

AST251 Project 3 – Evaluating Claims of  
Extraterrestrial Messaging  
ozerdend  
Planet 1

Friday 11<sup>th</sup> September, 2099

**We have identified what may be an indication of extraterrestrial intelligence, as well as the planet where it may have originated. This document summarizes the information gathered so far about the candidate message and its candidate planet of origin.**

**Potential evidence for extraterrestrial intelligence**

Astronomers have detected a narrowband optical transmission that appears to have originated from this planet’s solar system. The transmission is believed to contain an image and is displayed below with the most likely aspect ratio. The transmission lasted a short duration and then stopped. The transmission is shown below:

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111100000000000000000000000000000000000000000000000000000000000000000000000000001111
11110111111111111111111101011111100000001000000011111111101011111100111111001111
1111011111111111111111110110111110000011100000111111111101101111100011111001111
11110111111111111111111101101111100001111100001111111111011101111100001111001111
1111011111111111111111110111101110001111111000111111111101111011101110000111001111
1111011111111111111111110111110110011111111100111111111101111101110110000011001111
1111011111111111111111110111111010011111111110011111111111011111010000001001111
1111000000000000000000000000000000000000000000000000000000000000000000000000000001111
```

This signal was first noticed at UTC 2096-04-21/10:43.

**Parameters of the candidate planet of origin and its host star**

Spectral Type	K
Stellar Luminosity (Solar Units)	0.273
Stellar Mass (Solar Masses)	0.723
Distance to Star (lightyears)	27.7
Planet Mass (Earth masses)	0.3
Atmospheric Pressure (atm)	5.9

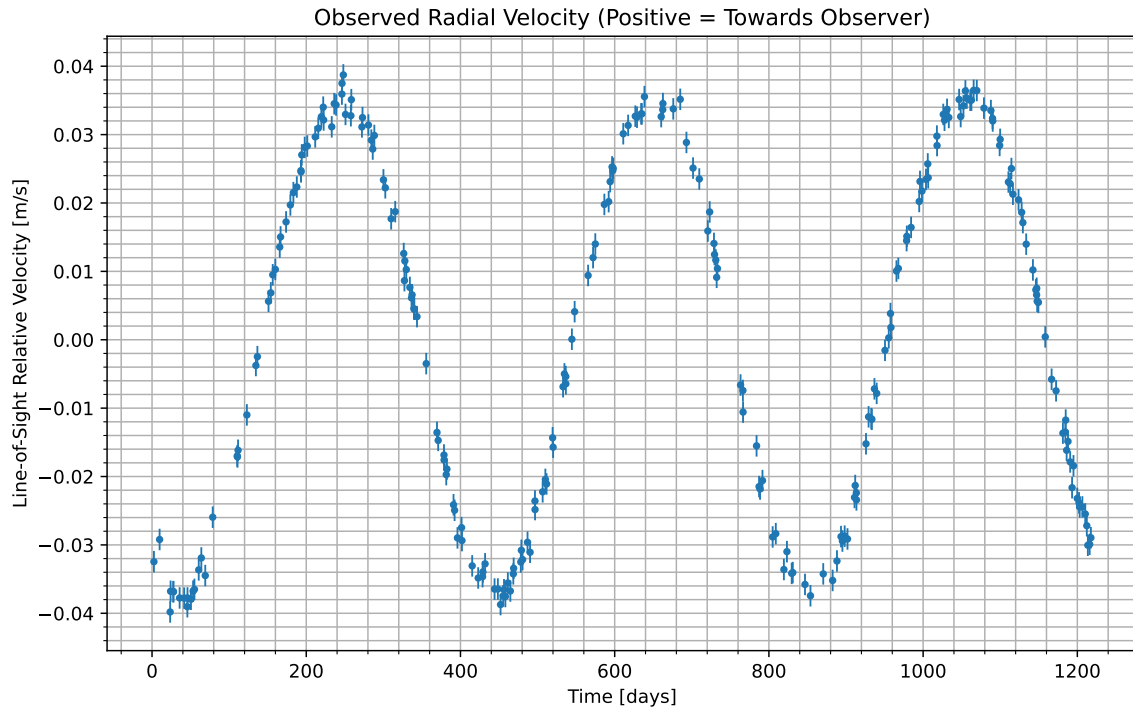


Figure 1: We have isolated the radial velocity of the host star due to the candidate planet. Data begins at UTC 2096-04-23/14:55. Positive values indicate the velocity at which the star is moving towards us; negative indicate the velocity at which it is moving away.

**Atmospheric composition of the candidate planet (percent by volume)**

Molecule	Concentration
$N_2$	49.5
$CO$	14.9
$CO_2$	3.74
$SiH_4$	11.1
$H_2S$	20.7

