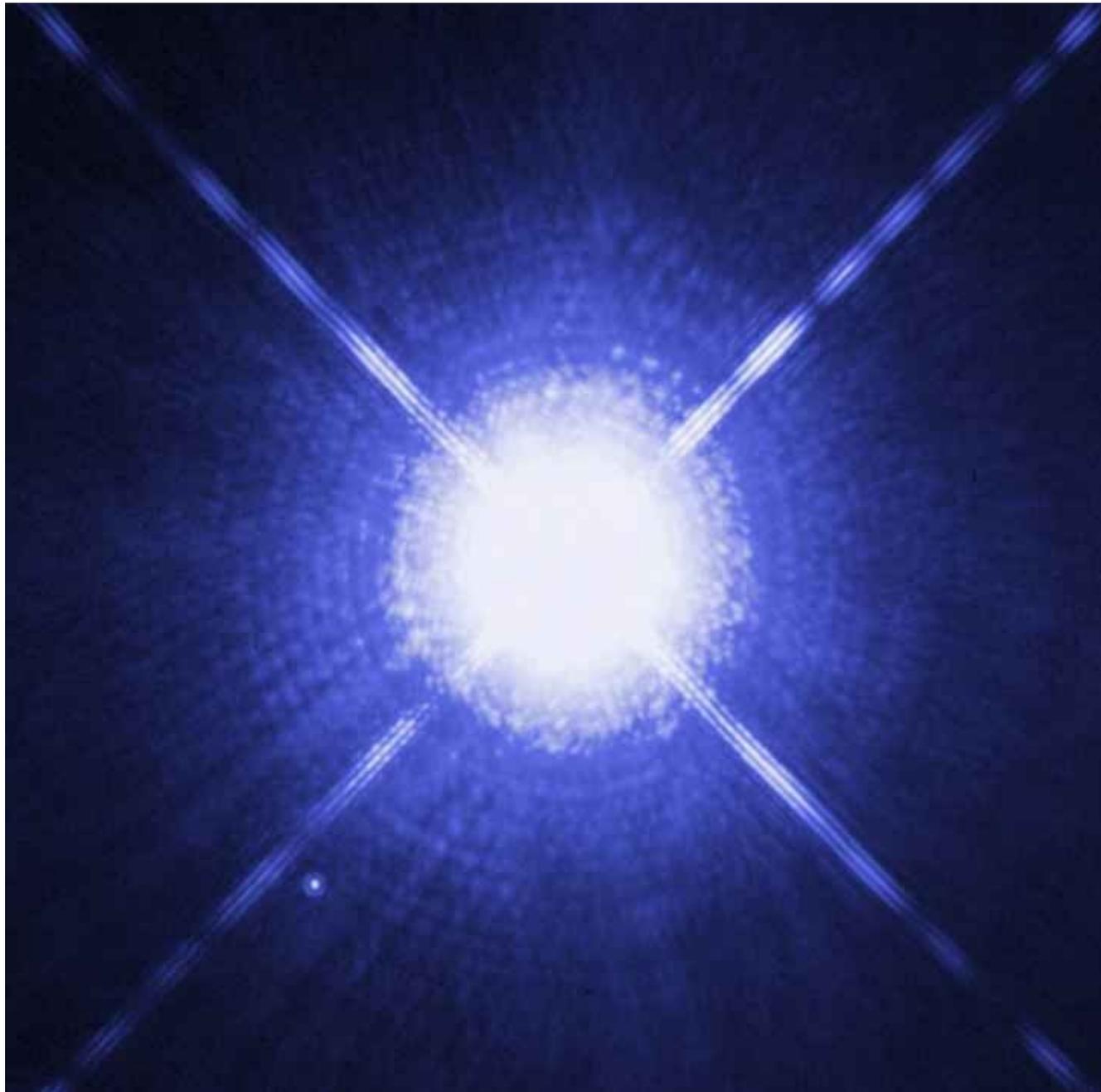
Mass ~ Sun,  
Radius ~ Earth

⇒ very dense

(degenerate matter)

