

# AST251 Project 3 – Evaluating Claims of Extraterrestrial Messaging

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### Planet 3

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**We have identified what may be an indication of extraterrestrial intelligence, as well as the planet where it may have originated. This document summarizes the information gathered so far about the candidate message and its candidate planet of origin.**

### **Potential evidence for extraterrestrial intelligence**

Astronomers have detected a broadband microwave transmission that appears to have originated from this planet's solar system. The transmission is believed to contain an image and is displayed below with the most likely aspect ratio. The transmission lasted a short duration and then stopped. The transmission is shown below:

```
011110010101100110001011111101001110001111110111110101101111000000000
000100011011101111001111001011010111100110011010110110000100001001010100
001111100101000111000110110001001010001110100100001001000011011011011010
010101001100100010000000011101101110001001010010010000100000100011101100
000001101110110100111111011001101111001100001101000010111100111000101011
010010011111000000110100100111000110000101001011110010101010110001011100
11011111011100011011001110000001011010111000010000000011111111111011100
```

This signal was first noticed at UTC 2080-07-02/05:30.

### **Parameters of the candidate planet of origin and its host star**

Spectral Type	K
Stellar Luminosity (Solar Units)	0.125
Stellar Mass (Solar Masses)	0.595
Distance to Star (lightyears)	95.0
Planet Mass (Earth masses)	3.8
Atmospheric Pressure (atm)	0.8

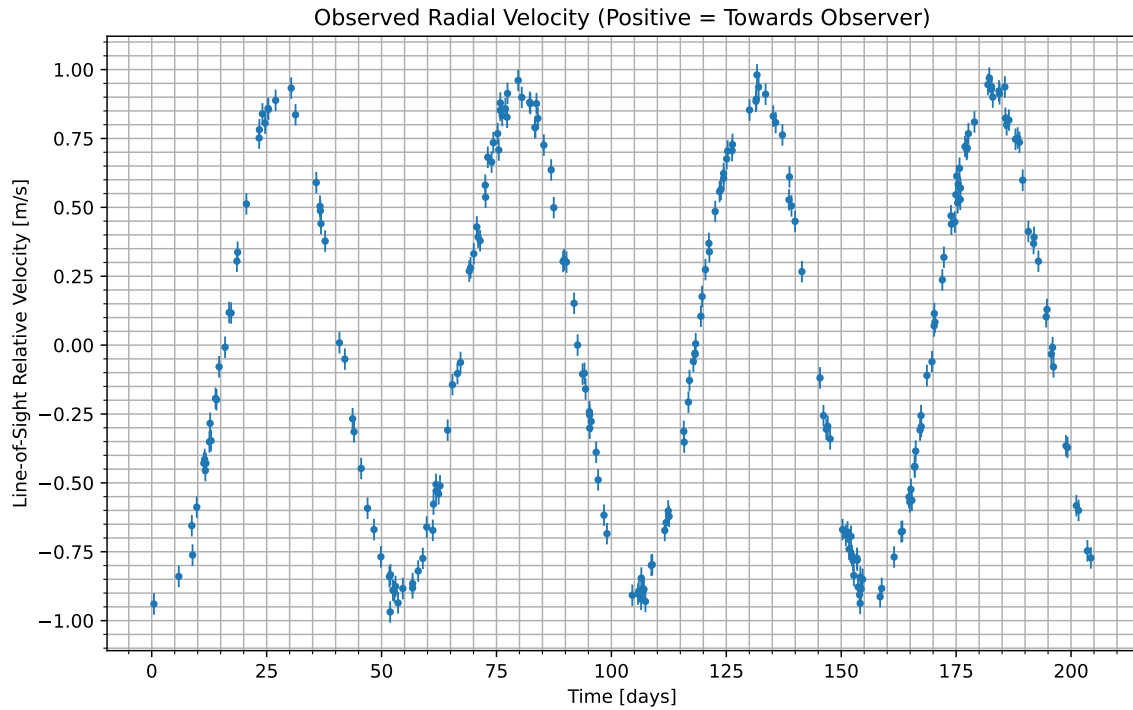


Figure 1: We have isolated the radial velocity of the host star due to the candidate planet. Data begins at UTC 2080-07-02/12:14. Positive values indicate the velocity at which the star is moving towards us; negative indicate the velocity at which it is moving away.