### Gas Abundance (percent by volume)

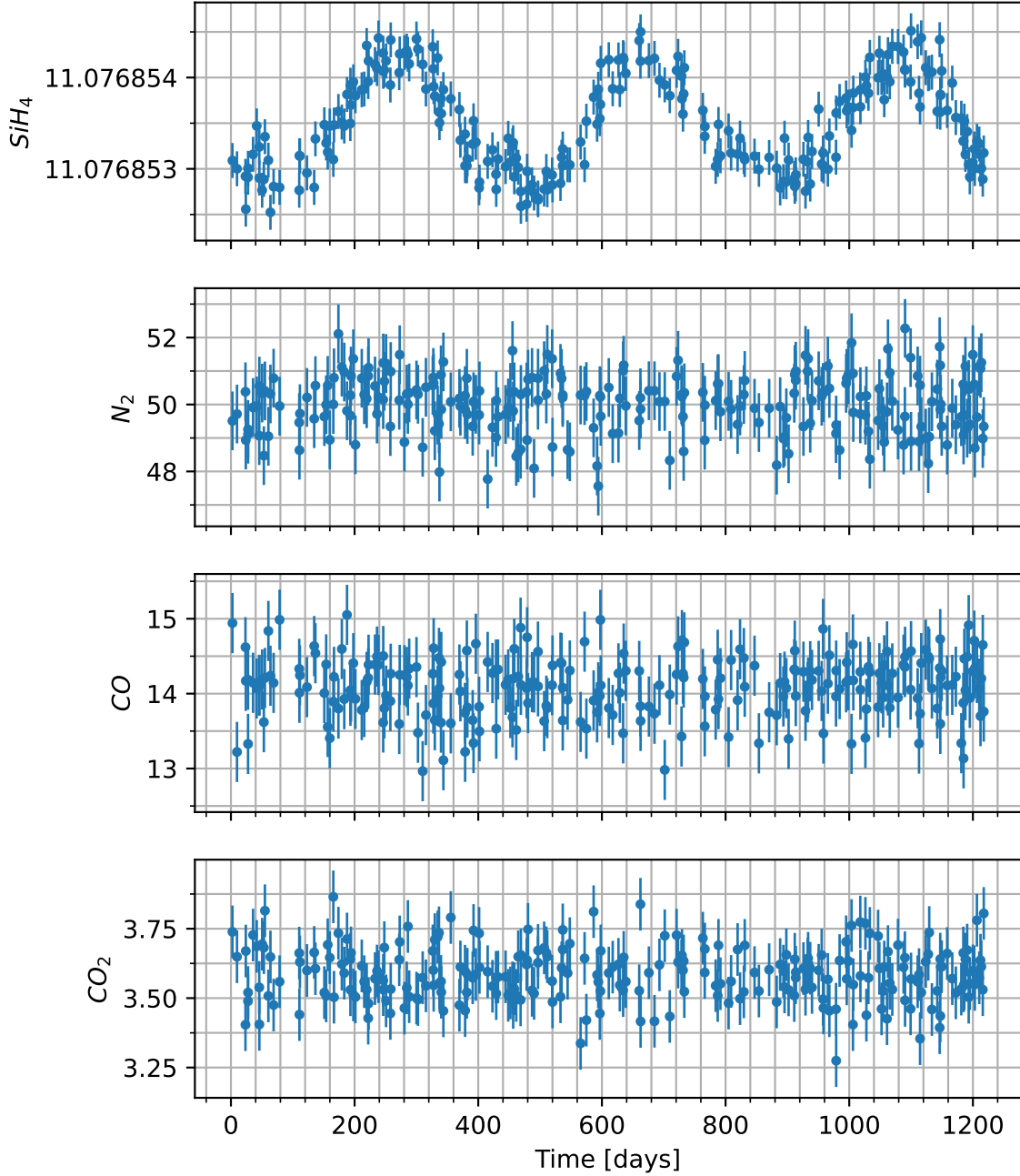


Figure 2: Concentration of various gases in the atmosphere of the candidate planet versus time. Note that the y-axis will usually only show the variation multiplied by some factor, shown in the upper left, and then added to some normal amount, also in the upper-left.

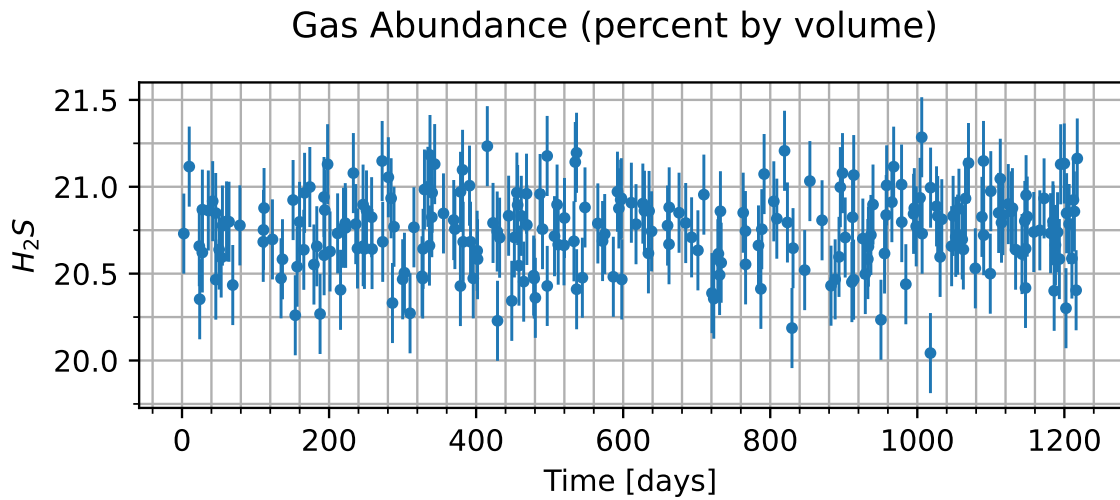


Figure 3: Concentration of various additional gases in the atmosphere of the candidate planet versus time. Note that the y-axis will usually only show the variation multiplied by some factor, shown in the upper left, and then added to some normal amount, also in the upper-left.

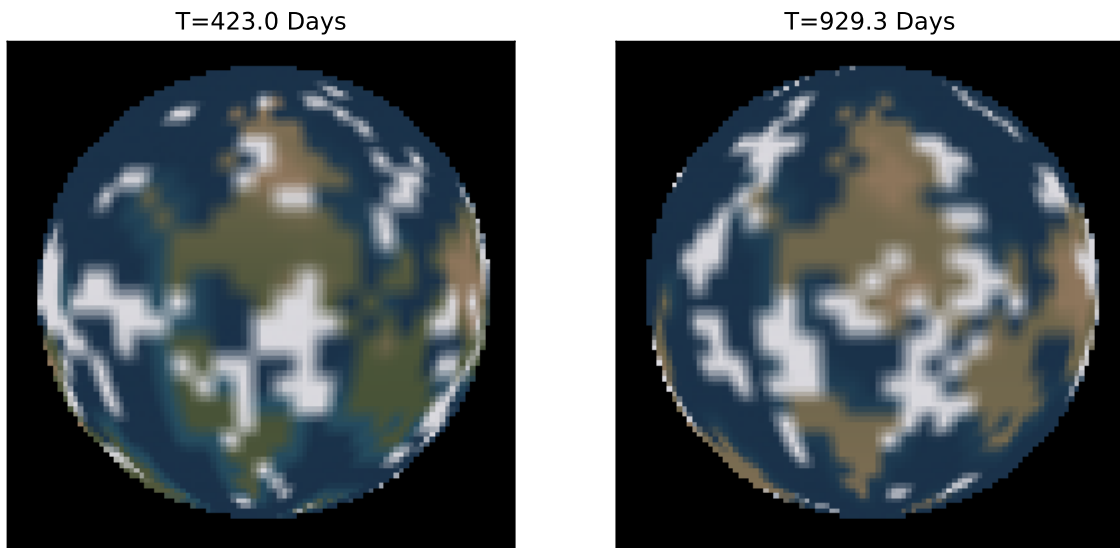


Figure 4: Maps of the surface of the candidate planet taken at two different times. Times are indicated above each image relative to the times shown in the radial velocity curve. Those maps are shown here. Tan areas indicate what we believe to be land, while blue-ish areas indicate what we believe to be liquid regions of some kind. Other colors present reflect the visible color as best as we are able to measure.