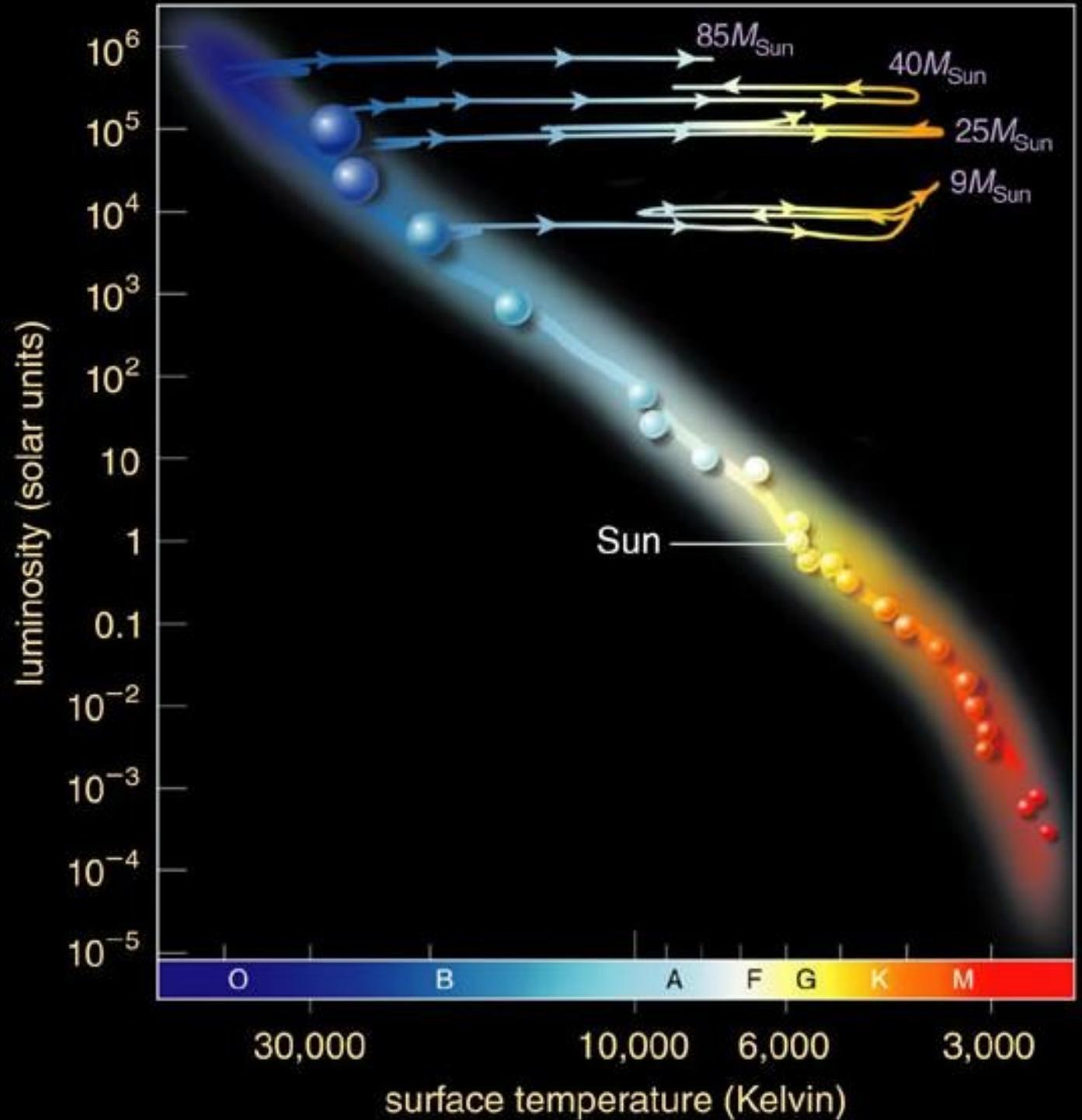
density ~ 1 ton/cm<sup>3</sup>

*(a million times denser than water,  