### Atmospheric composition of the candidate planet (percent by volume)

Molecule	Concentration
$N_2$	12.7
$CO_2$	80.6
$H_2O$	6.72

### Gas Abundance (percent by volume)

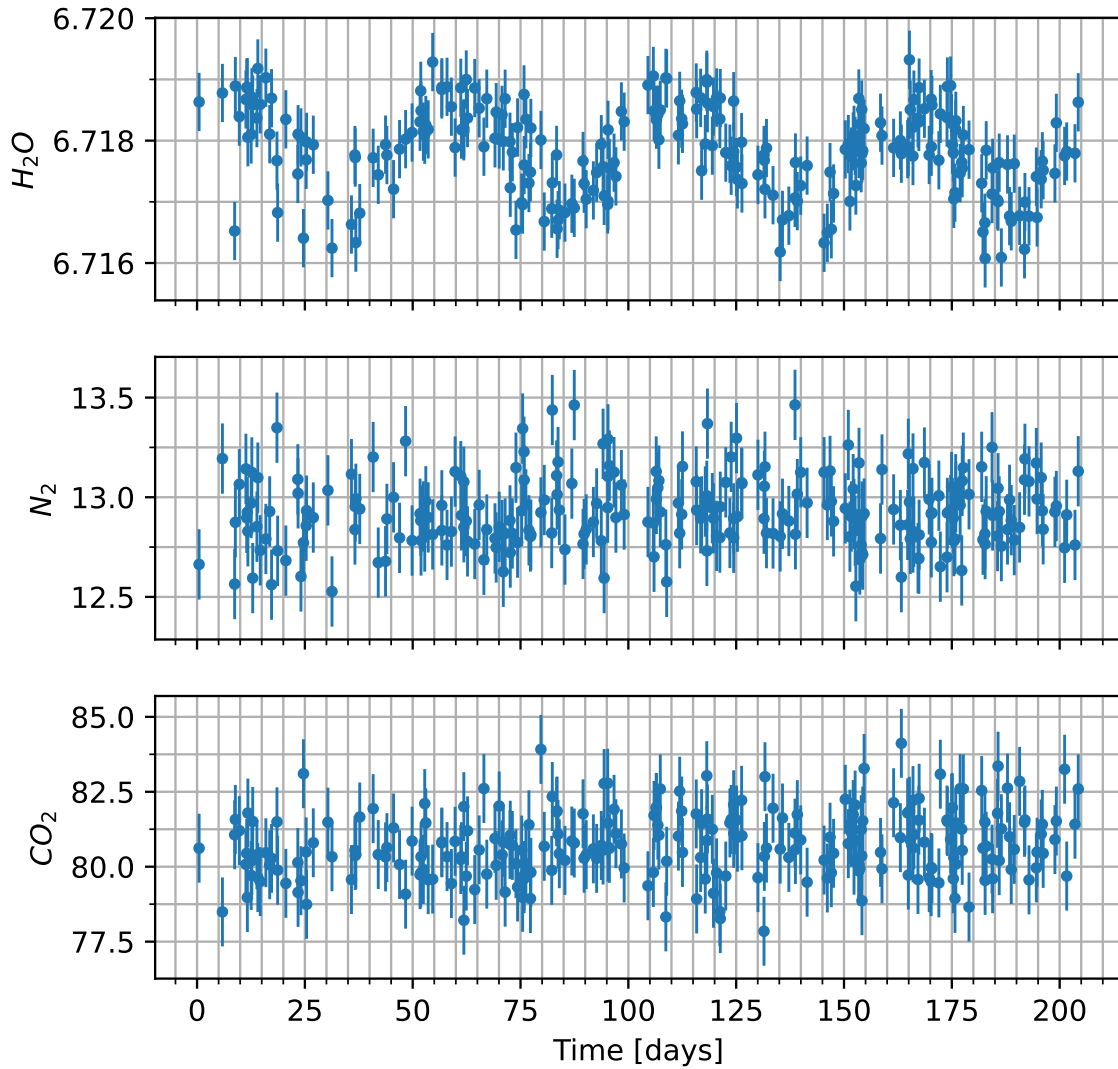


Figure 2: Concentration of various gases in the atmosphere of the candidate planet versus time. Note that the y-axis will usually only show the variation multiplied by some factor, shown in the upper left, and then added to some normal amount, also in the upper-left.

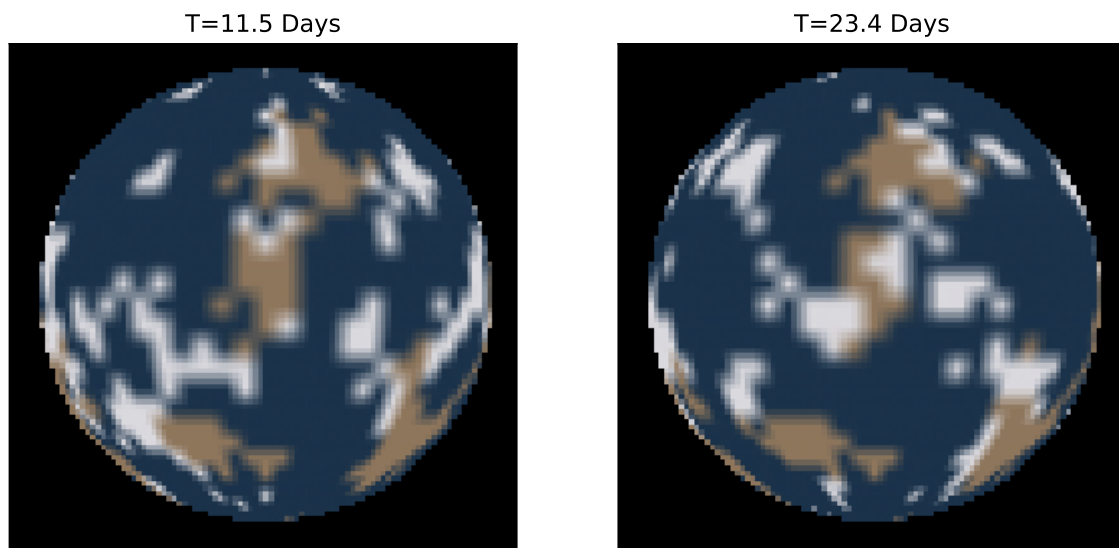


Figure 3: Maps of the surface of the candidate planet taken at two different times. Times are indicated above each image relative to the times shown in the radial velocity curve. Those maps are shown here. Tan areas indicate what we believe to be land, while blue-ish areas indicate what we believe to be liquid regions of some kind. Other colors present reflect the visible color as best as we are able to measure.