pack an elephant into a sugar cube)*



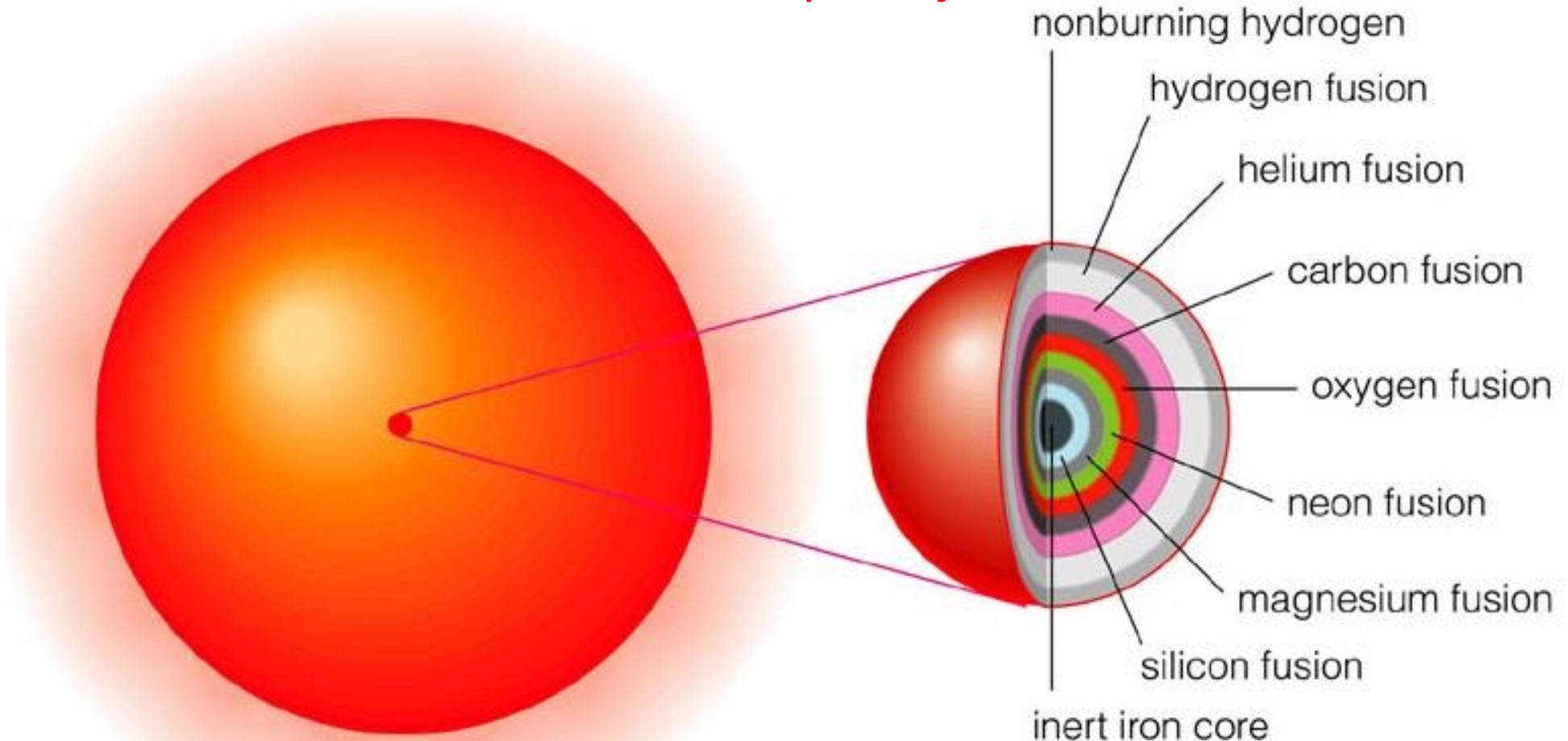
What about higher mass stars?

# Stellar life tracks



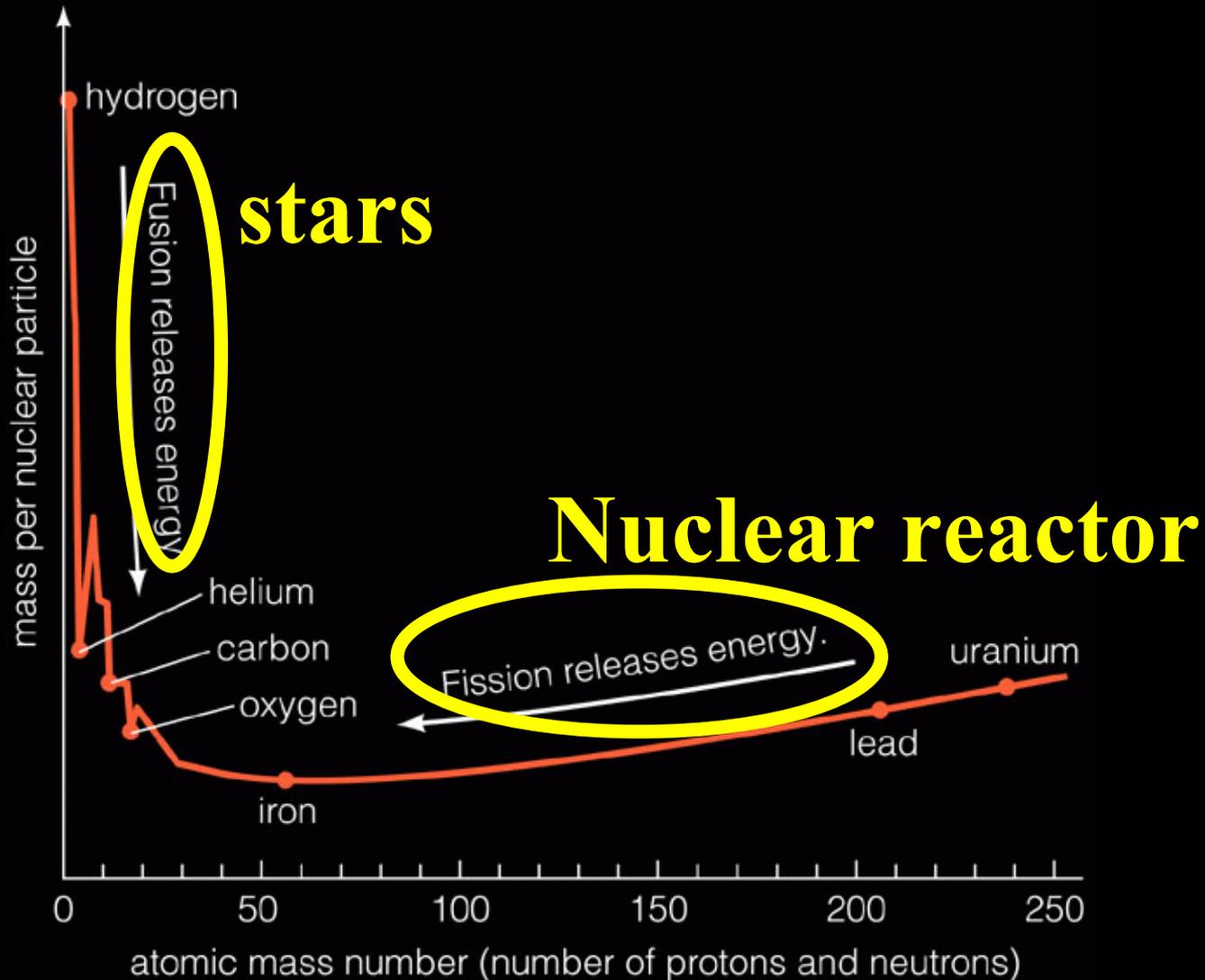
# Stellar Onion

Unlike low-mass stars which can only process hydrogen and helium to carbon and oxygen, high-mass stars have hotter cores and therefore can burn more completely.



However, when silicon is burned to iron in the stellar core, no more nuclear fusion possible!!!!

# Nuclear energy



# They go **SUPERNOVA!**

the explosion of a star, possibly caused by gravitational collapse, during which the star's luminosity increases by as much as 20 magnitudes and most of the star's mass is blown away at very high velocity, sometimes leaving behind an extremely dense core. --- Random House dictionary

**SN 1987A**



First discoverer: Dr. I. Shelton (UofT)

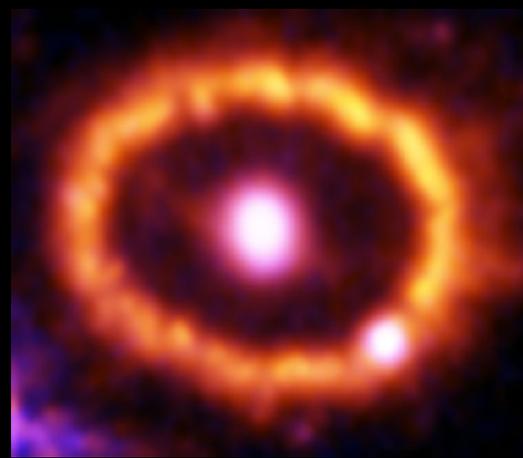
Astronomers are trying to reconstruct  
this most energetic event using computers  
(movie).

So far we haven't totally succeeded.

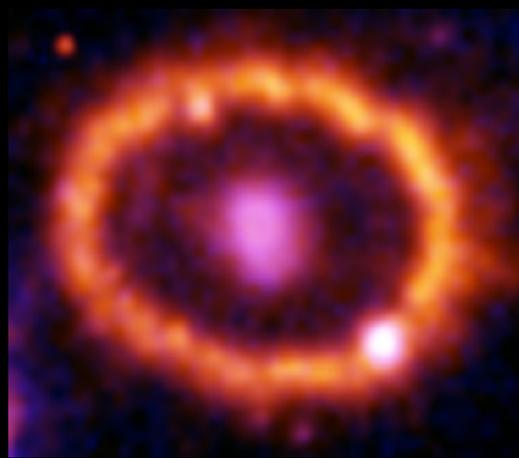
# SN 1987A



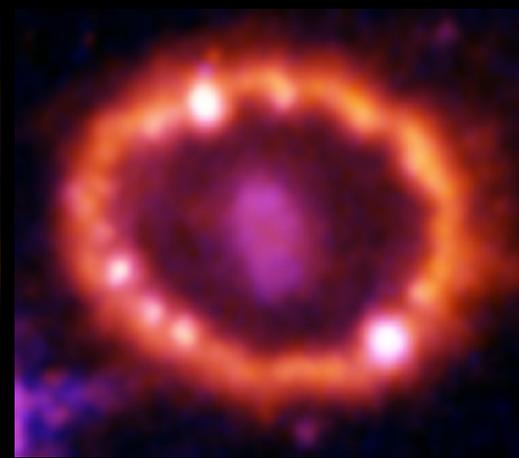
# SN 1987A



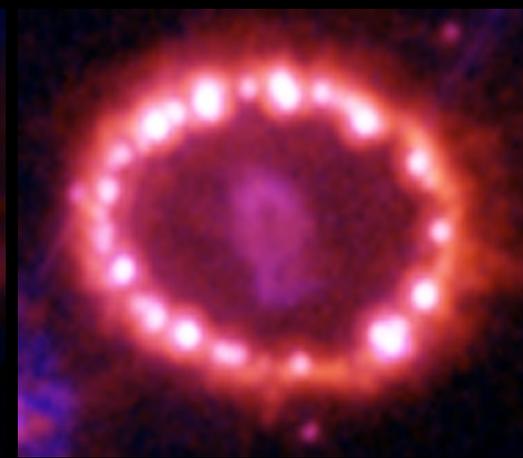
1994



1997



2000



2003

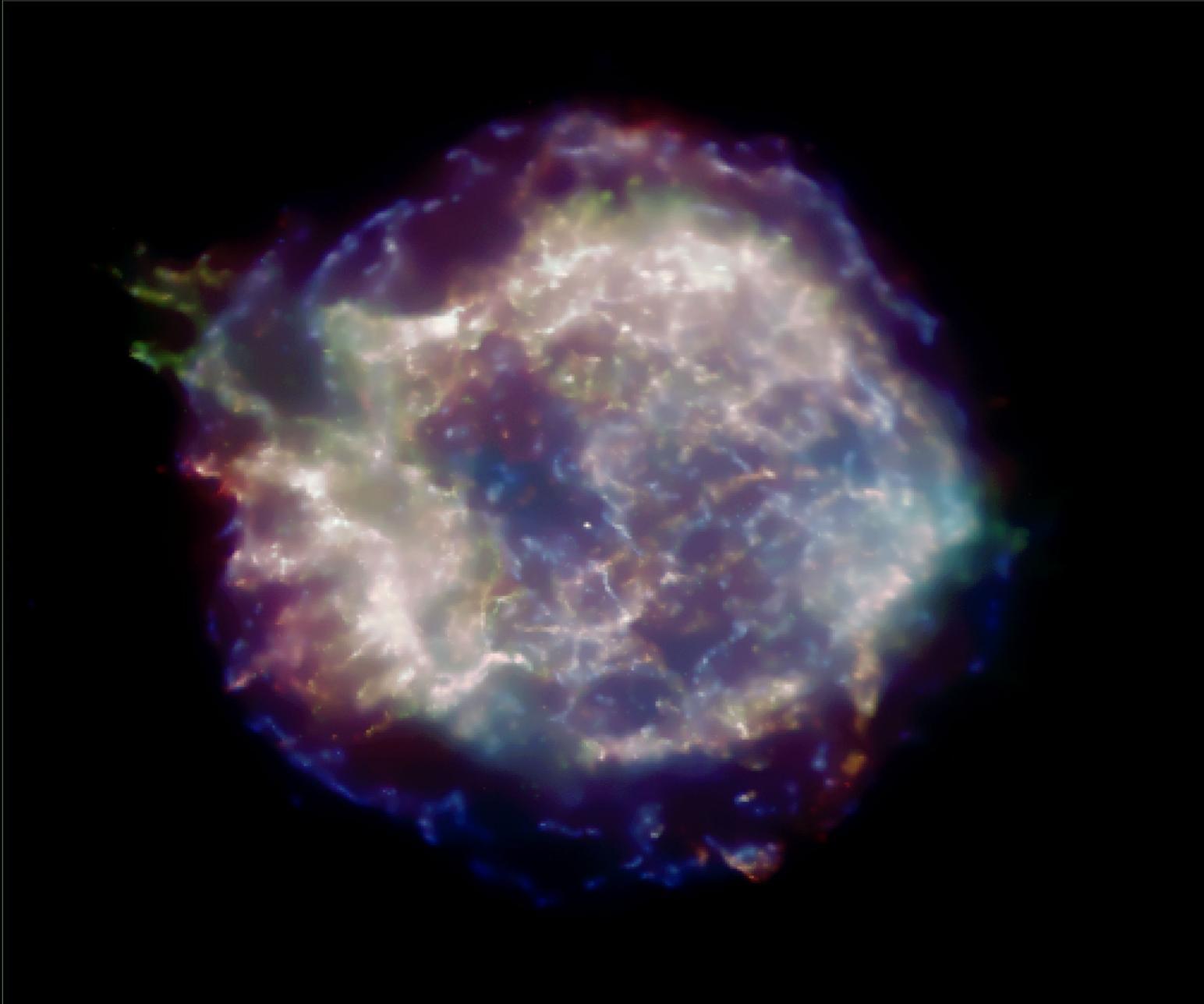


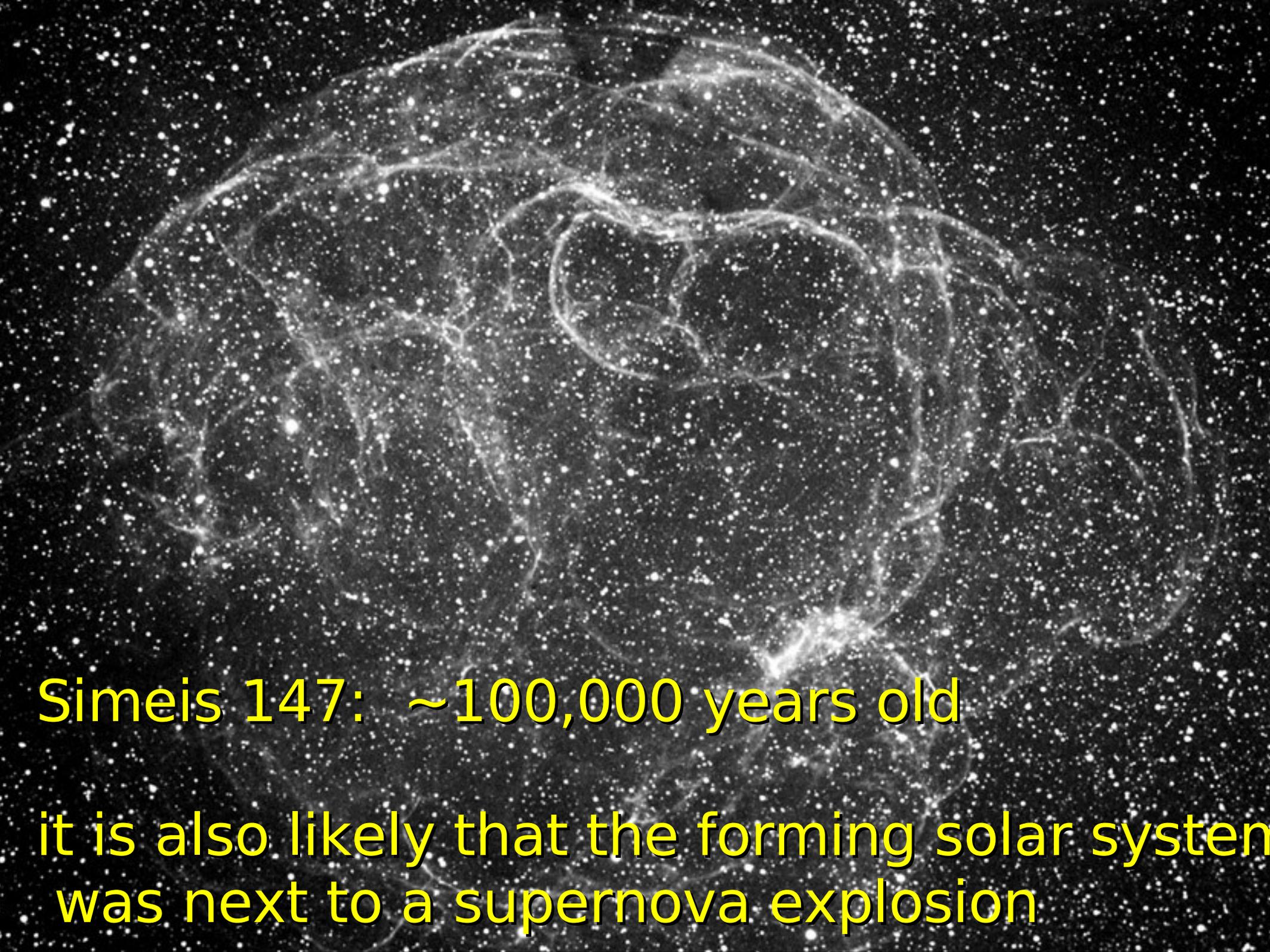
Crab supernova remnant:  
exploded July 4th, AD1054



*“In the 1st year of the period Chih-ho, the 5th moon, the day chi-ch'ou, a guest star appeared approximately several inches south-east of Tien-Kuan [Zeta Tauri]. After more than a year, it gradually became invisible .. “ (Chinese records)*

Cassiopeia A **supernova remnant:**  
Milky way, 300 years old





Simeis 147: ~100,000 years old

it is also likely that the forming solar system  
was next to a supernova explosion

An extra-galactic supernova (2005cs) – spot it if you can.

